

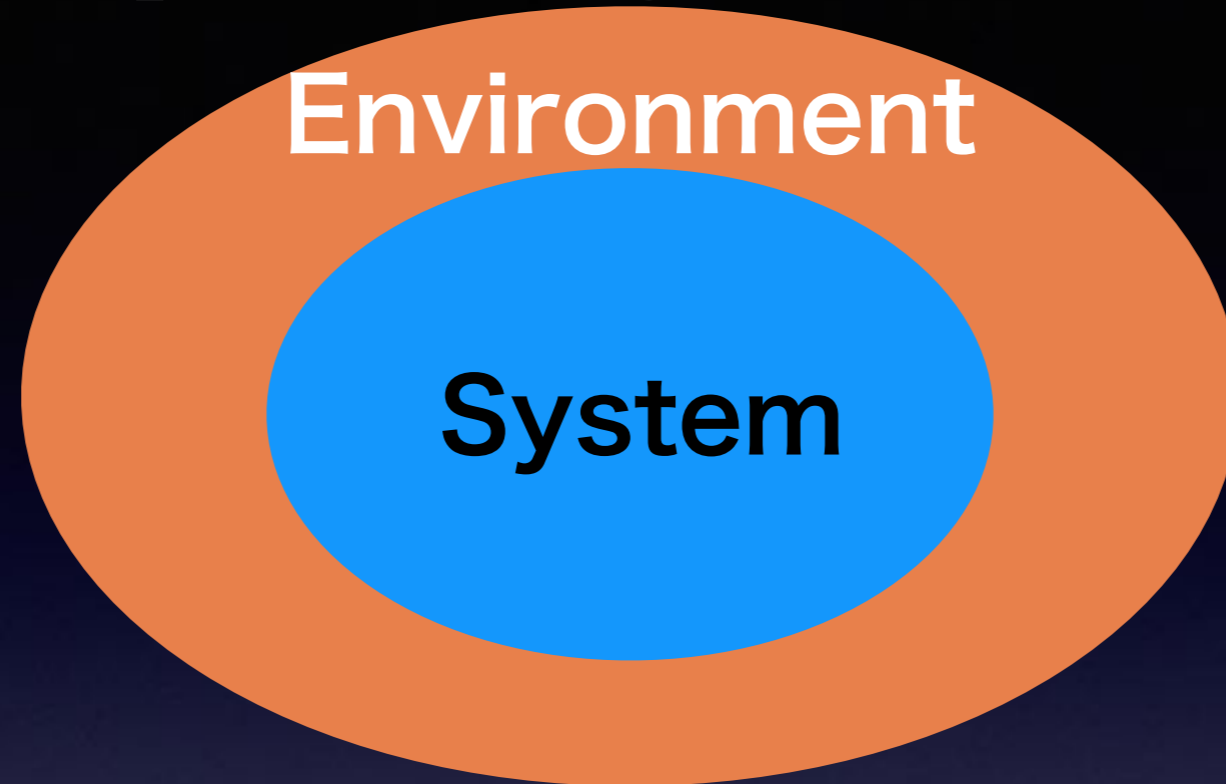
Spontaneous symmetry breaking in open systems

Yoshimasa Hidaka
RIKEN

Collaboration with Yuki Minami

1509.05042[cond-mat.stat-mech], 1712.xxxx

Open systems



Environment: QGP
System: Heavy quarks

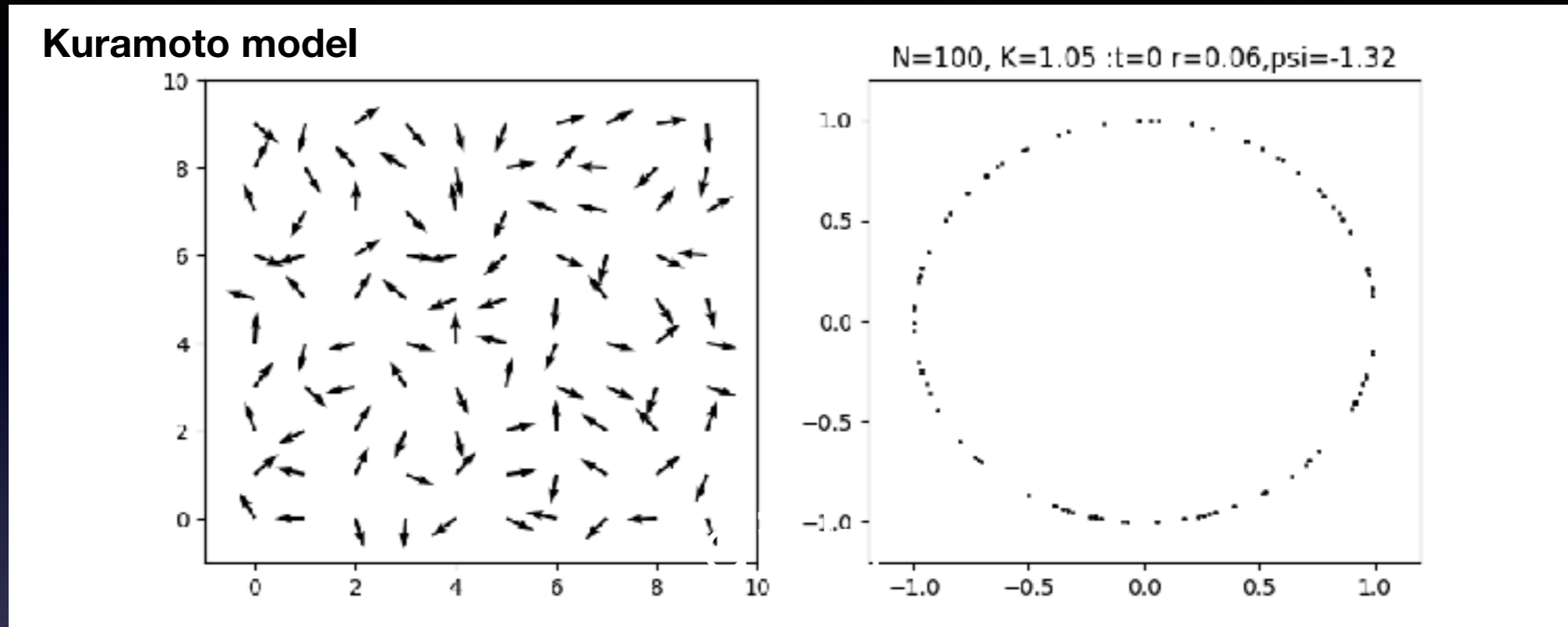


CC BY-SA 2.0

Environment: Air
System: flock of birds

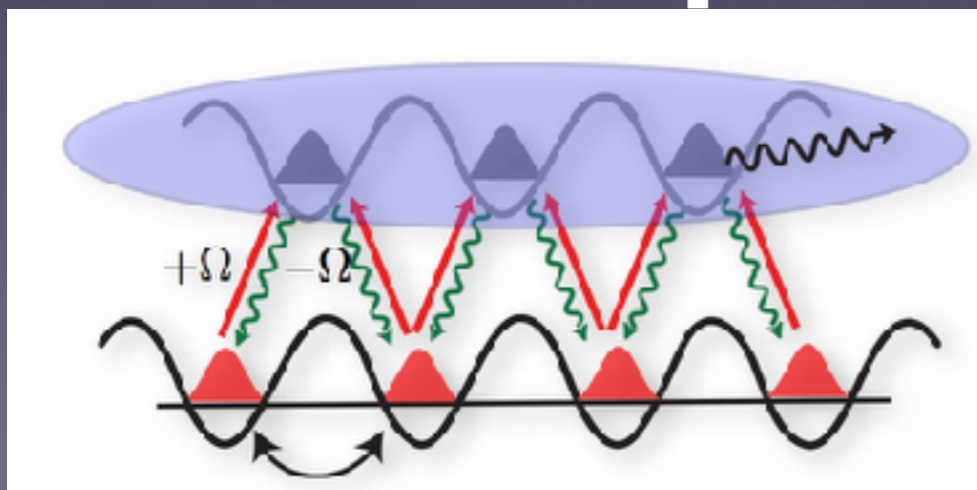
Symmetry breaking in open systems

Synchronization



Metronome, fireflies, ...

Driven dissipative condensate



Driving force and dissipative causes a condensate.

Diehl, Micheli, Kantian, Kraus, Büchler, Zoller, Nature Physics 4, 878 (2008);
Kraus, Diehl, Micheli, Kantian, Büchler, Zoller, Phys. Rev. A 78, 042307 (2008).

figure is taken from Diehl's website

Questions

Hamiltonian systems

**Continuum
symmetry**



$$\partial_{\mu} J^{\mu} = 0$$

Open systems

$\partial_{\mu} J^{\mu} \neq 0$ because of friction

What is the symmetry?

Is there any symmetry breaking?

Does a NG mode appear?

Nambu-Goldstone theorem

Nambu('60), Goldstone(61), Nambu Jona-Lasinio('61),
Goldstone, Salam, Weinberg('62).

For Lorentz invariant vacuum

Spontaneous breaking of global symmetry



$$N_{NG} = N_{BS}$$

of NG modes

of broken symmetries

Dispersion relation: $\omega = c|k|$

Gapless modes in nature

pion

superfluid phonon

U(1) symm.

spin waves

spin symm.

Need generalization

We focus on open systems

NG theorem OK nodes

photon

U(1) 1-form symm.

surface waves

translation symm.

diffusive modes

rotation symm.

U(1) symm.



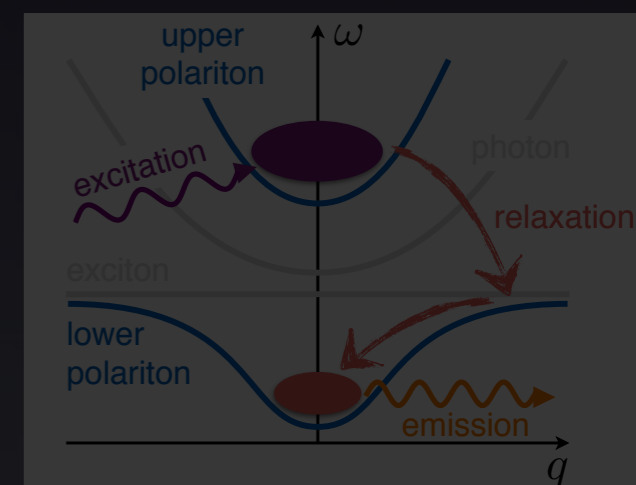
CC by Zouavman Le Zouave



CC by-sa Roger McLassus



CC BY-SA 2.0



Classification of Nambu-Goldstone modes in Hamiltonian system

Exception of NG theorem

NG modes with $N_{\text{BS}} \neq N_{\text{NG}}$ and $\omega \neq k$ exist

NG modes in Kaon condensed CFL phase

Miransky, Shovkovy ('02) Schafer, Son, Stephanov, Toublan, and Verbaarschot ('01)

$$SU(2)_I \times U(1)_Y \rightarrow U(1)_{\text{em}}$$

$$N_{\text{BS}} = 3, \quad N_{\text{NG}} = 2$$

Dispersion: $\omega \propto k$ and $\omega \propto k^2$

Magnon



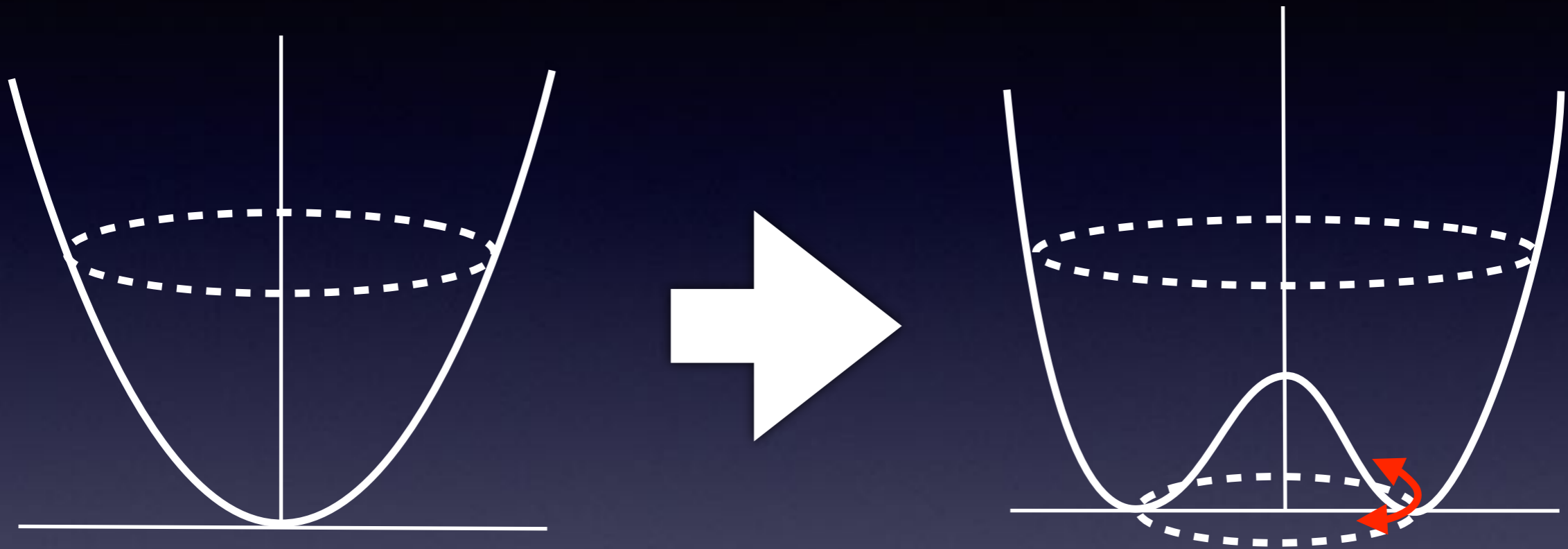
spin rotation $SO(3) \rightarrow SO(2)$

$$N_{\text{BS}} = \dim(G/H) = 2 \quad N_{\text{NG}} = 1$$

Dispersion: $\omega \propto k^2$

Internal symmetry breaking

Symmetry group $G \Rightarrow H$



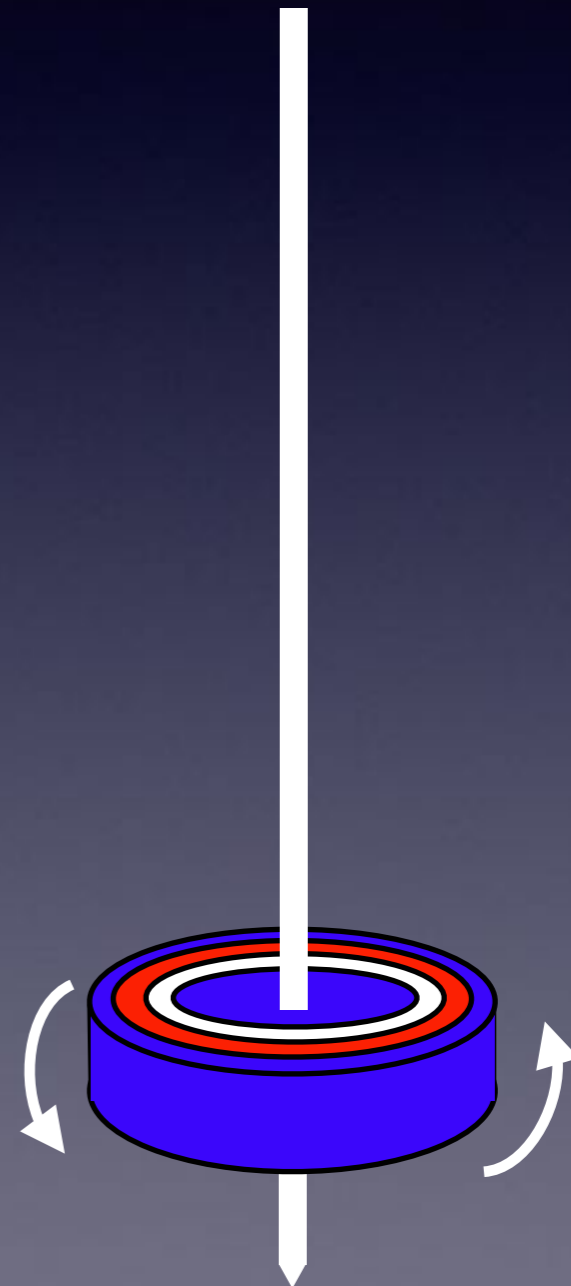
of flat direction
 $N_{\text{BS}} = \dim(G/H)$

**This does work in nonrelativistic system
at zero and finite temperature**

Intuitive example for type-B NG modes

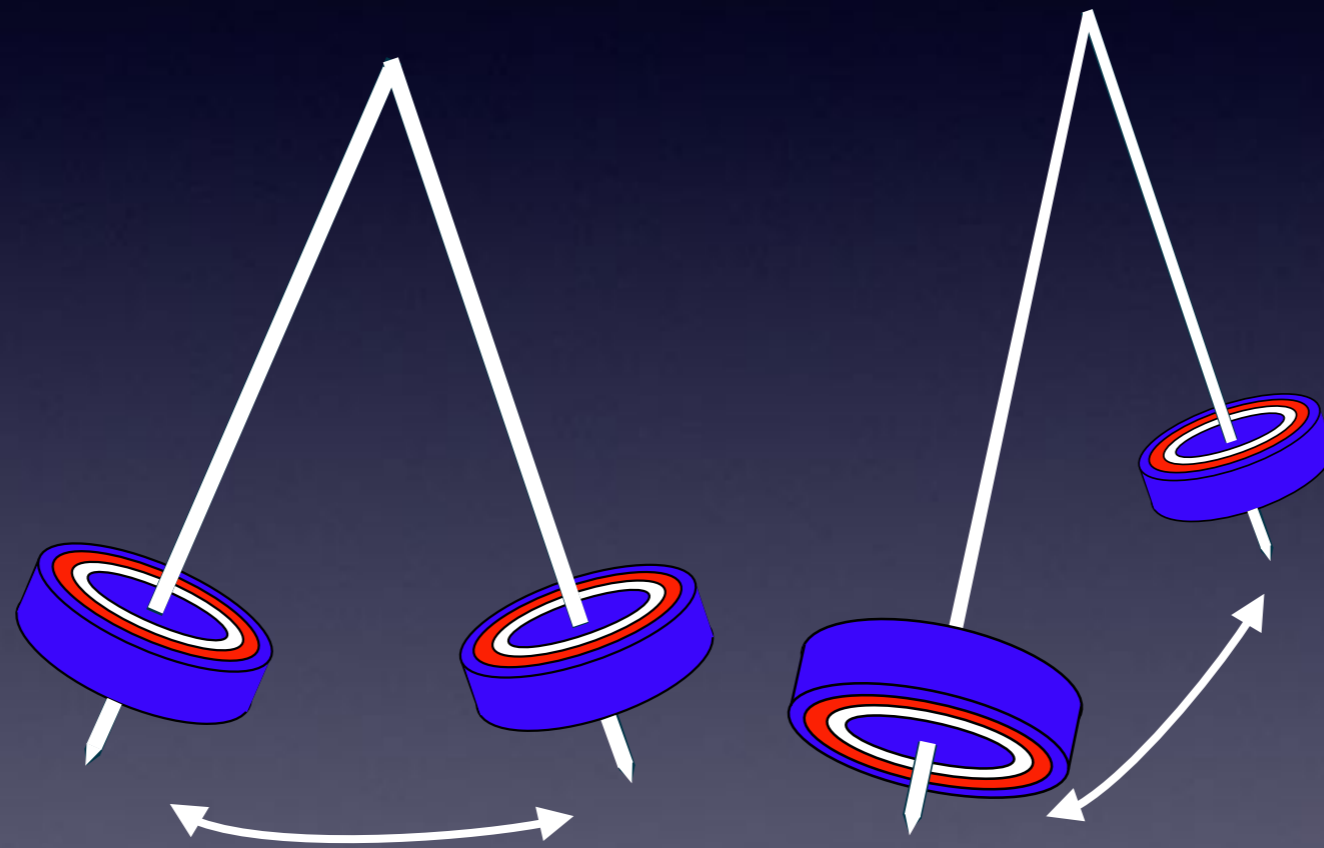
Pendulum with a spinning top

- Rotation symmetry is explicitly broken by a weak gravity
- Rotation along with z axis is unbroken.
- Rotation along with x or y is broken.
- The number of broken symmetry is two.



Intuitive example for type-B NG modes

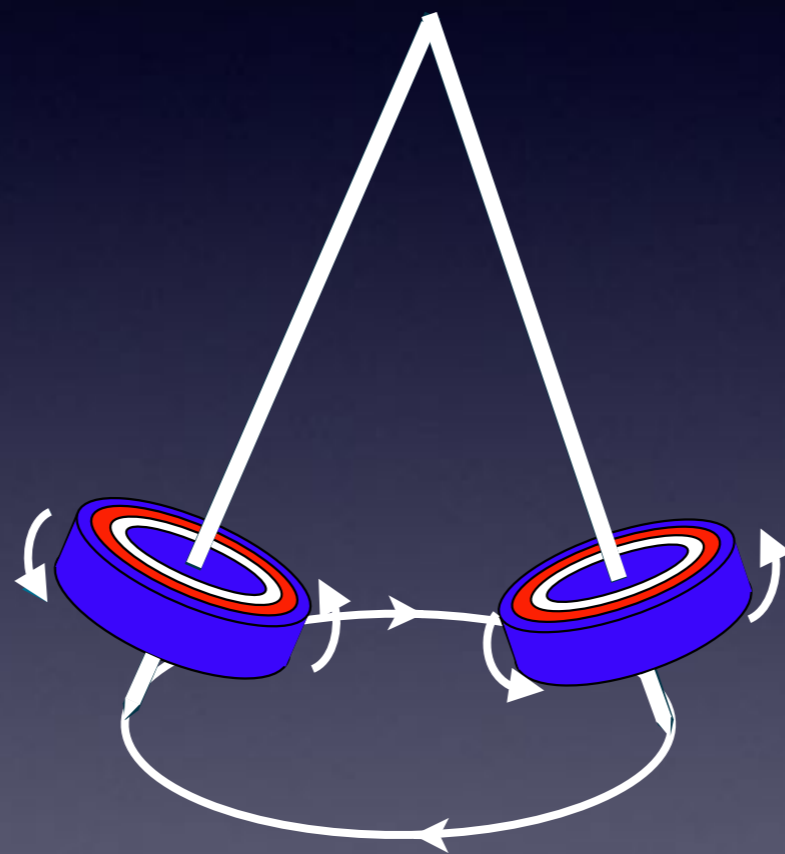
Pendulum has two oscillation motions



if the top is not spinning.

Intuitive example for type-B NG modes

If the top is spinning,



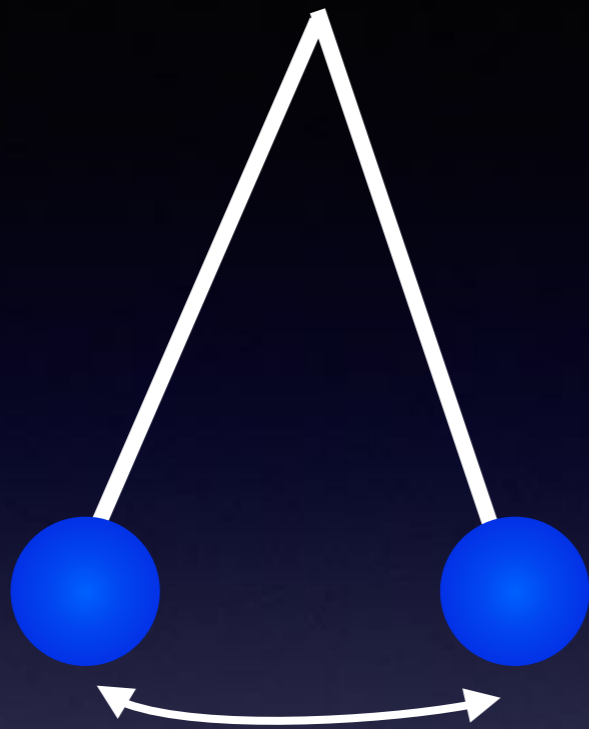
the only one rotation motion (Precession) exists.

In this case, $\{L_x, L_y\}_P = L_z \neq 0$

Classification of NG modes

Watanabe, Murayama ('12), YH ('12)

cf. Takahashi, Nitta ('14), Beekman ('14)

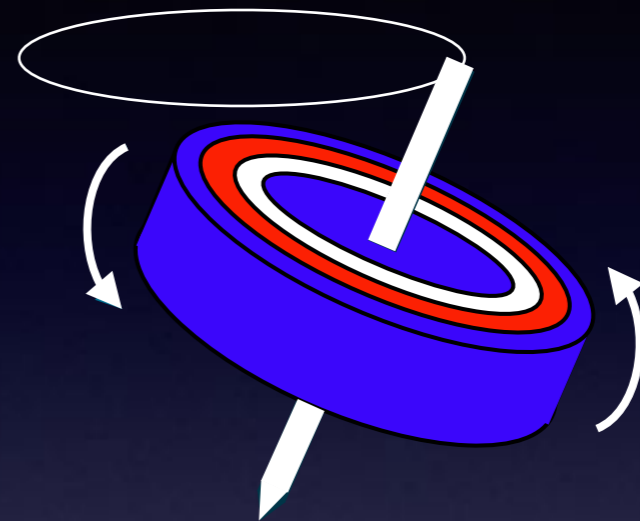


Type-A

Harmonic oscillation

$$N_A = N_{\text{BS}} - \text{rank} \langle [iQ_a, Q_b] \rangle$$

Ex.) superfluid phonon

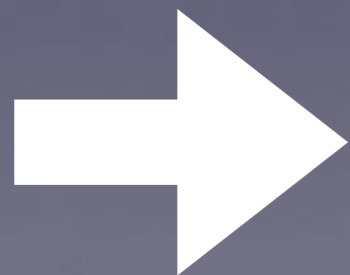


Type-B

Precession

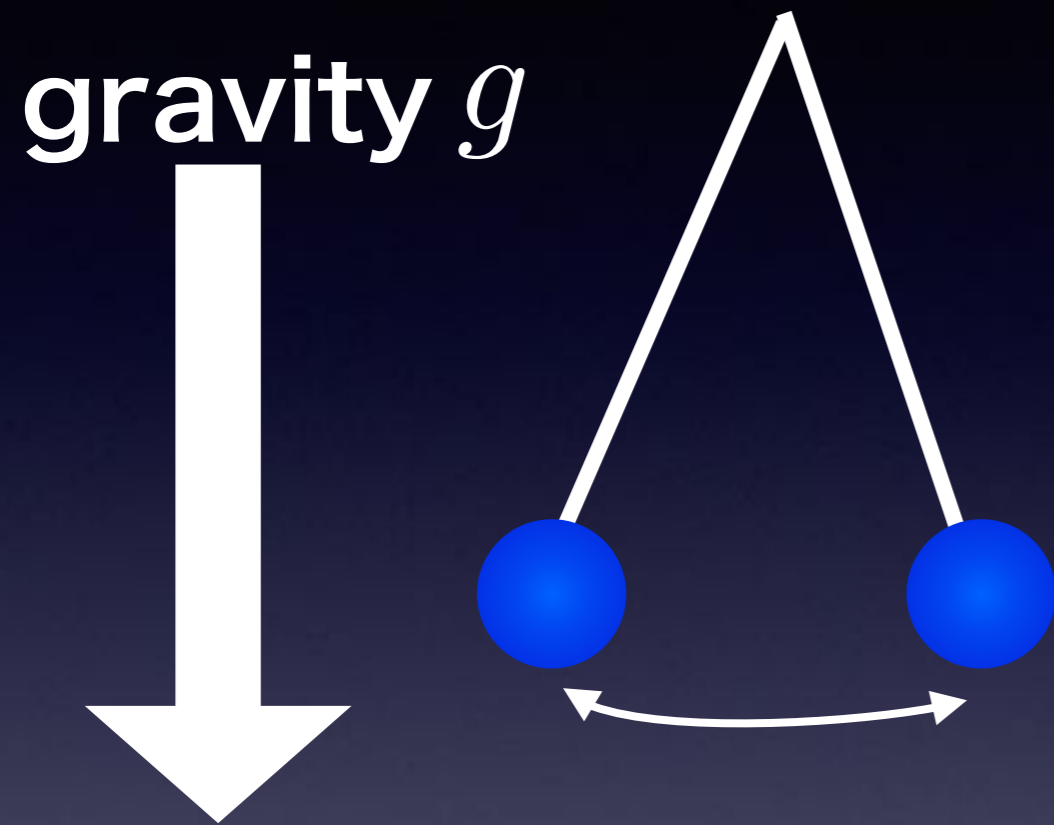
$$N_B = \frac{1}{2} \text{rank} \langle [iQ_a, Q_b] \rangle$$

Ex.) magnon



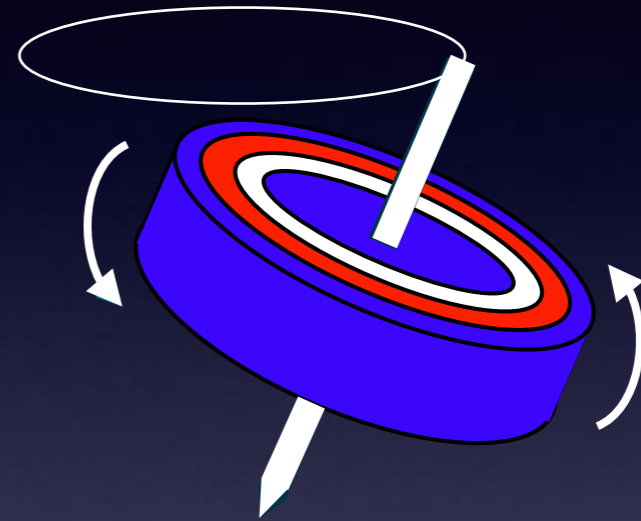
$$N_{\text{NG}} = N_{\text{BS}} - \frac{1}{2} \langle i[Q_a, Q_b] \rangle$$

Dispersion relation



Type-A

$$\omega \sim \sqrt{g} \sim \sqrt{k^2}$$



Type-B

$$\omega \sim g \sim k^2$$

Examples of Type-B NG modes

	N_{BS}	$N_{\text{type-A}}$	$N_{\text{type-B}}$	$\frac{1}{2}\text{rank}\langle[Q_a, Q_b]\rangle$	$N_{\text{type-A}} + 2N_{\text{type-B}}$
Spin wave in ferromagnet $\text{SO}(3) \rightarrow \text{SO}(2)$	2	0	1	1	2
NG modes in Kaon condensed CFL $\text{SU}(2) \times \text{SU}(1)_Y \rightarrow \text{U}(1)_{\text{em}}$	3	1	1	1	3
Spinor BEC $\text{SO}(3) \times \text{U}(1) \rightarrow \text{U}(1)$	3	1	1	1	3
nonrelativistic massive CP^1 model $\text{U}(1) \times \mathbf{R}^3 \rightarrow \mathbf{R}^2$	2	0	1	1	2

$$N_{\text{type-A}} + 2N_{\text{type-B}} = N_{\text{BS}} \qquad N_{\text{BS}} - N_{\text{NG}} = \frac{1}{2}\text{rank}\langle[Q_a, Q_b]\rangle$$

At finite temperature

Hayata, YH ('14)

**The interaction with thermal particles
modifies the dispersion relation**

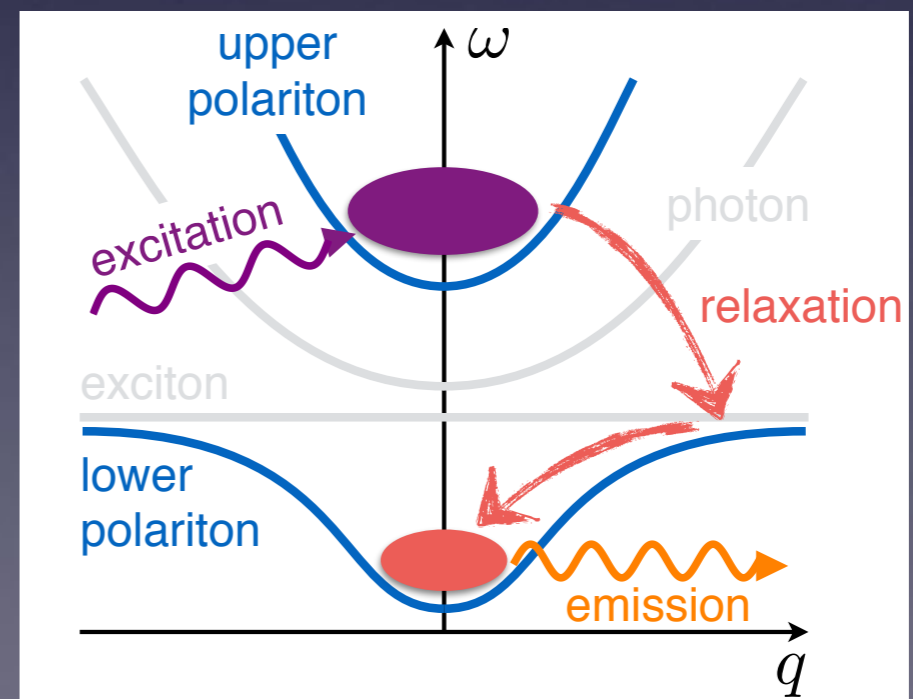
Type-A: $\omega = ak - ibk^2$

Type-B: $\omega = a'k^2 - ib'k^4$

Open system

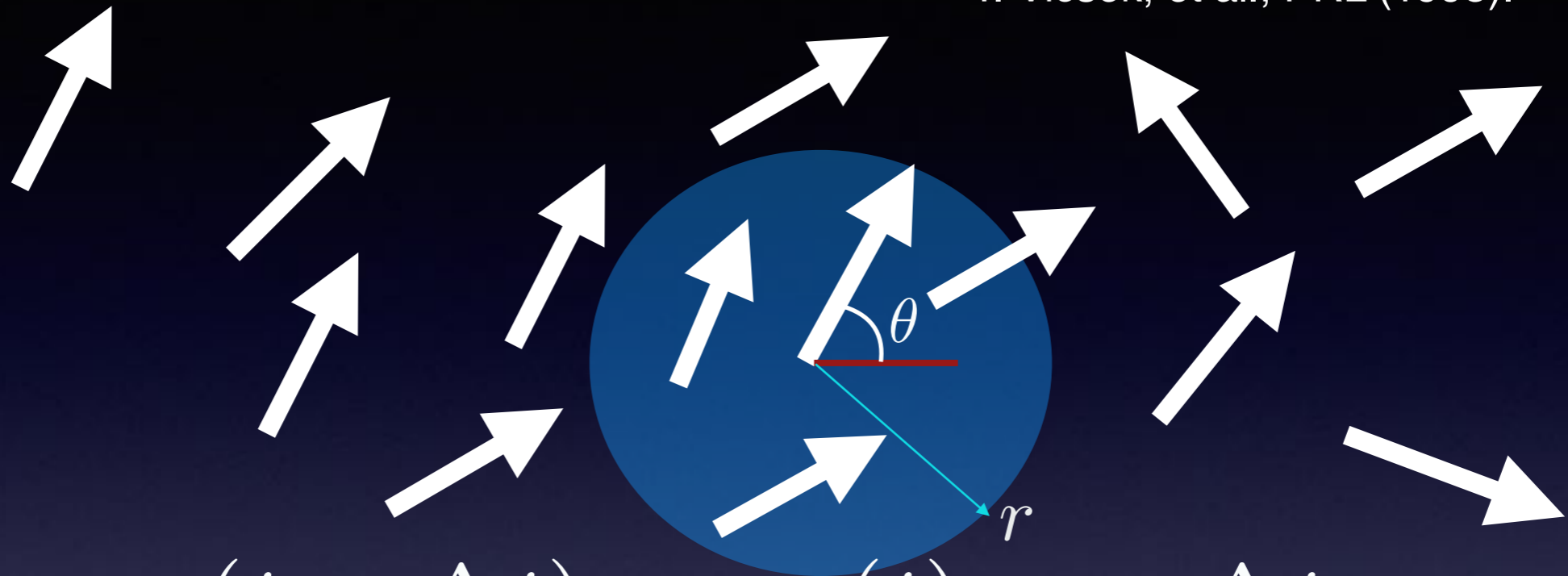


CC BY-SA 2.0



Ex2) Vicsek model

T. Vicsek, et al., PRL (1995).



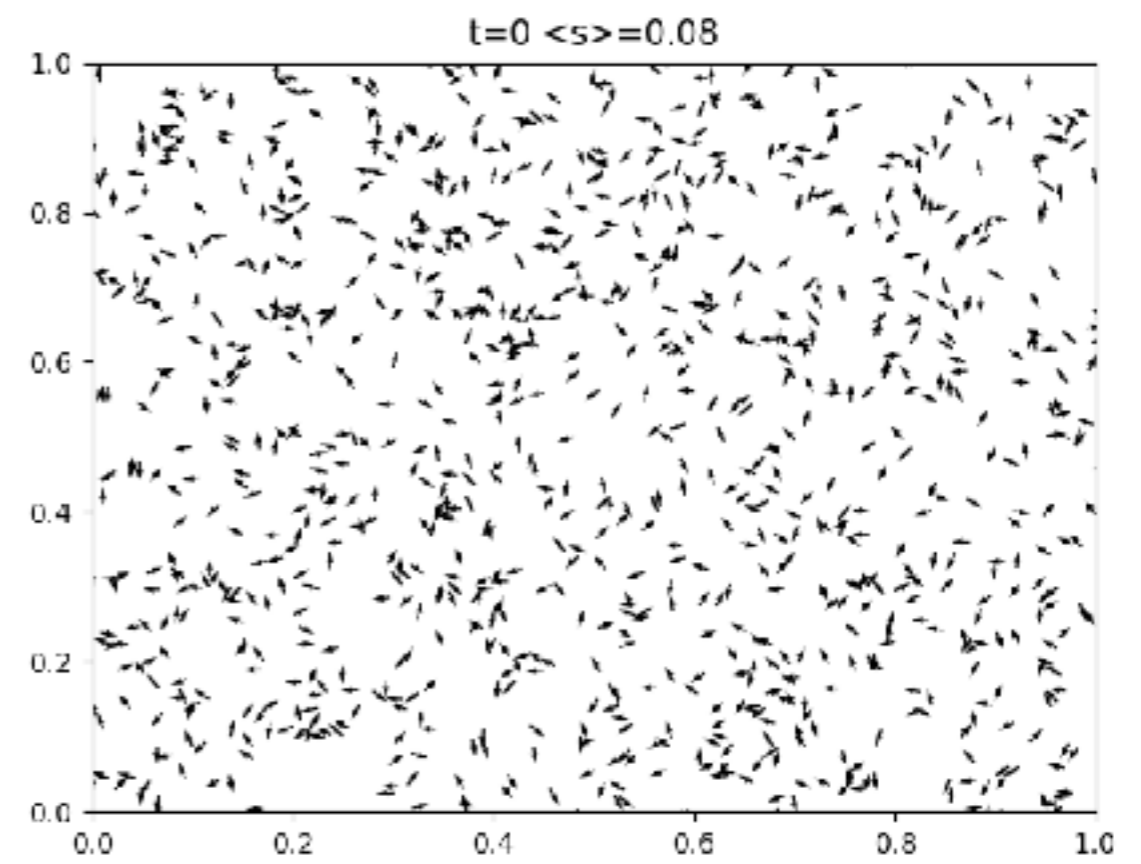
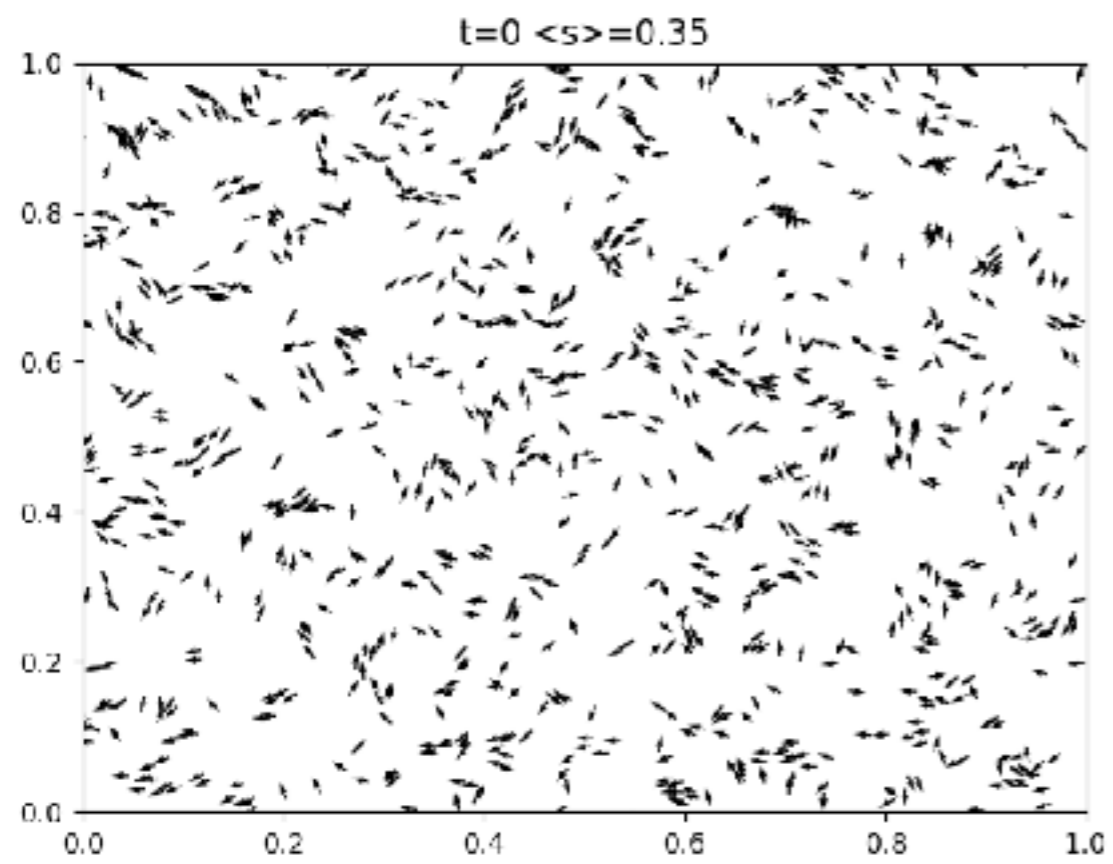
$$x_i(t + \Delta t) = x_i(t) + v_i \Delta t$$

velocity

$$v_i = v_0 (\cos \theta_i, \sin \theta_i)$$

$$\theta_i(t + \Delta t) = \langle \theta_i(t) \rangle_r + \xi_i$$

angle of velocity **average** **noise**
angle

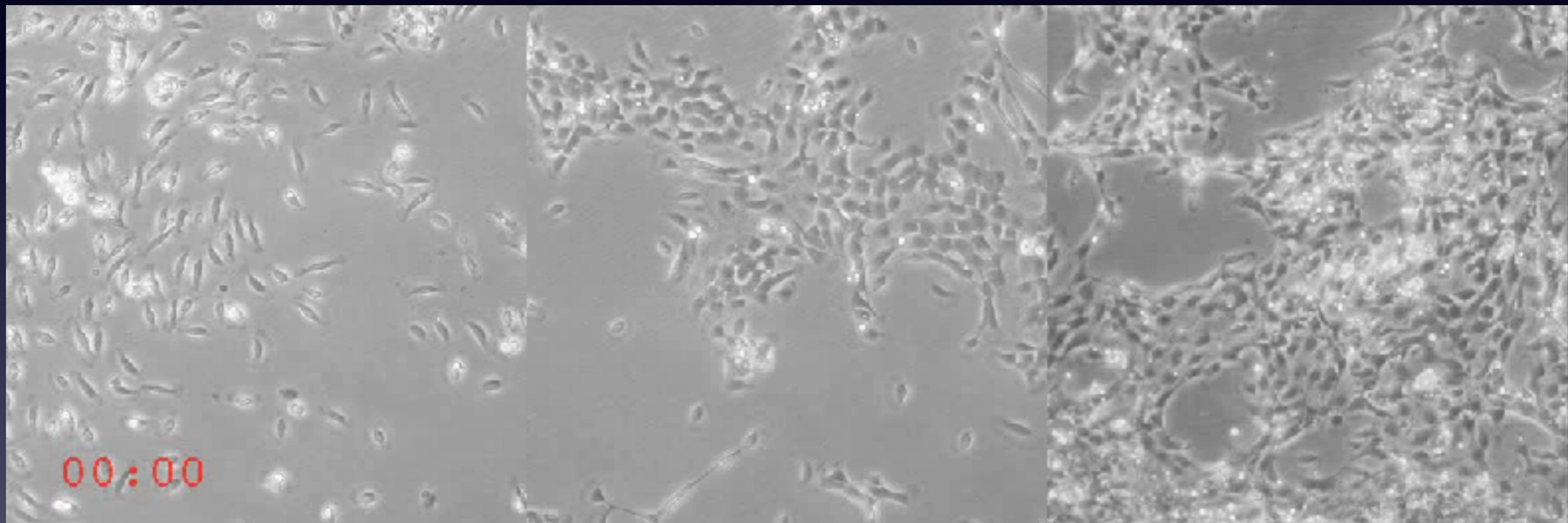


An example of SSB in open systems

Some cells on fish skin

low density

high density



B. Szabo, et al., Phys. Rev. E 74, 061908 (2006)

Model of active matter: Vicsek model, Active hydrodynamics,

T. Vicsek, et al., PRL (1995). J. Toner, and Y. Tu, PRL (1995).

Field theoretical model

Ex.) NG mode in Active hydrodynamics

J. Toner, and Y. Tu, PRE (1998)

$$\partial \rho + \nabla \cdot (\rho \mathbf{v}) = 0$$

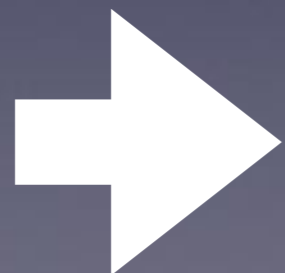
$$\partial_t \mathbf{v} + (\mathbf{v} \cdot \nabla) \mathbf{v} = \alpha \mathbf{v} - \beta \mathbf{v}^2 \mathbf{v} - \nabla P + D_L \nabla (\nabla \cdot \mathbf{v}) + D_l (\mathbf{v} \cdot \nabla)^2 \mathbf{v} + f$$

nonconserved term **noise**

Steady state solution: $v^2 = \alpha/\beta \equiv v_0^2$

Symmetry breaking: $O(3) \rightarrow O(2)$

Fluctuation: $\mathbf{v} = (v_0 + \delta v_x, \delta v_y, \delta v_z)$



$\omega = ck$ $\omega = i\Gamma k^2$ **NG modes**
propagating diffusive

**Can we discuss symmetry breaking
without ordinary conservation law?**

Ex) Symmetry of Brownian motion

Langevin equation

$$\frac{d}{dt}x(t) = u(t)$$

$$\frac{d}{dt}u(t) = -\gamma u(t) + \xi(t)$$

$$\langle \xi_i(t) \xi_j(t') \rangle = 2\delta_{ij}\gamma T \delta(t - t')$$



Angular momentum $\mathbf{L} = \mathbf{x} \times \mathbf{u}$

$$\frac{d}{dt}\langle \mathbf{L}(t) \rangle = -\gamma \langle \mathbf{x} \times \mathbf{u}(t) \rangle \neq 0$$

not conserved

Langevin equation

$$\frac{d}{dt}u(t) = -\gamma u(t) + \xi(t)$$



Fokker-Planck equation

$$\partial_t P(t, u) = \frac{\partial}{\partial u_i} \left(\gamma T \frac{\partial}{\partial u_i} + \gamma u_i \right) P(t, u)$$



Path integral Martin-Siggia-Rose formalism

$$Z = \int \mathcal{D}\chi \mathcal{D}u e^{iS[\chi, u]}$$

Dynamic action: $iS = \int dt \left[i\chi_i \left(\frac{d}{dt}u_i + \gamma u_i \right) - T\gamma\chi_i^2 \right]$

Symmetry of Dynamic action

$$iS = \int dt \left[i\chi_i \left(\frac{d}{dt} u_i + \gamma u_i \right) - T\gamma \chi_i^2 \right]$$

O(3) symmetry

$$\chi_i \rightarrow R_{ij} \chi_j \quad u_i \rightarrow R_{ij} u_j \quad \text{with } R_{ik} R_{kj} = \delta_{jk}$$

Noether charge

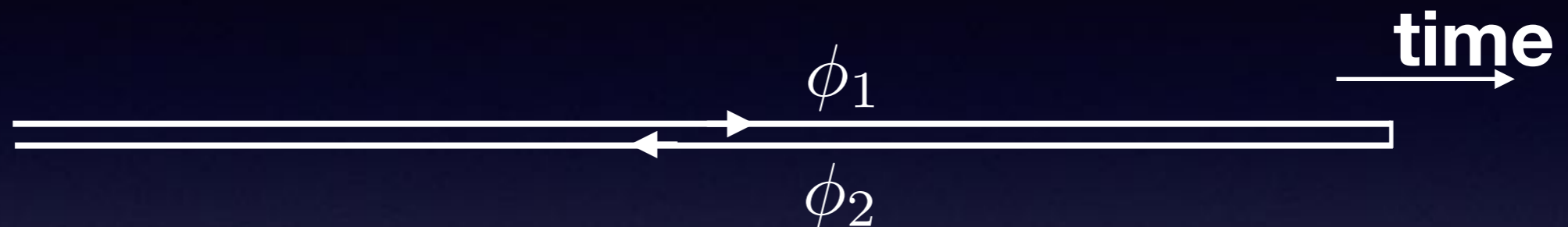
$$\mathbf{L}_{\text{MSR}} = \boldsymbol{\chi} \times \mathbf{u} \quad \mathbf{L} = \mathbf{x} \times \mathbf{u}$$

$$\mathbf{L}_{\text{MSR}} \neq \mathbf{L}$$

Open quantum system

cf. for review, Sieberer, Buchhold, Diehl, 1512.00637

Schwinger-Keldysh Path integral



$$Z = \int \mathcal{D}\phi_1 \mathcal{D}\phi_2 \exp \left[iS[\phi_1] - iS[\phi_2] + iS_{12}[\phi_1, \phi_2] \right]$$

complex

$S[\phi_1]$: **forward evolution**

$S[\phi_2]$: **backward evolution**

$S_{12}[\phi_1, \phi_2]$: **Interaction with environment**

Example

Lindblad equation

$$\partial_t \rho = -i[H, \rho] + \gamma \left(L \rho L^\dagger - \frac{1}{2} (L^\dagger L \rho + \rho L^\dagger L) \right)$$

fluctuation and dissipation

Action

$$S[\phi] = \frac{1}{2} (\partial_\mu \phi)^2 - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \phi^4$$

$$iS_{12}[\phi_1, \phi_2] = \gamma L(\phi_1) L^\dagger(\phi_2) - \frac{\gamma}{2} \left(L(\phi_1) L^\dagger(\phi_1) + L(\phi_2) L^\dagger(\phi_2) \right)$$

R/A basis

$$\phi_R = \frac{1}{2}(\phi_1 + \phi_2)$$

classical field

$$\phi_A = \phi_1 - \phi_2$$

fluctuation

$$S[\phi_1] - S[\phi_2] = \int d^4x \phi_A \left(-\partial_\mu^2 \phi_R - m^2 \phi_R - \frac{\lambda}{3!} \phi_R^3 \right) + \int d^4x \frac{\lambda}{24} \phi_A^3 \phi_R$$

classical Equation of motion

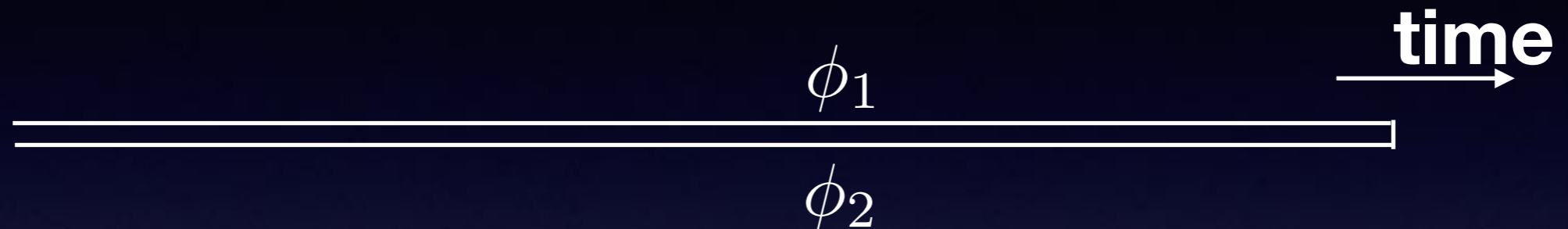
For example,

$$iS_{12}[\phi_R, \phi_A] = -\gamma \phi_A \partial_0 \phi_R - \frac{A}{2} \phi_A^2 + \dots$$

Symmetry of Open quantum system

cf. for review, Sieberer, Buchhold, Diehl, 1512.00637

Schwinger-Keldysh Path integral



$$Z = \int \mathcal{D}\phi_1 \mathcal{D}\phi_2 \exp \left[iS[\phi_1] - iS[\phi_2] + iS_{12}[\phi_1, \phi_2] \right]$$

complex

Q_1, Q_2 :Symmetry generators:

$S[\phi_1], S[\phi_2]$ are invariant.

Suppose $S_{12}[\phi_1, \phi_2]$ is invariant under $Q_A = \frac{Q_1 - Q_2}{2}$

We also define $Q_R = \frac{Q_1 + Q_2}{2}$

Spontaneous symmetry breaking

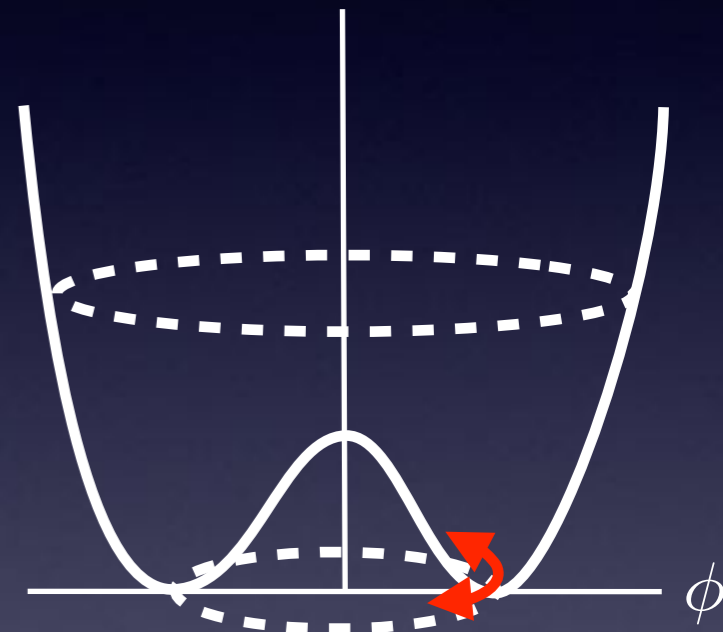
Minami, YH ('15)

Ex1) $SU(2) \times U(1)$ model Type-A $V(\phi)$

$$iS = \int d^4x \left[\phi_A^\dagger (-\partial_0^2 + \nabla^2 - \gamma \partial_0) \varphi_R - 2\lambda |\varphi_R|^2 \varphi_R - A \phi_A^\dagger \varphi_A \right] + \dots$$

$\varphi_{R/A}$ two component complex field

$$\varphi_R = (\pi_1 + i\pi_2, v + h + i\pi_3)$$



Linear analysis

$$(\partial_0^2 + \gamma \partial_0 - \nabla^2) \pi_a = 0 \quad \text{NG type-A mode}$$

$$\Rightarrow -\omega^2 - i\gamma\omega + k^2 = 0$$

$$\Rightarrow \omega = \frac{-i\gamma}{2} \pm \frac{i}{2} \sqrt{\gamma^2 - 4k^2} \sim -\frac{i}{\gamma} k^2, -i\gamma + \frac{i}{\gamma} k^2$$

diffusion mode

Spontaneous symmetry breaking

Minami, YH ('15)

Ex2) $SU(2) \times U(1)$ model Type-B

with chemical potential μ

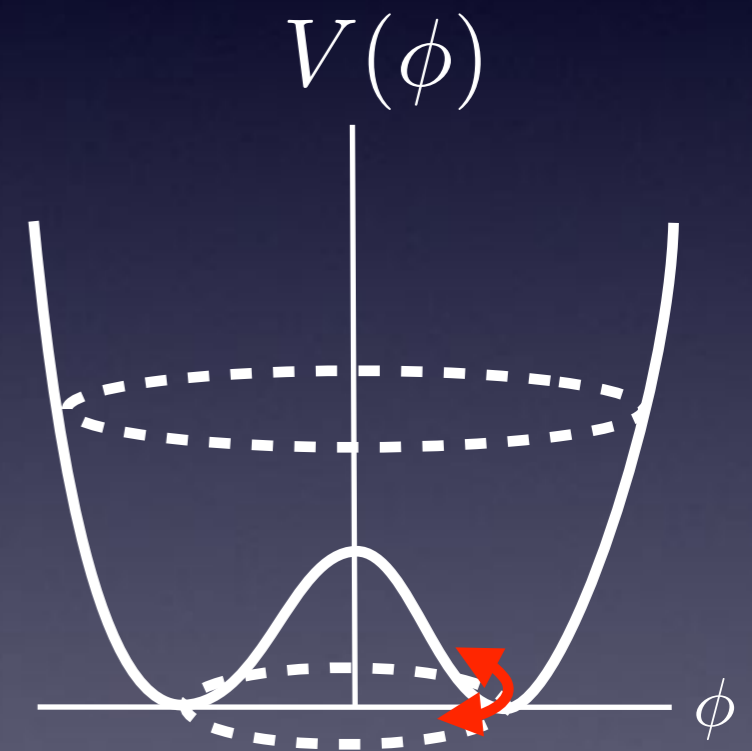
$$iS = \int d^4x \left[\phi_A^\dagger (-(\partial_0 + i\mu)^2 + \nabla^2 - \gamma\partial_0) \varphi_R - 2\lambda |\varphi_R|^2 \varphi_R - A \phi_A^\dagger \varphi_A \right] + \dots$$

$$\begin{pmatrix} -\partial_0^2 - \gamma\partial_0 + \nabla^2 & 2\mu\partial_0 \\ -2\mu\partial_0 & -\partial_0^2 - \gamma\partial_0 + \nabla^2 \end{pmatrix} \begin{pmatrix} \pi_1 \\ \pi_2 \end{pmatrix} = 0,$$

➔ $\omega = \frac{k^2}{4\mu^2 + \gamma^2} (\pm 2\mu - i\gamma)$

quadratic dispersion

$$\langle [iQ_A^1, Q_R^2] \rangle \neq 0 \quad \textbf{Type-B}$$



Spontaneous symmetry breaking

Minami, YH ('17)

Ex3) $SU(2) \times U(1)$ model Type-B

with complex potential

$$iS = \int d^4x \left(i\varphi_A^\dagger \left((-\partial_0^2 + \nabla^2 - (\gamma + 2i\mu)\partial_0 - m_r^2 - im_i^2) \varphi_R \right. \right. \\ \left. \left. - 2(\lambda_r + i\lambda_i)(\varphi_R^\dagger \varphi_R) \varphi_R \right) - A\varphi_A^\dagger \varphi_A \right) + \dots$$

Assuming $\varphi_R = (0, ve^{-i\omega_0 t})$

Gap equation

$$(\omega_0^2 - 2\mu\omega_0 - m_r^2 - 2\lambda_r v^2 + i(\gamma\omega_0 - m_i^2 - 2\lambda_i v^2))v = 0$$

Symmetric phase $v = 0$

Broken phase $v \neq 0, \quad \omega_0 \neq 0$

Spontaneous symmetry breaking

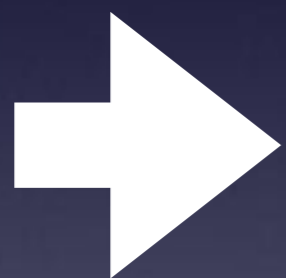
Minami, YH ('17)

Ex3) $SU(2) \times U(1)$ model Type-B

with complex potential

Linear analysis

$$\begin{pmatrix} -\partial_0^2 - \gamma\partial_0 + \nabla^2 & 2(\mu - \omega_0)\partial_0 \\ -2(\mu - \omega_0)\partial_0 & -\partial_0^2 - \gamma\partial_0 + \nabla^2 \end{pmatrix} \begin{pmatrix} \pi_1 \\ \pi_2 \end{pmatrix} = 0,$$



$$\omega = \frac{k^2}{4(\mu - \omega_0)^2 + \gamma^2} (\pm 2(\mu - \omega_0) - i\gamma)$$

We still have quadratic dispersion

Similarly, we find

$$\omega = -i \frac{k^2}{\gamma + 2(\mu - \omega_0)\lambda_i/\lambda_r}$$

Diffusive mode for type-A

Inverse propagator and dispersion

Minami, YH ('17)

$$[G_{\pi}^{-1}(k)]^{\beta\alpha} = iC^{\mu;\beta\alpha}k_{\mu} + C^{\mu\nu;\beta\alpha}k_{\mu}k_{\nu} + \cdots$$

Hamiltonian system

$$C^{\mu;\beta\alpha} = -\langle [iQ_R^{\alpha}, j_A^{\beta\mu}(0)] \rangle$$

$$-i \int d^D x \langle [iQ_R^{\alpha}, \mathcal{L}_{12}(x)] j_A^{\beta\mu}(0) \rangle_{\pi c}$$

$$C^{\mu\nu;\beta\alpha} = i \int d^D x \langle j_R^{\alpha\mu}(x) j_A^{\beta\nu}(0) \rangle_{\pi c}$$

$$- \lim_{k \rightarrow 0} \frac{\partial}{\partial k_{\nu}} i \int d^D x e^{ik_{\rho}x^{\rho}} \langle [iQ_R^{\alpha}, \mathcal{L}_{12}(x)] j_A^{\beta\mu}(0) \rangle_{\pi c}$$

Our result is too general

Need to impose symmetry of S_{12}

Ex)'Standard' Fokker-Plank eq.

Type-A mode

Type-B mode

Diffusive $\omega = -ik^2\Gamma$

$\omega = ak^2 - ik^2\Gamma'$

$$N_A = N_{BS} - \text{rank}\langle [iQ_R^\alpha, Q_A^\beta] \rangle$$

$$N_B = \frac{1}{2}\text{rank}\langle [iQ_R^\alpha, Q_A^\beta] \rangle$$

Next step: classification

Summary



**Spontaneous breaking of
symmetry of Dynamic action**

Two-type of diffusive NG modes

Type-A mode

Diffusive $\omega = -ik^2\Gamma$

Type-B mode

$\omega = ak^2 - ik^2\Gamma'$

Next step: classification

What is the condition satisfying this table?

Type	Dispersion		Conserved charge	Examples
	Re	Im		
A	k	k^2	Q_A, Q_R	Superfluid, etc.
	0	k^2	Q_A	flock of birds, Exciton-polariton condensates
B $\langle [Q_A, Q_R] \rangle \neq 0$	k^2	k^4	Q_A, Q_R	Ferromagnet
	k^2	k^2	Q_A	Spinor BEC in open quantum system? Magnetotactic bacteria?

Possible active matter with type B modes? Collective motion of Magnetotactic bacteria

