

Chiral Matter and Topology in Astrophysics

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“Recent Developments in Chiral Matter and Topology”
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Main topics

- Core-collapse supernova
- Chiral hydrodynamics
- Chiral turbulence in supernovae
- Photonic chiral vortical effect in pulsars

Units: $\hbar = c = k_B = e = 1$

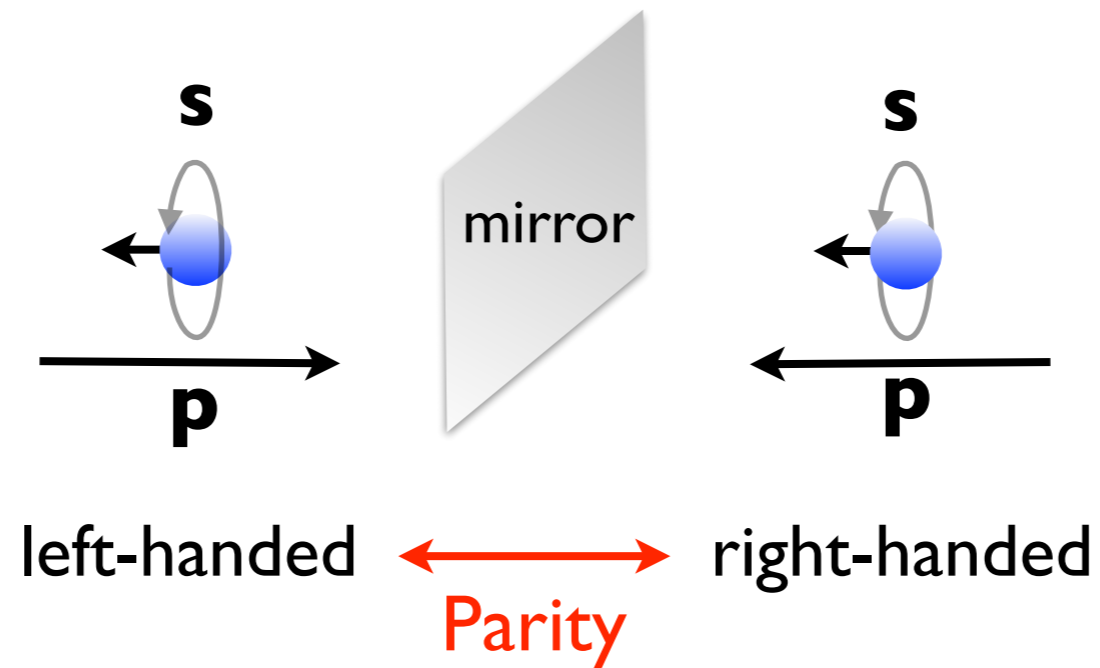
Core-collapse supernovae explosions

Core-collapse supernova explosions

- One of the most energetic phenomena in the Universe
- Transition to neutron stars & origin of heavy elements
- But explosion is difficult in conventional 3D hydrodynamic theory

One of the puzzles in astrophysics

Chirality of fermions



Why is “God” left-handed?

The laws of physics are left-right symmetric except for the **weak interaction** that acts only on **left-handed particles**.



W. Pauli

“God is just a **weak left-hander**.”

From **micro** to **macro**

Microscopic parity violation is reflected in **macroscopic** behavior:

Micro

Chirality of fermions (e, ν) in Standard Model



Chiral kinetic theory



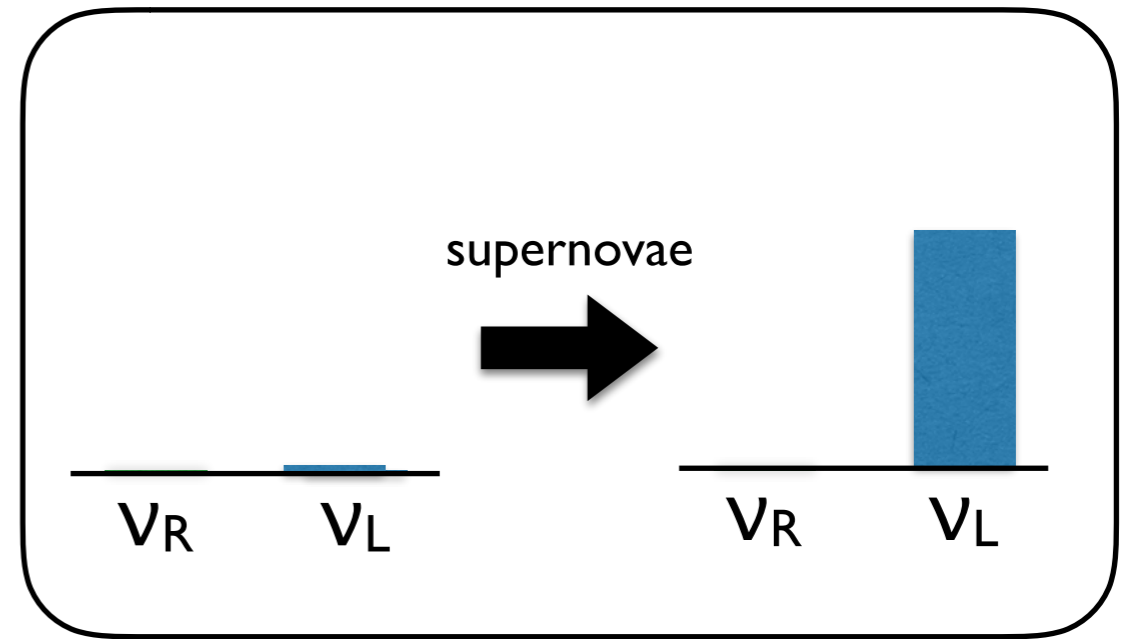
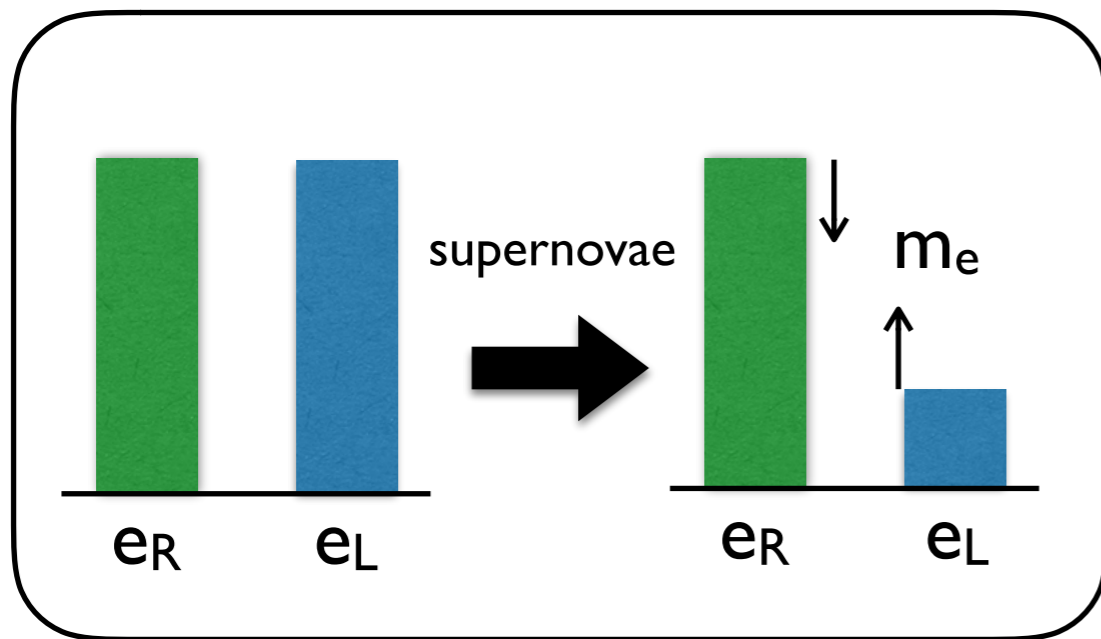
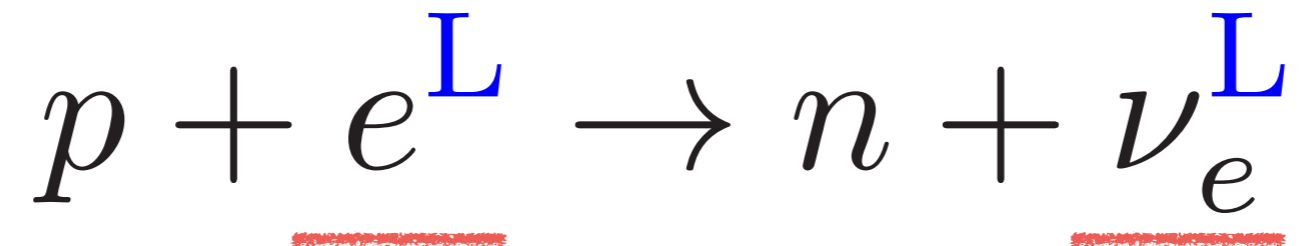
Son, Yamamoto (2012); Stephanov, Yin (2012); J. W. Chen, S. Pu, Q. Wang, X. N. Wang (2013), ...

Macro

Evolution of core-collapse supernovae (**giant P violation**)

Yamamoto (2016)

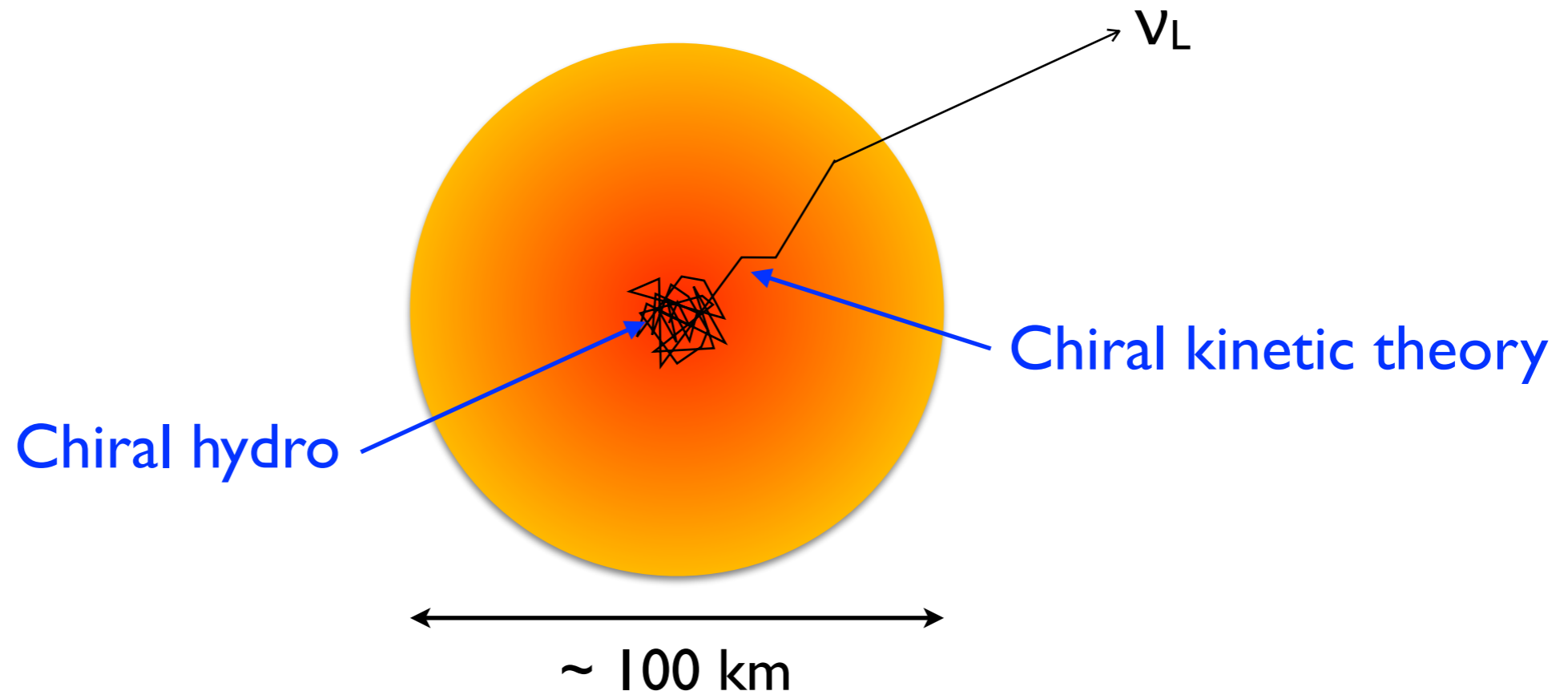
Supernova = Giant Parity Breaker



Ohnishi, Yamamoto (2014); Grabowska, Kaplan, Reddy (2015); Sigl, Leite (2016), ...

Neutrino matter in supernovae

- Neutrino mean free path ~ 1 cm at core ($\rho_N \sim 10^{15}$ g/cm³).
- Neutrino matter = **Chiral liquid** ($\mu_\nu \sim 200$ MeV \gg $T \sim 10$ MeV)
= **3D topological matter**



Chiral hydrodynamics

Chiral magnetic effect

$$\mathbf{j} = \frac{\mu_R - \mu_L}{4\pi^2} \mathbf{B} \equiv \frac{\mu_5}{2\pi^2} \mathbf{B}$$

$$\mathbf{j}_5 = \frac{\mu_R + \mu_L}{4\pi^2} \mathbf{B} \equiv \frac{\mu}{2\pi^2} \mathbf{B}$$

Vilenkin (1980); Nielsen, Ninomiya (1983); Fukushima, Kharzeev, Warringa (2008), ...

Chiral vortical effect

$$\mathbf{j} = \frac{\mu\mu_5}{2\pi^2}\boldsymbol{\omega}$$

$$\mathbf{j}_5 = \left(\frac{\mu^2 + \mu_5^2}{4\pi^2} + \frac{T^2}{12} \right) \boldsymbol{\omega}$$

vorticity $\boldsymbol{\omega} \equiv \nabla \times \mathbf{v}$

Vilenkin (1979); Erdmenger et al. (2009); Banerjee et al. (2011);
Son, Surowka (2009); Landsteiner et al. (2011)

Lorentz covariant chiral hydro

Energy-momentum conservation: $\partial_\mu T^{\mu\nu} = F^{\nu\lambda} j_\lambda$

Anomaly relation: $\partial_\mu j_5^\mu = C E^\mu B_\mu$

$$T^{\mu\nu} = (\epsilon + P) u^\mu u^\nu - P g^{\mu\nu} + (\text{dissipation})$$

$$j^\mu = n u^\mu + \xi_B B^\mu + \xi \omega^\mu + (\text{dissipation})$$

$$j_5^\mu = n_5 u^\mu + \xi_{B5} B^\mu + \xi_5 \omega^\mu + (\text{dissipation})$$

$$B^\mu = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} u_\nu F_{\alpha\beta}, \quad \omega^\mu = \epsilon^{\mu\nu\alpha\beta} u_\nu \partial_\alpha u_\beta$$

Son, Surowka (2009); Sadofyev, Isachenkov (2011); Neiman, Oz (2011)

Helicity conservation

Yamamoto (2016); see also Avdoshkin et al., (2016)

$$\partial_{\mu} j_5^{\mu} = C \mathbf{E} \cdot \mathbf{B}$$

Helicity conservation

Yamamoto (2016); see also Avdoshkin et al., (2016)

$$\frac{d}{dt} \int d^3 \boldsymbol{x} \left(j_5^0 + \frac{C}{2} \boldsymbol{A} \cdot \boldsymbol{B} \right) = 0$$

Helicity conservation

Yamamoto (2016); see also Avdoshkin et al., (2016)

$$\frac{d}{dt} \int d^3 \boldsymbol{x} \left(\underline{j_5^0} + \frac{C}{2} \boldsymbol{A} \cdot \boldsymbol{B} \right) = 0$$

CVE

CME

$$j_5^0 = n_5 + \xi_5 \boldsymbol{v} \cdot \boldsymbol{\omega} + \xi_{B5} \boldsymbol{v} \cdot \boldsymbol{B}$$

Helicity conservation

Yamamoto (2016); see also Avdoshkin et al., (2016)

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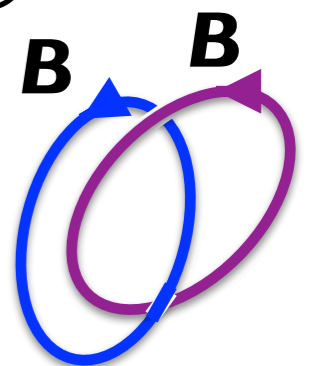
$$\frac{d}{dt} Q_{\text{tot}} = 0, \quad Q_{\text{tot}} \equiv Q_{\text{chi}} + Q_{\text{mag}} + Q_{\text{flu}} + Q_{\text{mix}}$$

chiral charge

$$Q_{\text{chi}} = \int d^3 \mathbf{x} n_5$$

magnetic helicity

$$Q_{\text{mag}} = \int d^3 \mathbf{x} \frac{C}{2} \mathbf{A} \cdot \mathbf{B}$$

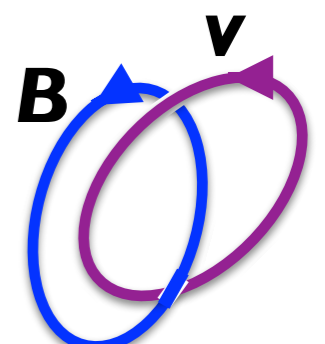
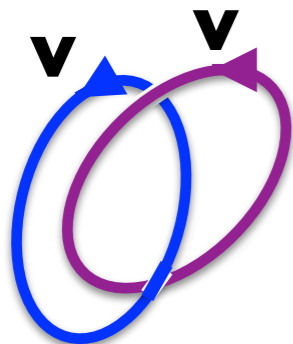


fluid helicity

$$Q_{\text{flu}} = \int d^3 \mathbf{x} \xi_5 \mathbf{v} \cdot \boldsymbol{\omega}$$

mixed helicity

$$Q_{\text{mix}} = \int d^3 \mathbf{x} \xi_{B5} \mathbf{v} \cdot \mathbf{B}$$



Neutrino chiral hydro

- Chiral hydrodynamic equations for pure neutrino matter:

$$(\epsilon + P)(\partial_t + \mathbf{v} \cdot \nabla)\mathbf{v} = -\nabla P + \nu \nabla^2 \mathbf{v}$$

$$\partial_t \underbrace{(n + \xi \mathbf{v} \cdot \boldsymbol{\omega})}_{\text{CVE}} + \nabla \cdot \mathbf{j} = 0, \quad \mathbf{j} = n\mathbf{v} + \underbrace{\xi \boldsymbol{\omega}}_{\text{CVE}}$$

- Neutrino number + fluid helicity is conserved.
- Generation of fluid helicity is numerically observed.

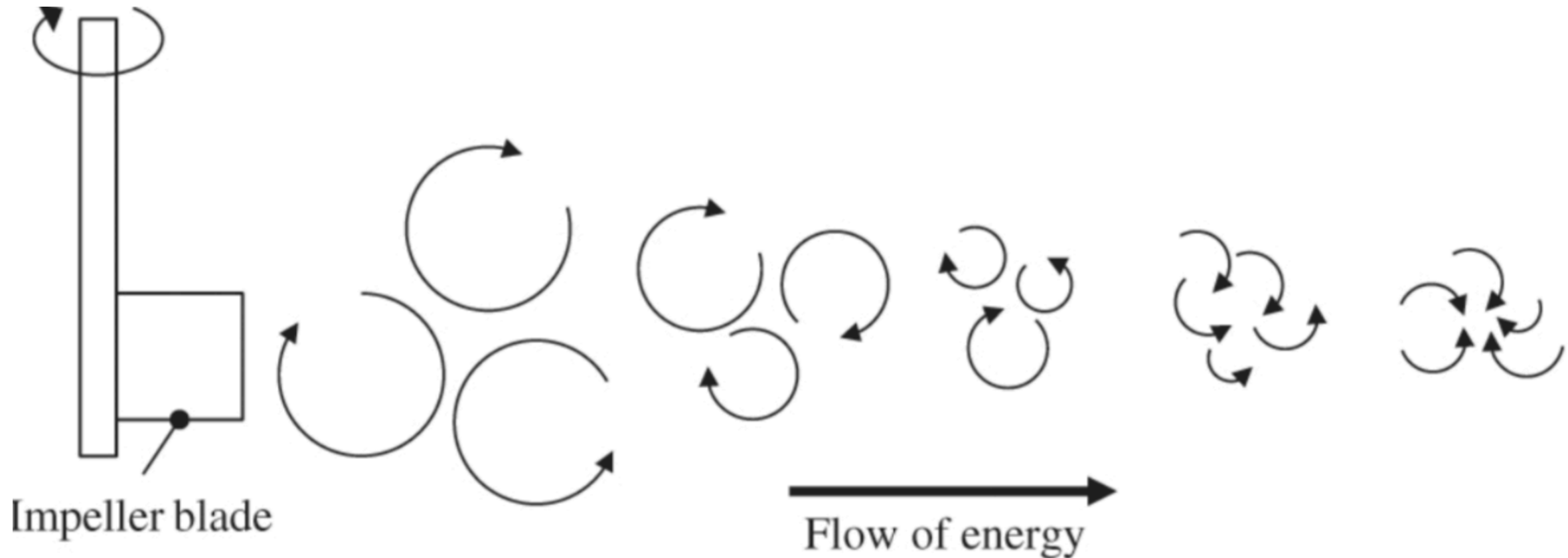
Kobayashi, Okuno, Yamamoto, in preparation

- When coupled to charged sector, fluid helicity $\sim \mu_5$ for electrons

$$\mathbf{j} \sim (\mathbf{v} \cdot \boldsymbol{\omega}) \mathbf{B}$$

Chiral MHD turbulence in supernovae

Turbulence and cascade

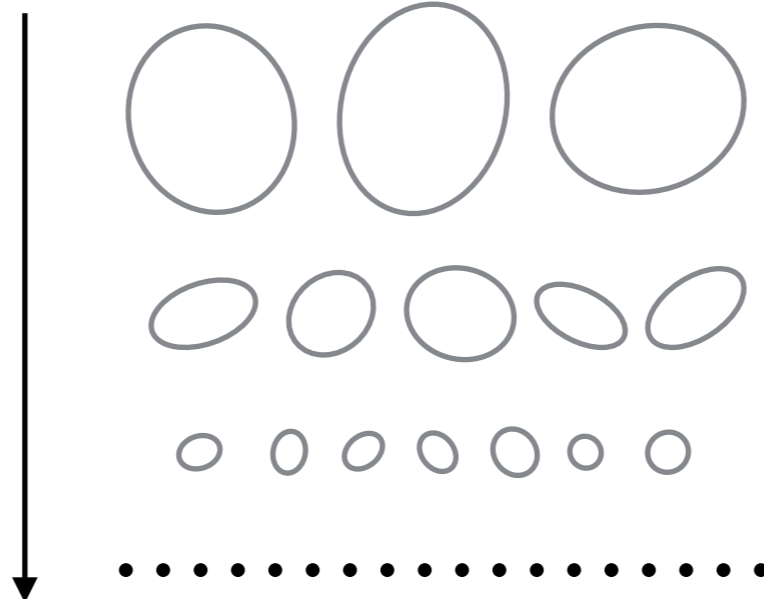
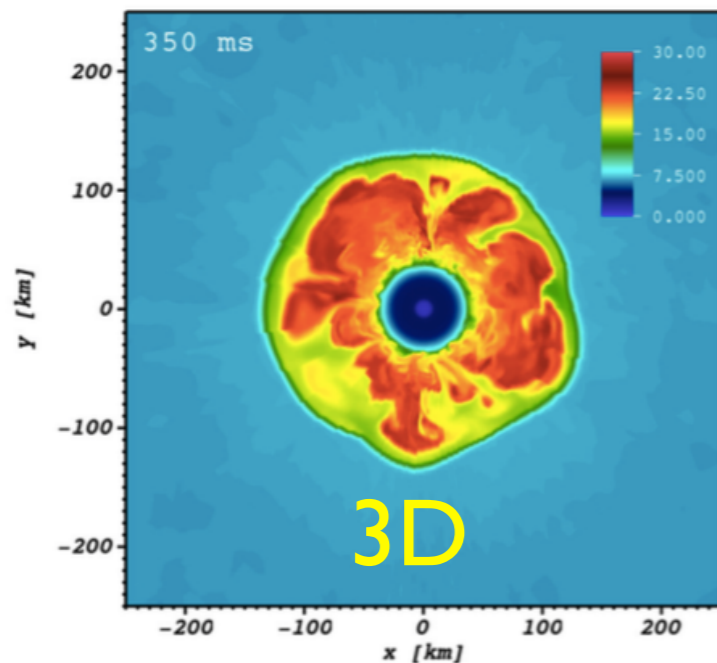


<https://doi.org/10.1515/htmp-2016-0043>

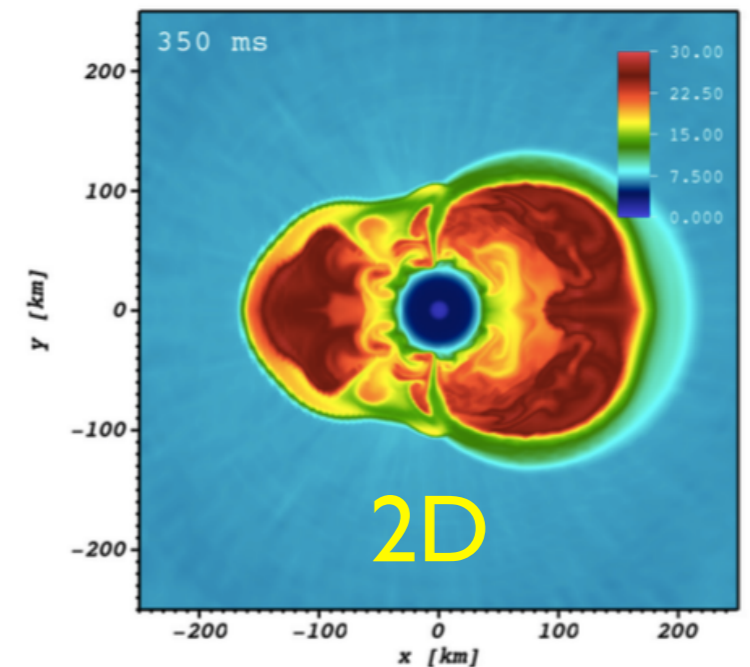
- The structure becomes smaller, and eventually dissipates (**direct cascade**)
- Similar in magneto-hydrodynamics (MHD)

Cascade and explosion

Direct cascade
(3D usual matter)
explosion difficult



Inverse cascade
(2D usual matter)
explosion easier

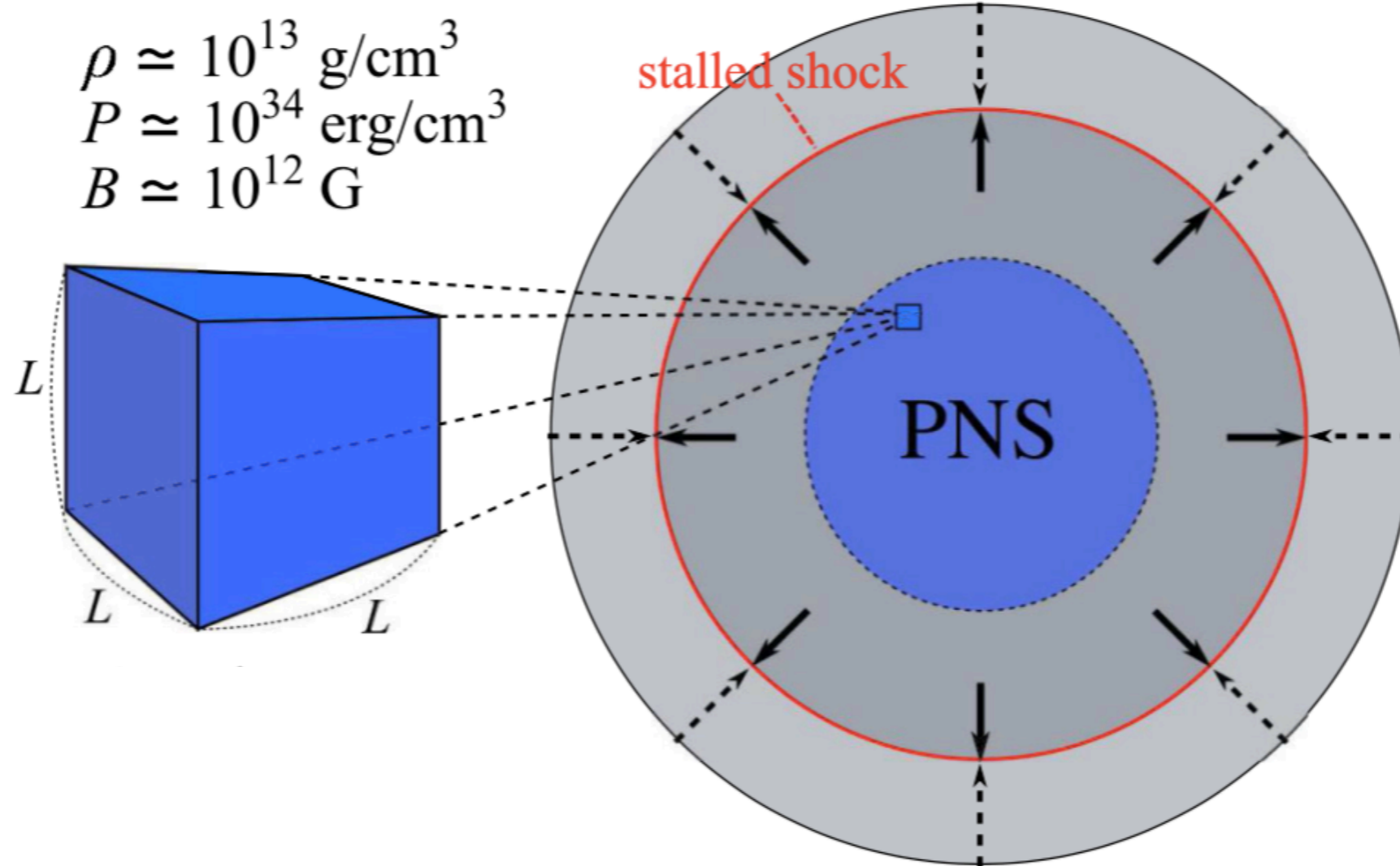


F. Hanke (2014)

What about 3D chiral matter?

Chiral MHD for supernovae

Masada, Kotake, Takiwaki, Yamamoto, arXiv:1805.10419



Proto-neutron star (PNS)

Chiral MHD for supernovae

Masada, Kotake, Takiwaki, Yamamoto, arXiv:1805.10419

- Chiral MHD w/o vorticity at the core (proton, e_R , e_L):

$$\partial_t \rho + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$\partial_t (\rho \mathbf{v}) + \nabla \cdot (\rho \mathbf{v} \mathbf{v}) = -\nabla P + \mathbf{J} \times \mathbf{B} + (\text{dissipation})$$

$$\partial_t \mathbf{B} = \nabla \times (\mathbf{v} \times \mathbf{B}) + \eta \nabla^2 \mathbf{B} + \eta \nabla \times \underbrace{(\xi_B \mathbf{B})}_{\text{"CME"}}$$

$$\partial_t n_5 = \frac{\eta}{2\pi^2} \underbrace{(\nabla \times \mathbf{B} - \xi_B \mathbf{B}) \cdot \mathbf{B}}_{\text{chiral anomaly}}$$

- Setup for proto-neutron stars (100 MeV = 1) :

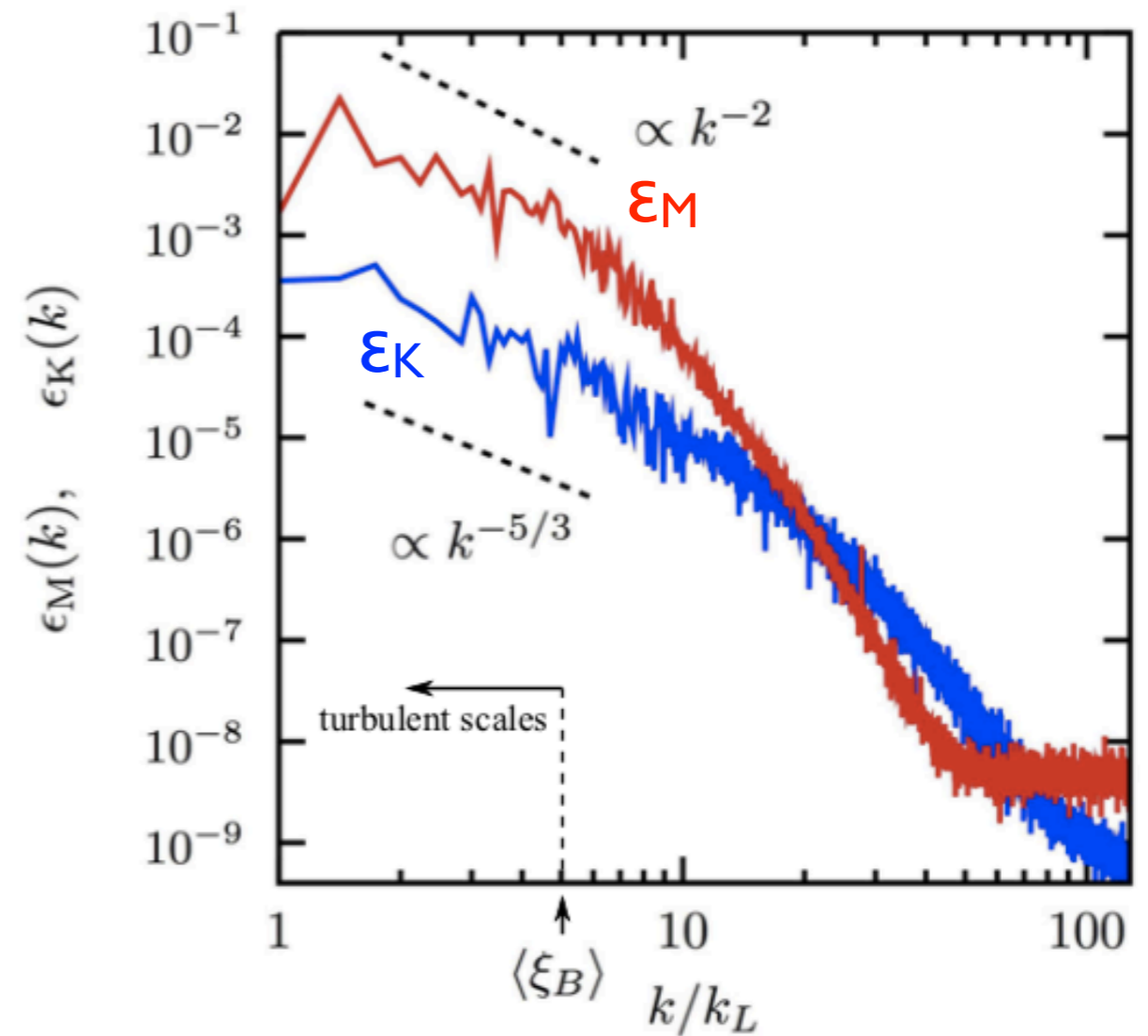
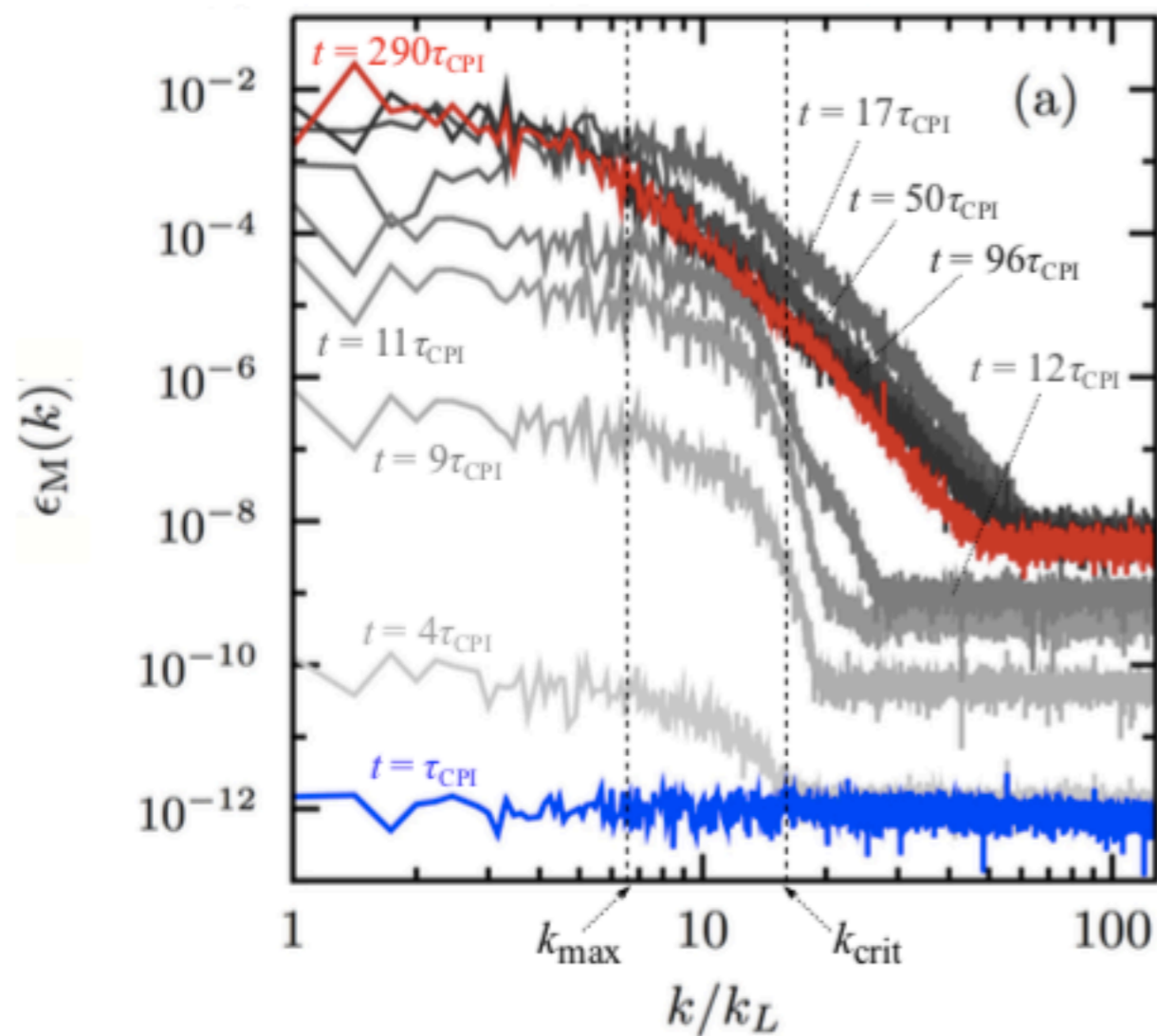
$$\rho_0 = 5.0, \quad P_0 = 1.0, \quad \xi_{B0} = 4.2 \times 10^{-3}, \quad \eta = 100.0$$

Movies of 3D simulations are available at:

<http://www.kusastro.kyoto-u.ac.jp/~masada/movie.mp4>

Masada, Kotake, Takiwaki, Yamamoto, arXiv:1805.10419

Energy spectra



- As time passes, energy in small- k and large- k regions grows
- Eventually, $\epsilon_M \sim k^{-2}$, $\epsilon_K \sim k^{-5/3}$

Masada et al., arXiv:1805.10419; see also Brandenburg et al., arXiv:1707.03385

Neutrino **chiral** radiation hydro

Yamamoto, work in progress

$$\nabla_{\alpha} (T_{\text{hyd}}^{\alpha\beta} + T_{\nu}^{\alpha\beta}) = 0$$

Stress tensor for N & e

(Hydro)

Stress tensor for ν

(Chiral kinetic theory)

$$T_{\nu}^{ij} = \int_{\mathbf{p}} |\mathbf{p}| \left(\hat{p}^i \hat{p}^j n_{\nu} - \frac{1}{2} p^i \epsilon^{jkl} \Omega_{\mathbf{p}}^k \partial_{\ell} n_{\nu} - \frac{1}{2} p^j \epsilon^{ikl} \Omega_{\mathbf{p}}^k \partial_{\ell} n_{\nu} \right)$$

Berry curvature of ν : $\Omega_{\mathbf{p}} = -\frac{\hat{\mathbf{p}}}{2|\mathbf{p}|^2}$

Photonic chiral vortical effect

Avkhadiev-Sadofiev (2017); Yamamoto (2017); V.A. Zyuzin (2017);
Chernodub, Cortijo, Landsteiner (2018), ...

Helicity and Berry curvature

- Spin-momentum locking \Leftrightarrow helicity λ
 - chiral fermions ($\lambda=\pm 1/2$)
 - photons ($\lambda=\pm 1$) e.g., Onoda, Murakami, Nagaosa (2004)
 - gravitons ($\lambda=\pm 2$) Yamamoto (2018)
- Berry curvature (adiabatic approximation): $\Omega_p = \lambda \frac{\hat{p}}{|\mathbf{p}|^2}$

Photon gas under rotation

Yamamoto, arXiv:1702.08886

- Semi-classical equations of motion in a rotating frame:

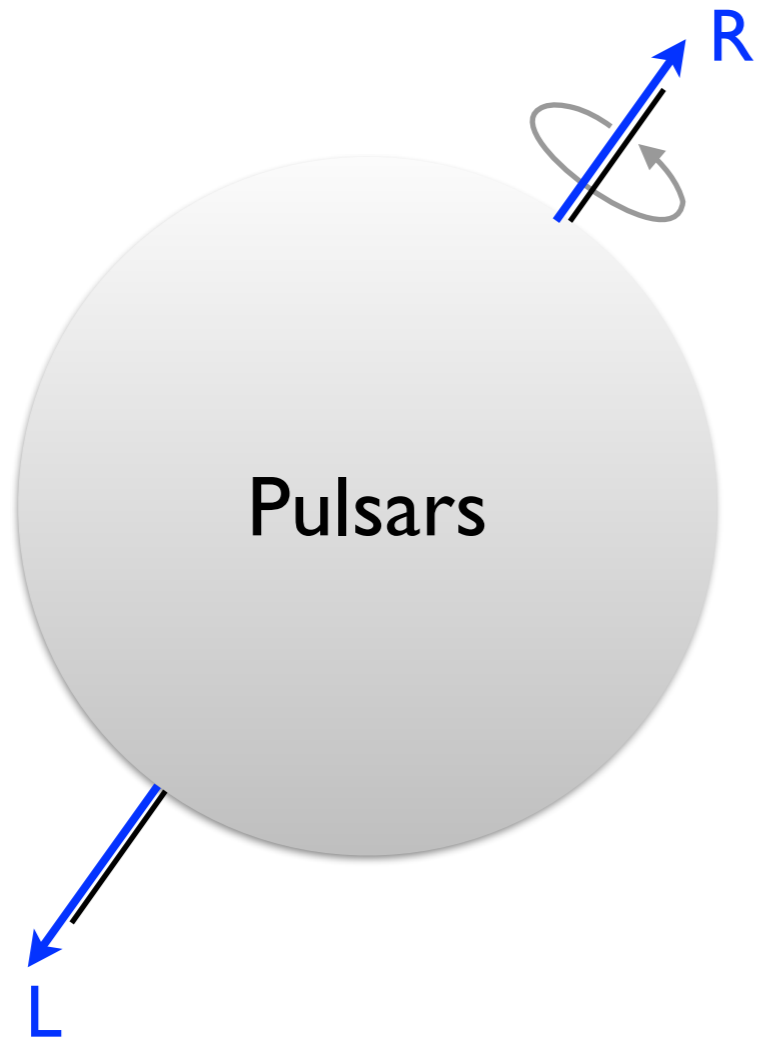
$$\begin{aligned} \dot{\mathbf{x}} &= \hat{\mathbf{p}} + \dot{\mathbf{p}} \times \boldsymbol{\Omega}_p & \longrightarrow & \sqrt{G} \dot{\mathbf{x}} = \hat{\mathbf{p}} + 2\omega |\mathbf{p}| (\hat{\mathbf{p}} \cdot \boldsymbol{\Omega}_p) \\ \dot{\mathbf{p}} &= 2|\mathbf{p}| \dot{\mathbf{x}} \times \boldsymbol{\omega} + O(\omega^2) & & G = (1 + 2|\mathbf{p}| \boldsymbol{\omega} \cdot \boldsymbol{\Omega}_p)^2 \\ & \text{Coriolis force} & & \end{aligned}$$

- Photonic chiral current along a rotation:

$$j_{\text{CVE}}^{\pm} = 2\omega \int \frac{d^3\mathbf{p}}{(2\pi)^3} |\mathbf{p}| (\hat{\mathbf{p}} \cdot \boldsymbol{\Omega}_p) n_{\mathbf{p}}^{\pm} = \pm \frac{T^2}{6} \omega$$

non-equilibrium equilibrium

X-ray pulsars



$$T \sim 10 \text{ keV}, \omega \sim 10^3 \text{ Hz}$$

→ Polarized photon flux: $f^{\pm} \sim 10^{21} / \text{s} \cdot \text{cm}^2$

cf) photon flux from sun: $f_{\odot} \sim 10^{17} / \text{s} \cdot \text{cm}^2$

Conclusion

- Chiral effects of e & ν may help the supernova explosion
- Photonic chiral vortical effect in pulsars
- Quantum correction to gravit. lensing of gravit. waves
~ Berry curvature

Yamamoto, arXiv:1708.03113