# Are the standard-model parameters free?

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#### Standard Model



physicists are curious about flavor structure: mass hierarchy, mixing patterns,... puzzles for decades Textbooks: Chapter 15 in Quarks & Leptons by Halzen and Martin: .... the (fermion) masses depend on the arbitrary (Yukawa) couplings and cannot be predicted.

ChatGPT:

The Standard Model parameters are free in the sense that their values cannot be determined by the theory alone, and experimental measurements play a crucial role in determining their values. To explain them, need new physics, but...

Today's talk:

At least some of the SM parameters are not free, but constrained dynamically via analyticity for internal consistency

## Your discretion is advised

#### Higgs mechanism Before and after symmetry breaking VEV v physical world $V(\varphi)$ $arphi = rac{1}{\sqrt{2}}$ , 3X3 Yukawa matrix lefthanded $(\overline{Q_L}Y_u u_R \varphi + \overline{Q_L}Y_d d_R \tilde{\varphi})$ doublet (u,c,t) (d,s,b)d(c)

- Massless particles → massive particles
- Flavor changing via Yukawa couplings 🔿 diagonal Yukawa matrices
- Quarks in weak eigenstates → quark mixing

$$\begin{array}{l} \begin{array}{l} \mathsf{CKM\ matrix}\\ \mathsf{for\ weak\ int.} & \begin{bmatrix} d'\\s'\\b' \end{bmatrix} = \begin{bmatrix} V_{\mathrm{ud}} & V_{\mathrm{us}} & V_{\mathrm{ub}}\\ V_{\mathrm{cd}} & V_{\mathrm{cs}} & V_{\mathrm{cb}}\\ V_{\mathrm{td}} & V_{\mathrm{ts}} & V_{\mathrm{tb}} \end{bmatrix} \begin{bmatrix} d\\s\\b \end{bmatrix}, \quad \sum_{k} |V_{ik}|^2 = \sum_{k} |V_{ki}|^2 = 1\\ \sum_{k} |V_{ki}|^2 = 0. \end{array}$$
$$\begin{array}{l} \begin{array}{l} \mathbf{unitarity}\\ \mathbf{unitarity}\\ \end{bmatrix} \\ \begin{bmatrix} 1 & 0 & 0\\0 & c_{23} & s_{23}\\0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{13}}\\0 & 1 & 0\\-s_{13}e^{i\delta_{13}} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0\\-s_{12} & c_{12} & 0\\0 & 0 & 1 \end{bmatrix} \\ \begin{array}{l} \begin{array}{l} \mathbf{c}_{12} & s_{12} & 0\\0 & 0 & 1 \end{bmatrix} \\ \mathbf{c}_{12} & c_{23} - c_{12} s_{23} s_{13} e^{i\delta_{13}} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta_{13}} & s_{23} c_{13}\\s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta_{13}} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta_{13}} & c_{23} c_{13}\\s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta_{13}} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta_{13}} & c_{23} c_{13}\\s_{13} (1 - \rho - i\eta) & -\lambda\lambda^2 & 1 \end{bmatrix} + O(\lambda^4) \,. \end{array}$$

#### Gauge interaction

• Gauge interaction like QED

**4-vector potential** generators  

$$D_{\mu}\psi(x) = \left(\partial_{\mu} - ig A^{a}_{\mu}(x) t^{a}\right)\psi(x)$$

- Vector interaction, photon has spin 1
- Gauge group fixes  $t^a$ : 1 for QED, U(1) group; Pauli matrices for weak interaction, SU(2) group; Gell-Mann matrices for QCD, SU(3) group
- Generators describe basic transformation, e.g.

Red = (1,0,0)  
Blue= (0,1,0) 
$$\lambda_1 = \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$
 Red  $\rightarrow$  Blue

• Only overall coefficient is free, single coupling g

#### Scalar interaction

• No such symmetry constraint on scalar coupling



- Yukawa couplings are arbitrary, so are quark masses, mixing matrix!
- But this observation made at Lagrangian level without considering dynamical interplay among various particles

#### Dispersion relation

• Example: mixing of D meson with mass squared s



 $D^0$ 

 $D^0$ 

#### Observations

- Fundamental parameters in theory (like Standard Model) usually constrained by symmetries at Lagrangian level
- Analyticity is crucial property of physical observables
- $\Gamma_{12}$  involves CKM matrix elements and fermion masses
- Additional dynamical constraints imposed by dispersion relations, if  $M_{12}$  is known ?
- Turn out that dispersive constraints are so strong that Yukawa couplings in SM are in fact not free parameters

#### Idea

 Neutral state mixing disappears at high energy, where electroweak symmetry is restored



#### Proof of $M_{12}(s) \approx 0$

- Consider mixing of  $Q_L ar{q}_L$  ,  $ar{Q}_L q_L$  neutral states
- Before breaking, all particles are massless, quarks in flavor eigenstates
- Mixing occurs via exchanges of charged or neutral scalars, whose strengths described by Yukawa matrices



- After breaking, particles get masses, quarks turned to mass eigenstates
- Mixing occurs via W boson exchanges, whose strengths described by CKM matrix

#### Mixing in symmetric phase

- Yukawa interaction  $\overline{Q_L}Y_u u_R \varphi + \overline{Q_L}Y_d d_R \tilde{\varphi} \qquad \varphi = \frac{1}{\sqrt{2}} \begin{pmatrix} \varphi^+ \\ \varphi^0 \end{pmatrix}$ , left-handed doublet
- In symmetric phase, implement quark field transformation adopted in broken phase  $u_L \rightarrow U_u u_L$   $u_R \rightarrow V_u u_R$   $d_L \rightarrow U_d d_L$   $d_R \rightarrow V_d d_R$
- Yukawa matrices diagonalized, but charged scalar currents exist
- down-type quarks, coupling to up-type quarks in mass eigenstates through charged scalar currents, are not in mass eigenstates



#### High-energy input

- Heavy quark Q provides large s in box diagrams. Symmetry restores and intermediate particles become massless,  $M_{12}(s) \approx 0$
- s' can be low, so  $\Gamma_{12}(s')$  depends on CKM matrix elements associated with massive intermediate quarks in broken phase.

Cheng 1982 Buras et al 1984

• Box-diagram contribution

$$\Gamma_{12}(s) \propto \sum_{i,j} \lambda_i \lambda_j \Gamma_{ij}(s),$$

$$\Gamma_{ij}(s) = \frac{1}{s^2} \frac{\sqrt{s^2 - 2s(m_i^2 + m_j^2) + (m_i^2 - m_j^2)^2}}{(m_W^2 - m_i^2)(m_W^2 - m_j^2)}$$

$$\times \left\{ \left( m_W^4 + \frac{m_i^2 m_j^2}{4} \right) [2s^2 - 4s(m_i^2 + m_j^2) + 2(m_i^2 - m_j^2)^2] + 3m_W^2 s(m_i^2 + m_j^2)(m_i^2 + m_j^2 - s) \right\}$$

for D mixing i, j = d, s, b  $\lambda_i \equiv V_{ci}^* V_{ui}$ 

#### Constraints

• How to diminish dispersive integral

$$\int ds' \frac{\Gamma_{12}(s')}{s-s'} ?$$

 $\lambda_i \lambda_j g_{ij} \approx 0$ 

• Asymptotic expansion

$$\Gamma_{ij}(s') \approx \Gamma_{ij}^{(1)}s' + \Gamma_{ij}^{(0)} + \frac{\Gamma_{ij}^{(-1)}}{s'} + \cdots \qquad \text{EW symmetry} \\ \text{restoration scale} \\ \Gamma_{ij}^{(1)} = \frac{4m_W^4 - 6m_W^2(m_i^2 + m_j^2) + 4m_i^2m_j^2}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}, \implies \Lambda^2/s \\ \Gamma_{ij}^{(0)} = -\frac{3(m_i^2 + m_j^2) \left[4m_W^4 - 4m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)} \implies (m_i^2 + m_j^2)\Lambda/s \\ \Gamma_{ij}^{(-1)} = \frac{3(m_i^4 + m_j^4) \left[4m_W^4 - 2m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}. \implies (m_i^4 + m_j^4) \ln \Lambda/s \\ \Gamma_{ij}^{(-1)} = \frac{3(m_i^4 + m_j^4) \left[4m_W^4 - 2m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}. \implies (m_i^4 + m_j^4) \ln \Lambda/s \\ \text{to diminish integral}$$

$$\int ds' \frac{\Gamma_{12}(s')}{s-s'} \approx \frac{1}{s} \sum_{i,j} \lambda_i \lambda_j g_{ij} \qquad g_{ij} \equiv \int_{t_{ij}}^{\infty} ds' \left[ \Gamma_{ij}(s') - \Gamma_{ij}^{(1)}s' - \Gamma_{ij}^{(0)} - \frac{\Gamma_{ij}^{(-1)}}{s'} \right]$$

#### Minimization

• Rewrite constrains  $r^2 R_{dd}^{(m)} + 2r R_{ds}^{(m)} + 1 \approx 0,, \quad m = 1, 0, -1, i$ 

$$R_{dd}^{(m)} = \frac{\Gamma_{dd}^{(m)} - 2\Gamma_{db}^{(m)} + \Gamma_{bb}^{(m)}}{\Gamma_{ss}^{(m)} - 2\Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}}, \quad R_{ds}^{(m)} = \frac{\Gamma_{ds}^{(m)} - \Gamma_{db}^{(m)} - \Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}}{\Gamma_{ss}^{(m)} - 2\Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}} \qquad m = 1, 0, -1$$

- Expression for m = i similar, but with  $g_{ij}$
- Ratio of CKM elements  $r = \frac{\lambda_d}{\lambda_s} = \frac{V_{cd}^* V_{ud}}{V_{cs}^* V_{us}} \equiv u + iv,$
- Tune u and v to minimize the sum (real parts of constraints)

$$\sum_{m=1,-1,i} \left[ (u^2 - v^2) R_{dd}^{(m)} + 2u R_{ds}^{(m)} + 1 \right]^2$$



$$r = \frac{V_{cd}^* V_{ud}}{V_{cs}^* V_{us}} = -1.0 + (6.2^{+1.2}_{-1.0}) \times 10^{-4} i \qquad u = -1.00029 \pm 0.00002, \qquad v = 0.00064 \pm 0.00002$$
  
variation of ms by 0.01 GeV they agree well

#### Analytical solution

• Insert u=-1 into m=1 constraint to get analytical expression of v

$$v \approx \frac{(m_W^2 - m_b^2)(m_s^2 - m_d^2)}{(m_W^2 - m_s^2)(m_b^2 - m_d^2)} \approx \frac{m_s^2}{m_b^2}$$

- In terms of Wolfenstein parameters  $v = A^2 \lambda^4 \eta$  Ahn et al, 2011
- Produce well-known numerical relation

$$\lambda = V_{us} \approx (A^2 \eta)^{-1/4} \sqrt{\frac{m_s}{m_b}} \approx \sqrt{\frac{m_s}{m_b}}$$

 $A \approx 0.826$   $\eta \approx 0.348$  $(A^2 \eta)^{-1/4} \approx 1.43 \sim O(1)$ 

Belfatto et al, 2023

Cheng, Sher 1987

#### Lepton mixing

- Pontecorvo–Maki–Nakagawa–Sakata
- Apply the same formalism to lepton  $\mu^-e^+-\mu^+e^-$  mixing through similar box diagrams with intermediate neutrino channels
- Correspondence  $m_{d,s,b} \leftrightarrow m_{1,2,3}$   $V_{cd}^*V_{ud}/(V_{cs}^*V_{us}) \leftrightarrow r = U_{\mu 1}^*U_{e1}/(U_{\mu 2}^*U_{e2})$
- Normal hierarchy (NH)  $m_1^2 = 10^{-6} \text{ eV}^2$  de Salas et al, 2018
  - $\Delta m_{21}^2 \equiv m_2^2 m_1^2 = (7.55^{+0.20}_{-0.16}) \times 10^{-5} \text{ eV}^2 \qquad \Delta m_{32}^2 \equiv m_3^2 m_2^2 = (2.424 \pm 0.03) \times 10^{-3} \text{ eV}^2$
- Predict  $r = \frac{U_{\mu 1}^* U_{e1}}{U_{\mu 2}^* U_{e2}} \approx -1.0 - 0.02i$   $r = -(0.738_{-0.048}^{+0.050}) - (0.179_{-0.125}^{+0.136})i$
- Inverted hierarchy (IH)  $r \approx -1.0 O(10^{-5})i$   $r = -(1.03^{+0.05}_{-0.16}) (0.356^{+0.015}_{-0.048})i$
- Quasi-degenerate  $r \approx -0.97 O(10^{-5})i$
- NH and observed PMNS matrix satisfy constraint at order of magnitude

#### PDG



	Ref. [188] w/o SK-ATM		Ref. [188] w SK-ATM		Ref. [189] w SK-ATM		Ref. [190] w SK-ATM	
NO	Best Fit Ordering		Best Fit Ordering		Best Fit Ordering		Best Fit Ordering	
Param	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range
$\frac{\sin^2 \theta_{12}}{10^{-1}}$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.04_{-0.13}^{+0.14}$	$2.65 \rightarrow 3.46$	$3.20^{+0.20}_{-0.16}$	$2.73 \rightarrow 3.79$
$\theta_{12}/^{\circ}$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.46^{+0.87}_{-0.88}$	$30.98 \rightarrow 36.03$	$34.5^{+1.2}_{-1.0}$	$31.5 \rightarrow 38.0$
$\frac{\sin^2 \theta_{23}}{10^{-1}}$	$5.58^{+0.20}_{-0.33}$	$4.27 \rightarrow 6.09$	$5.63^{+0.18}_{-0.24}$	$4.33 \rightarrow 6.09$	$5.51^{+0.19}_{-0.80}$	$4.30 \rightarrow 6.02$	$5.47^{+0.20}_{-0.30}$	$4.45 \rightarrow 5.99$
$\theta_{23}/^{\circ}$	$48.3^{+1.2}_{-1.9}$	$40.8 \rightarrow 51.3$	$48.6^{+1.0}_{-1.4}$	$41.1 \rightarrow 51.3$	$47.9^{+1.1}_{-4.0}$	$41.0 \rightarrow 50.9$	$47.7^{+1.2}_{-1.7}$	$41.8 \rightarrow 50.7$
$\frac{\sin^2 \theta_{13}}{10^{-2}}$	$2.241^{+0.066}_{-0.065}$	$2.046 \rightarrow 2.440$	$2.237^{+0.066}_{-0.065}$	$2.044 \rightarrow 2.435$	$2.14^{+0.09}_{-0.07}$	$1.90 \rightarrow 2.39$	$2.160^{+0.083}_{-0.069}$	$1.96 \rightarrow 2.41$
$\theta_{13}/^{\circ}$	$8.61^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.99$	$8.60^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.98$	$8.41^{+0.18}_{-0.14}$	$7.9 \rightarrow 8.9$	$8.45_{-0.14}^{+0.16}$	$8.0 \rightarrow 8.9$
$\delta_{\rm CP}/^{\circ}$	$222_{-28}^{+38}$	$141 \rightarrow 370$	$221_{-28}^{+39}$	$144 \rightarrow 357$	$238_{-33}^{+41}$	$149 \rightarrow 358$	$218^{+38}_{-27}$	$157 \rightarrow 349$
$\frac{\Delta m_{21}^2}{10^{-5}  {\rm gV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.34_{-0.14}^{+0.17}$	$6.92 \rightarrow 7.91$	$7.55_{-0.16}^{+0.20}$	$7.05 \rightarrow 8.24$
$\frac{\Delta m_{32}^2}{10^{-3} \text{ eV}^2}$	$2.449^{+0.032}_{-0.030}$	$2.358 \rightarrow 2.544$	$2.454_{-0.031}^{+0.029}$	$2.362 \rightarrow 2.544$	$2.419^{+0.035}_{-0.032}$	$2.319 \rightarrow 2.521$	$2.424 \pm 0.03$	$2.334 \rightarrow 2.524$
IO	$\Delta \chi^2 = 6.2$		$\Delta \chi^2 = 10.4$		$\Delta \chi^2 = 9.5$		$\Delta \chi^2 = 11.7$	
$\frac{\sin^2 \theta_{12}}{10^{-1}}$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.03^{+0.14}_{-0.13}$	$2.64 \rightarrow 3.45$	$3.20^{+0.20}_{-0.16}$	$2.73 \rightarrow 3.79$
$\theta_{12}/^{\circ}$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.75}$	$31.62 \rightarrow 36.27$	$33.40^{+0.87}_{-0.81}$	$30.92 \rightarrow 35.97$	$34.5^{+1.2}_{-1.0}$	$31.5 \rightarrow 38.0$
$\frac{\sin^2 \theta_{23}}{10^{-1}}$	$5.63^{+0.19}_{-0.26}$	$4.30 \rightarrow 6.12$	$5.65^{+0.17}_{-0.22}$	$4.36 \rightarrow 6.10$	$5.57^{+0.17}_{-0.24}$	$4.44 \rightarrow 6.03$	$5.51^{+0.18}_{-0.30}$	$4.53 \rightarrow 5.98$
$\theta_{23}/^{\circ}$	$48.6^{+1.1}_{-1.5}$	$41.0 \rightarrow 51.5$	$48.8^{+1.0}_{-1.2}$	$41.4 \rightarrow 51.3$	$48.2^{+1.0}_{-1.4}$	$41.8 \rightarrow 50.9$	$47.9^{+1.0}_{-1.7}$	$42.3 \rightarrow 50.7$
$\frac{\sin^2 \theta_{13}}{10^{-2}}$	$2.261^{+0.067}_{-0.064}$	$2.066 \rightarrow 2.461$	$2.259^{+0.065}_{-0.065}$	$2.064 \rightarrow 2.457$	$2.18^{+0.08}_{-0.07}$	$1.95 \rightarrow 2.43$	$2.220^{+0.074}_{-0.076}$	$1.99 \rightarrow 2.44$
$\theta_{13}/^{\circ}$	$8.65^{+0.13}_{-0.12}$	$8.26 \rightarrow 9.02$	$8.64^{+0.12}_{-0.13}$	$8.26 \rightarrow 9.02$	$8.49^{+0.15}_{-0.14}$	$8.0 \rightarrow 9.0$	$8.53^{+0.14}_{-0.15}$	$8.1 \rightarrow 9.0$
$\delta_{\rm CP}/^{\circ}$	$285^{+24}_{-26}$	$205 \rightarrow 354$	$282^{+23}_{-25}$	$205 \rightarrow 348$	$247^{+26}_{-27}$	$193 \rightarrow 346$	$281^{+23}_{-27}$	$202 \rightarrow 349$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.34_{-0.14}^{+0.17}$	$6.92 \rightarrow 7.91$	$7.55_{-0.16}^{+0.20}$	$7.05 \rightarrow 8.24$
$\frac{\Delta m_{32}^2}{10^{-3} \text{ eV}^2}$	$-2.509^{+0.032}_{-0.032}$	$-2.603 \rightarrow -2.416$	$-2.510^{+0.030}_{-0.031}$	$-2.601 \rightarrow -2.419$	$-2.478^{+0.035}_{-0.033}$	$-2.577 \rightarrow -2.375$	$-2.50\pm^{+0.04}_{-0.03}$	$-2.59 \rightarrow -2.39$

### Other observations

- Chau-Keung parametrization  $\operatorname{Im}(r) \propto \frac{s_{13}s_{23}}{s_{12}}$
- Larger mixing angles in lepton sector due to

$$\frac{m_2^2}{m_3^2} \approx 3.1 \times 10^{-2} \gg \frac{m_s^2}{m_b^2} \approx 9.0 \times 10^{-4}$$

- How about  $\tau^-e^+$ - $\tau^+e^-$  or  $t\bar{u}$ - $\bar{t}u$  mixing? Same intermediate channels, so same constraints. Additional solutions?
- Two possibilities: first, small  $\lambda_i \lambda_j$ , so that constraints  $|V_{cs}^* V_{us}|^2 = \lambda^2 \approx 5 \times 10^{-2}$ met automatically, happening to quark sector  $|V_{ts}^* V_{us}|^2 = A^2 \lambda^6 \approx 9 \times 10^{-5}$ .
- Second, two solutions of v with opposite signs: one for  $\mu^-e^+-\mu^+e^-$  another for  $\tau^-e^+-\tau^+e^-$
- Check data  $U_{\tau 1}^* U_{e1} / (U_{\tau 2}^* U_{e2})$   $r = U_{\mu 1}^* U_{e1} / (U_{\mu 2}^* U_{e2})$   $-(1.231_{-0.186}^{+0.078}) + (0.204_{-0.138}^{+0.085})i$   $r = -(0.738_{-0.048}^{+0.050}) - (0.179_{-0.125}^{+0.136})i$  de Salas et al, 2018  $-(1.139_{-0.207}^{+0.139}) + (0.266_{-0.124}^{+0.050})i$   $r = -(0.801_{-0.097}^{+0.219}) - (0.265_{-0.145}^{+0.090})i$  Capozzi et al, 2018

### Constraint on $\theta_{23}$

• Ratios in CK parametrization

$$\begin{aligned} & \text{roughly equal} \\ \text{atios in CK parametrization} \\ & \frac{U_{\mu 1}^{*}U_{e1}}{U_{\mu 2}^{*}U_{e2}} = -\frac{c_{12}}{s_{12}} \frac{c_{12}s_{12}(c_{23}^{2} - s_{13}^{2}s_{23}^{2}) + c_{23}s_{13}s_{23}c_{\delta}(c_{12}^{2} - s_{12}^{2}) - c_{23}s_{13}s_{23}s_{\delta}i}{(c_{12}c_{23} - s_{12}s_{13}s_{23})^{2} + 2c_{12}c_{23}s_{12}s_{13}s_{23}(1 - c_{\delta})} \\ & \frac{U_{\tau 1}^{*}U_{e1}}{U_{\tau 2}^{*}U_{e2}} = -\frac{c_{12}}{s_{12}} \frac{c_{12}s_{12}(s_{23}^{2} - c_{23}^{2}s_{13}^{2}) - c_{23}s_{13}s_{23}c_{\delta}(c_{12}^{2} - s_{12}^{2}) + c_{23}s_{13}s_{23}s_{\delta}i}{(c_{12}s_{23} - c_{23}s_{13}s_{23}c_{\delta}(c_{12}^{2} - s_{12}^{2}) + c_{23}s_{13}s_{23}s_{\delta}i} \\ & \text{equal} \end{aligned}$$

- The two ratios differ only by sign of Im
- Relation among mixing angles  $(c_{12}^2 - s_{12}^2 s_{13}^2)(c_{23}^2 - s_{23}^2) - 4c_{12}c_{23}s_{12}s_{13}s_{23}c_\delta \approx 0,$  $(c_{12}^2 + s_{12}^2 s_{13}^2)(c_{23}^2 - s_{23}^2) \approx 0 \quad \longleftarrow \quad c_{12}s_{12}(1 + s_{13}^2)(c_{23}^2 - s_{23}^2) + 2(c_{12}^2 - s_{12}^2)c_{23}s_{13}s_{23}c_\delta \approx 0$
- Indicate  $c_{23} \approx s_{23}$ , i.e.,  $\theta_{23} \approx 45^{\circ}$  .... maximal lepton mixing!

#### Summary and conjecture

- Dispersion relations for observables impose stringent constraint
- Fermion masses and mixing angles in both quark and lepton sectors constrained by SM dynamics itself
- Different mixing patterns due to different fermion mass ratios

$$\frac{m_2^2}{m_3^2} \approx 3.1 \times 10^{-2} \gg \frac{m_s^2}{m_b^2} \approx 9.0 \times 10^{-4}$$

- Normal hierarchy favored by dispersive constraint
- Maximal lepton mixing demanded by solutions for mixing between generations 1, 2 and generations 1, 3
- It is likely that SM has only three fundamental (gauge) parameters
- Scalar sector, coupling various generations, is not free