

Search for Dark Matter in pp Collisions with CMS

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Colloquium, Department of Physics
National Taiwan University

Outline

- Dark matter searches at colliders
 - Overview
 - Introduction to LHC and CMS
 - Experimental techniques
 - Interpretation of results
- Conclusion and outlook

How Do You “See” an Object?

Reflection



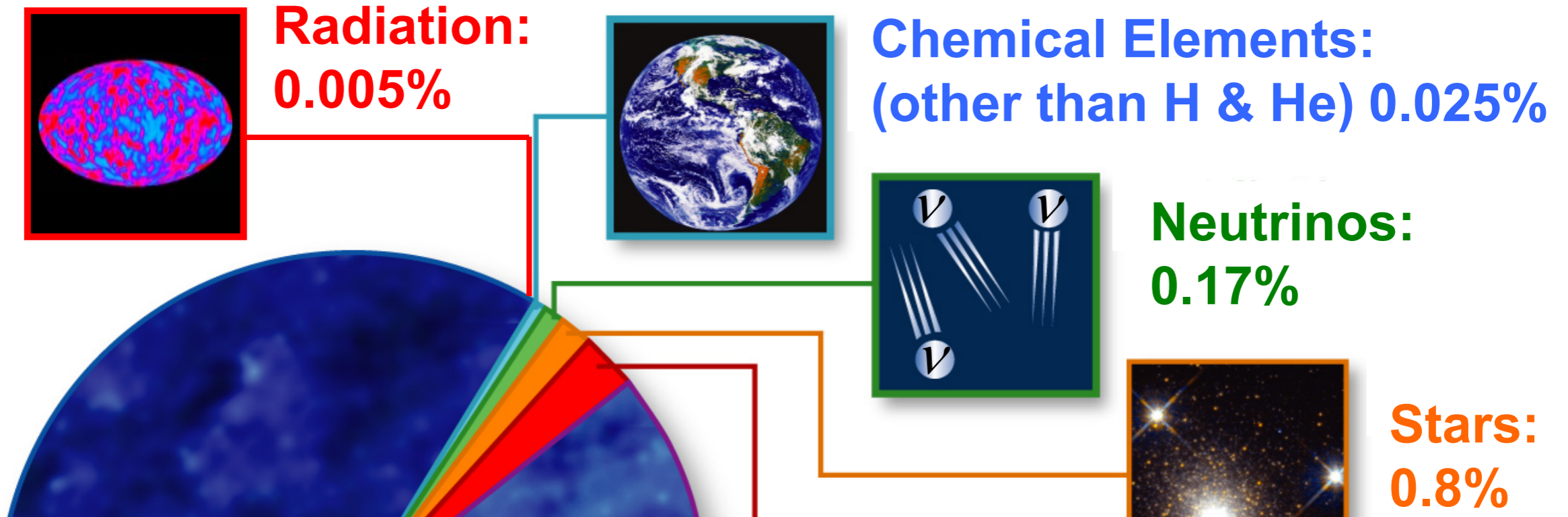
Thermal Radiation

Visible for $T=3800\sim 7600$ K

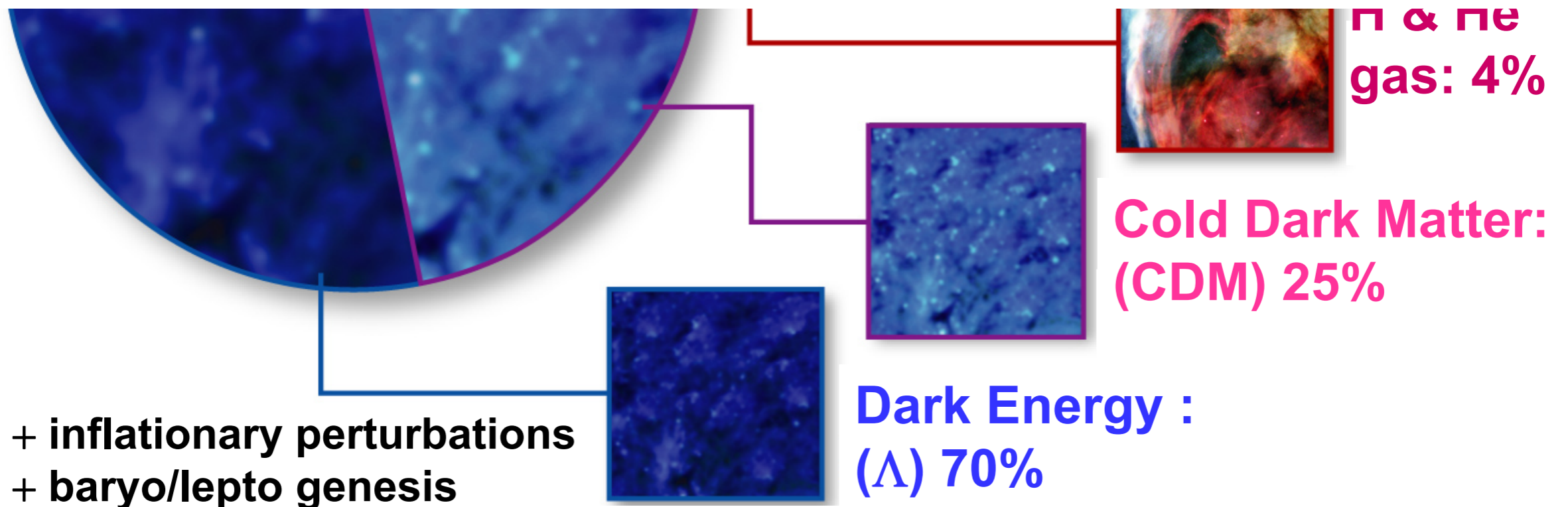


In our galaxy, besides visible stars, is there something else?



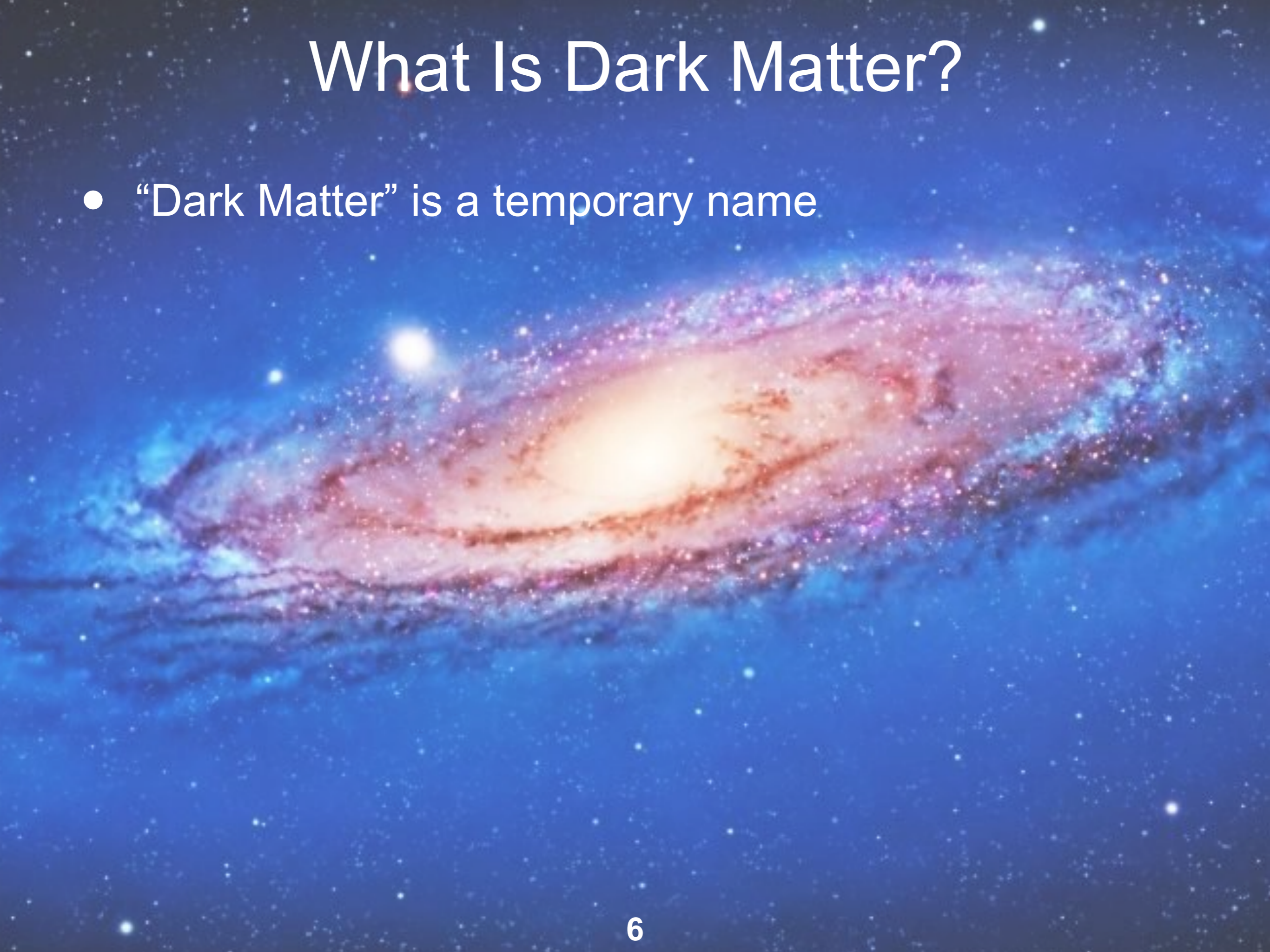


If I had been present at creation, I would have suggested a simpler scheme.
- Alfonse the Wise

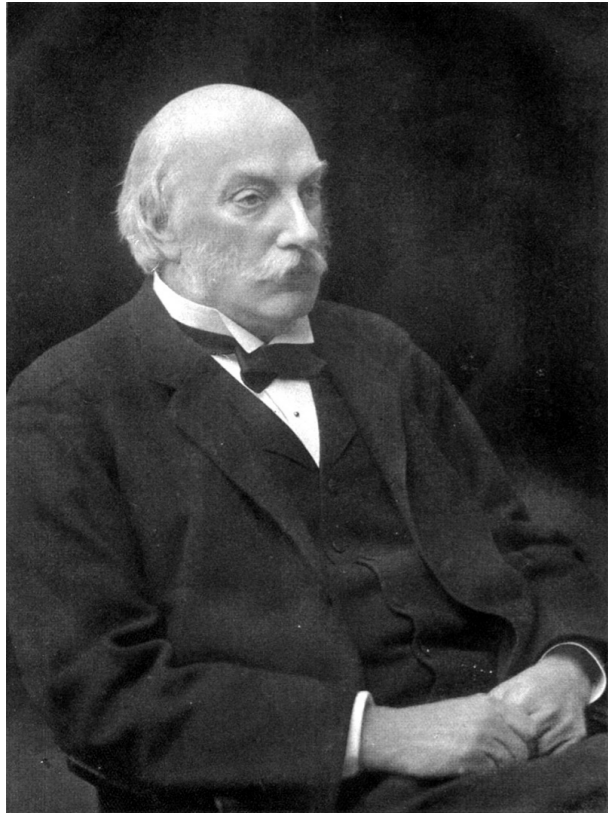


What Is Dark Matter?

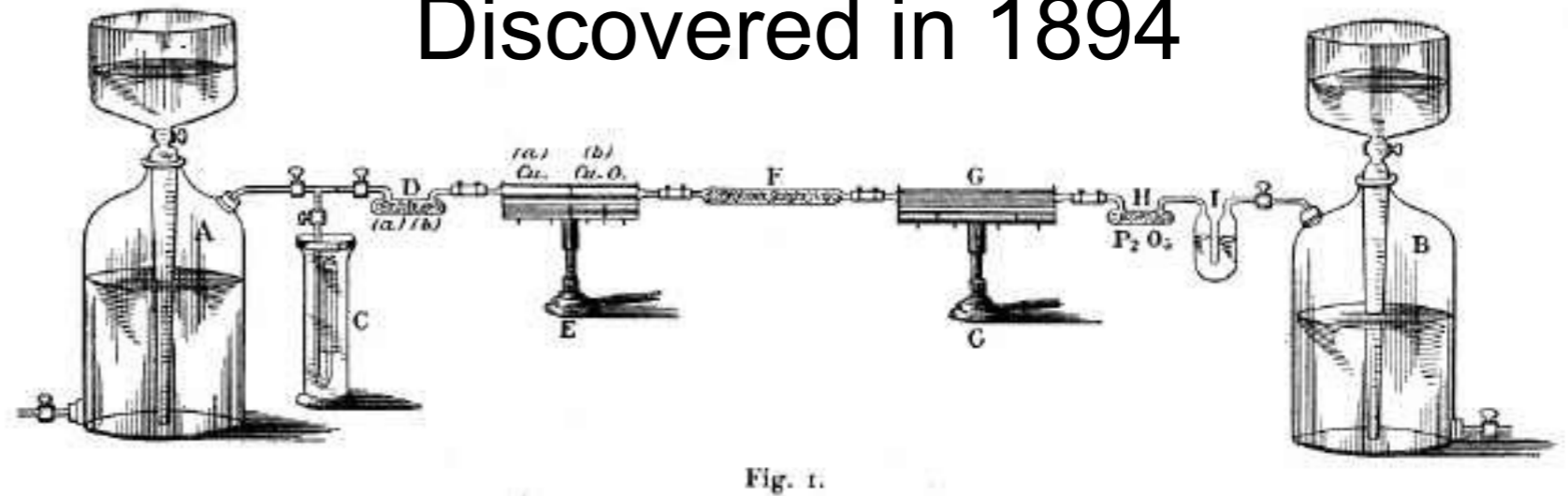
- “Dark Matter” is a temporary name



“Dark Matter” in Chemistry: Argon

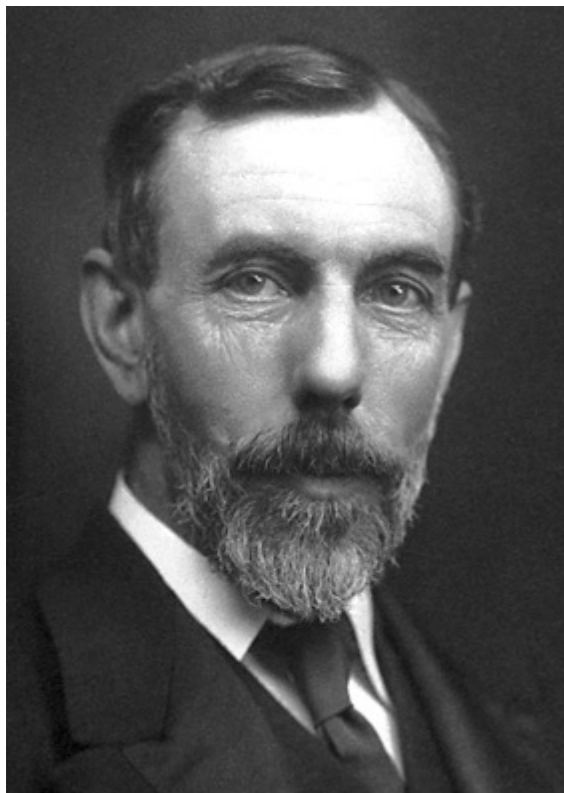


Lord Rayleigh



Nitrogen extracted from air is heavier than that extracted from the chemical reaction by 0.5%

👉 New Unknown Gas: Argon



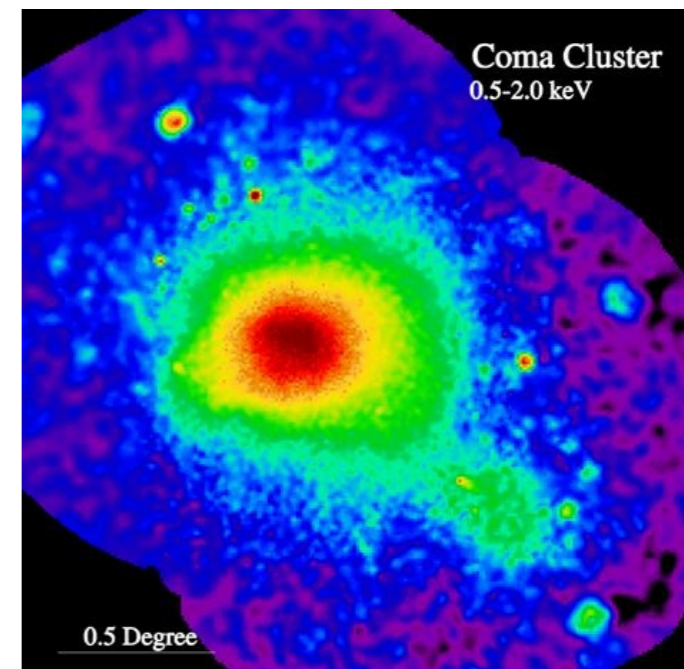
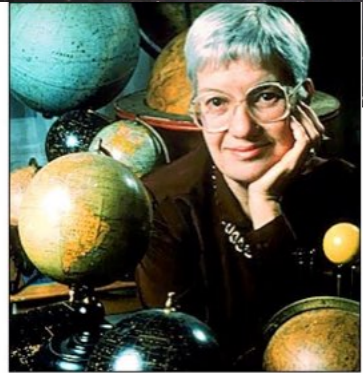
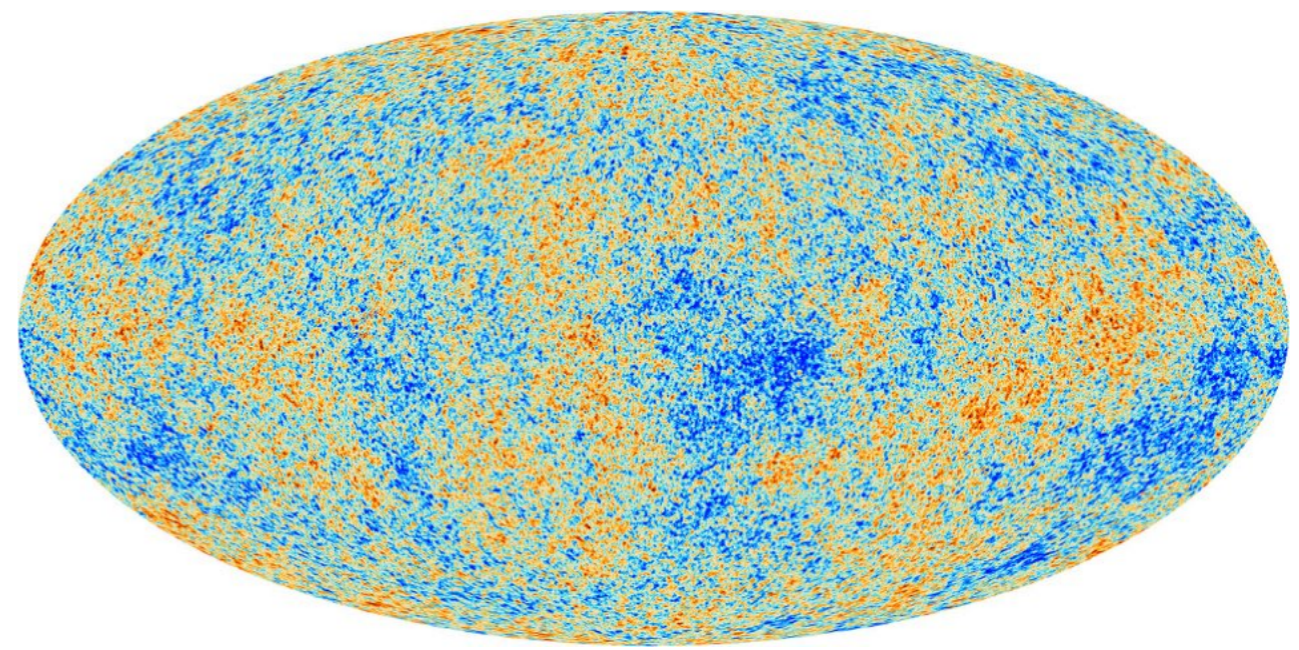
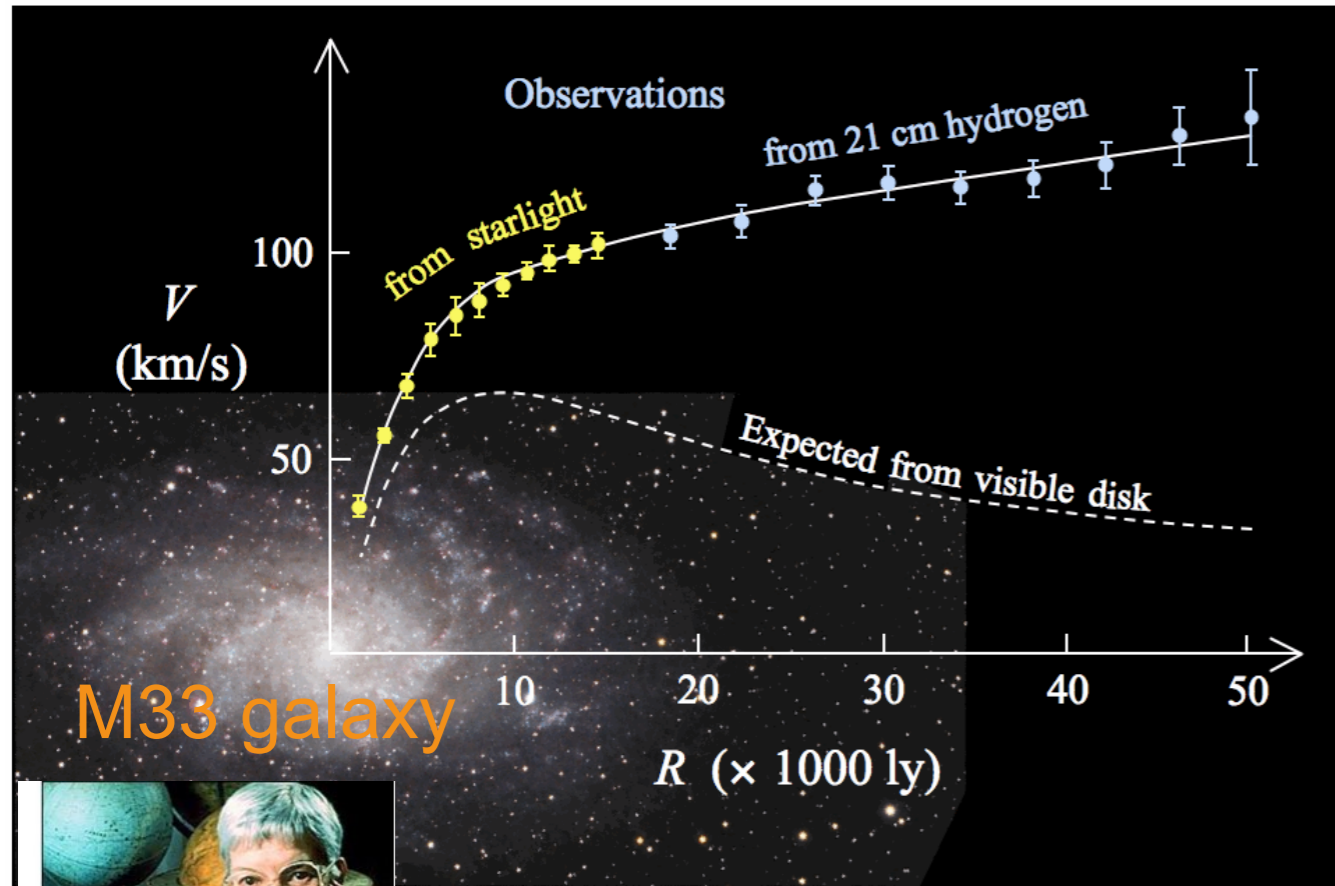
Sir William Ramsay

What Is Dark Matter?

- “Dark Matter” is a temporary name

- Influenced by gravitational interaction and no other standard model (SM) interactions
- Interact weakly with normal matter → may need a new type of interaction

Why Dark Matter?



Reminder of Gravitation Law (Outside the Earth)

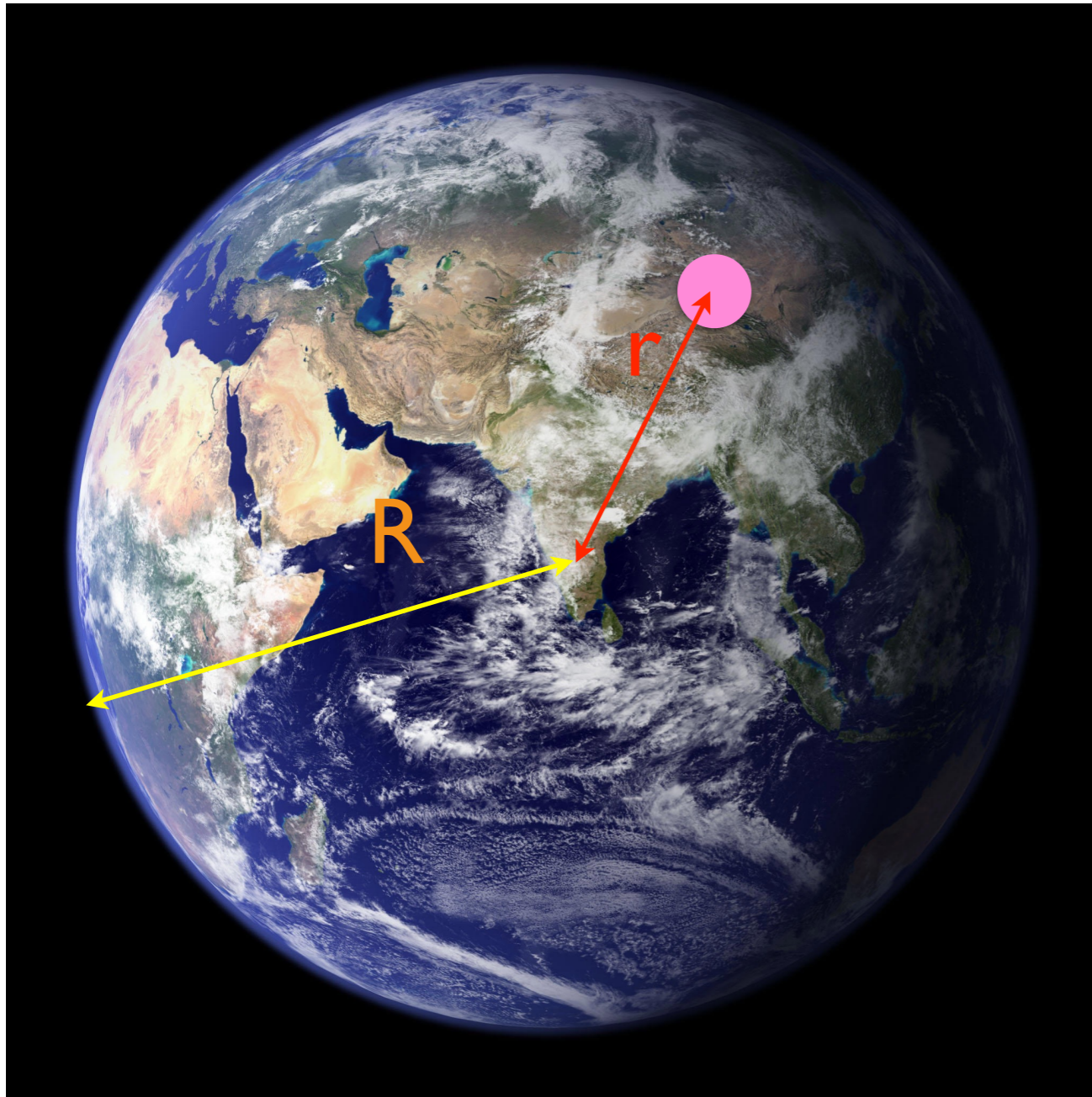


$$F(r > R) \propto \frac{M_{\text{total}}}{r^2}$$

$$m \frac{v_{\text{particle}}^2}{r} \propto \frac{M_{\text{total}}}{r^2}$$

$$\Rightarrow v_{\text{particle}} \propto \sqrt{\frac{M_{\text{total}}}{r}}$$

Reminder of Gravitation Law: Inside the Earth



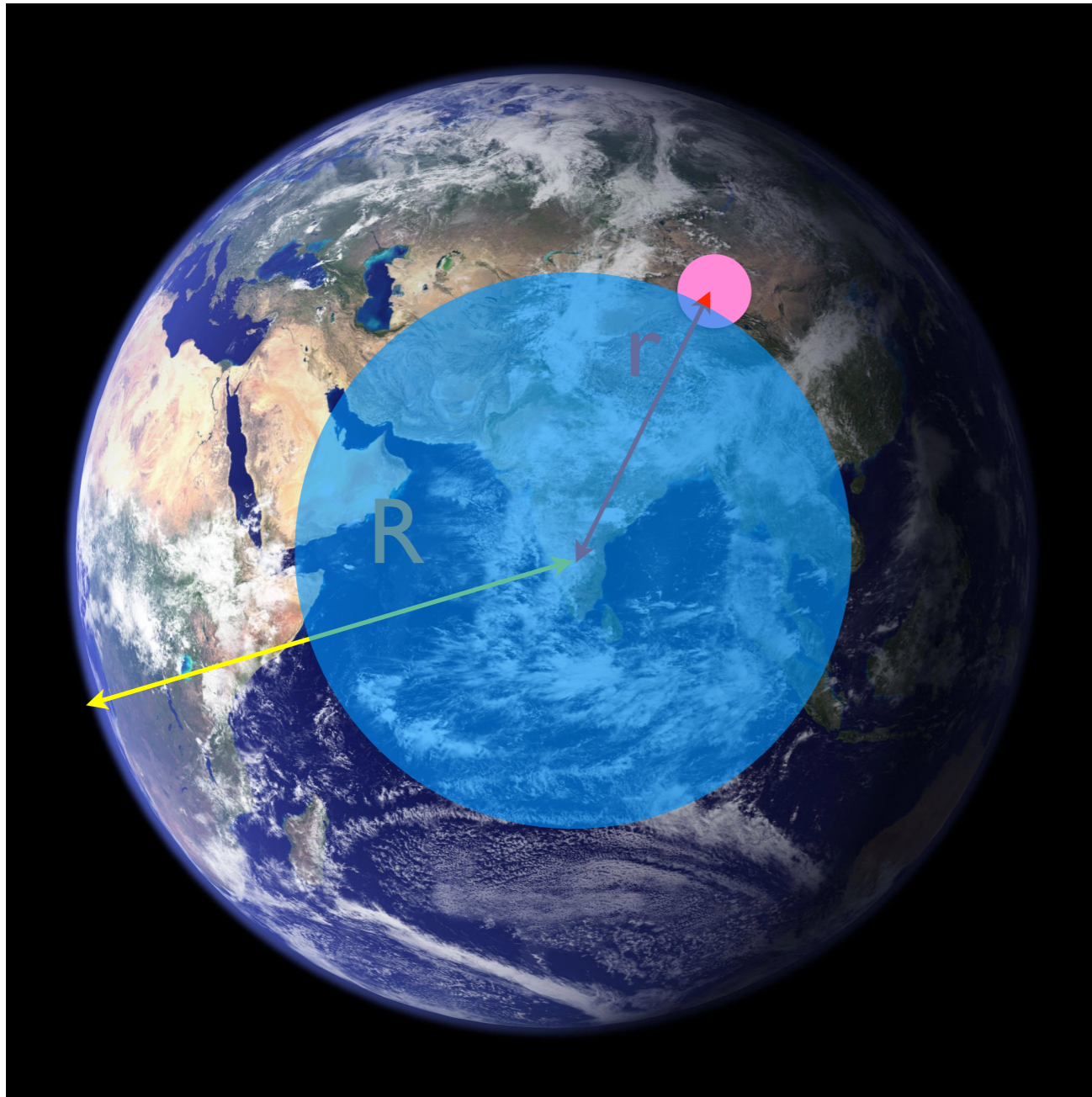
$$F(r < R) \propto \frac{M(r)}{r^2}$$

$$= \frac{\rho \left(\frac{4\pi}{3} r^3 \right)}{r^2} = r$$

$$m \frac{v_{\text{particle}}^2}{r} \propto r$$

$$\Rightarrow v_{\text{particle}} \propto r$$

Reminder of Gravitation Law: Inside the Earth



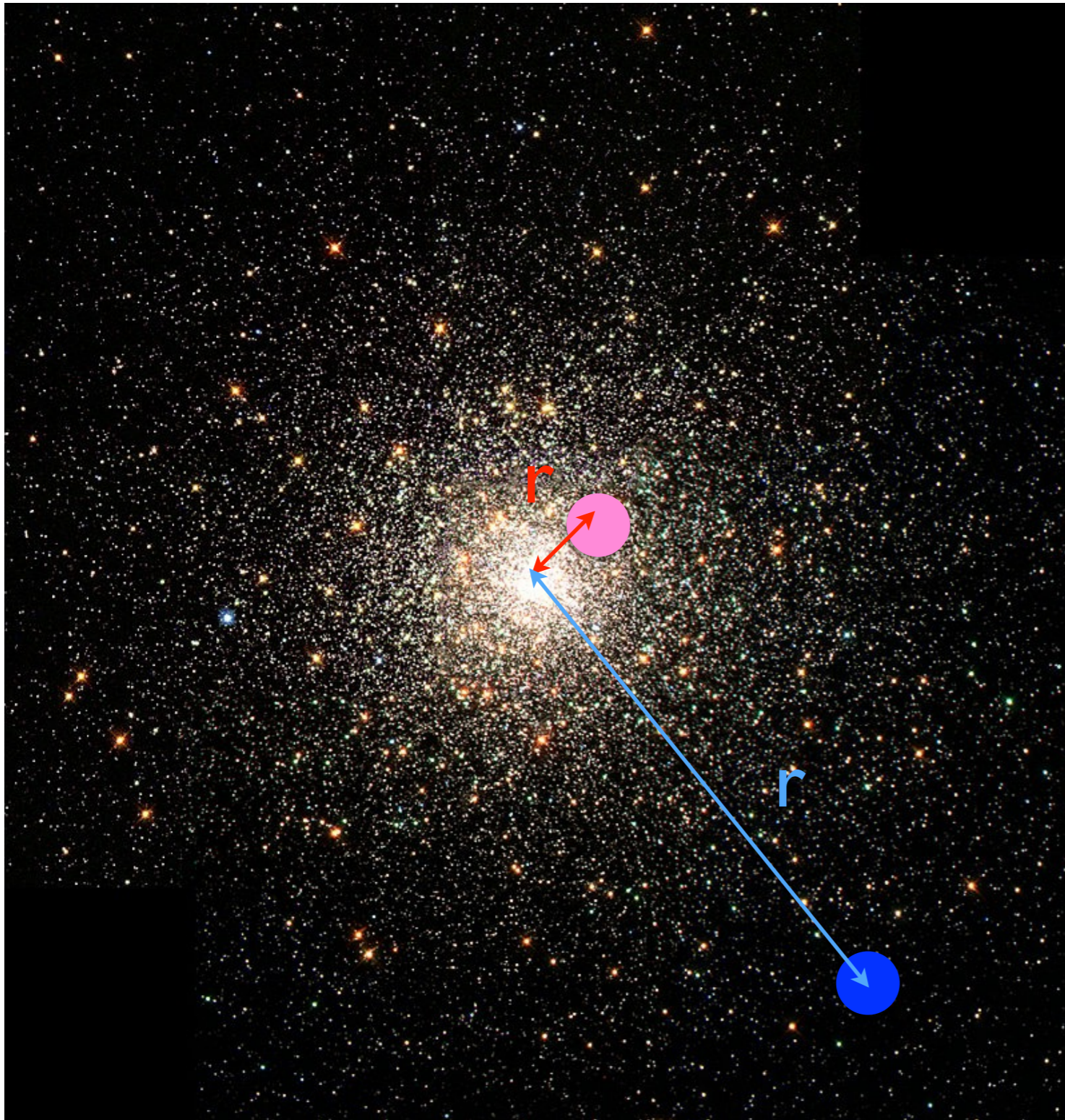
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Extended to the Stars in a Galaxy

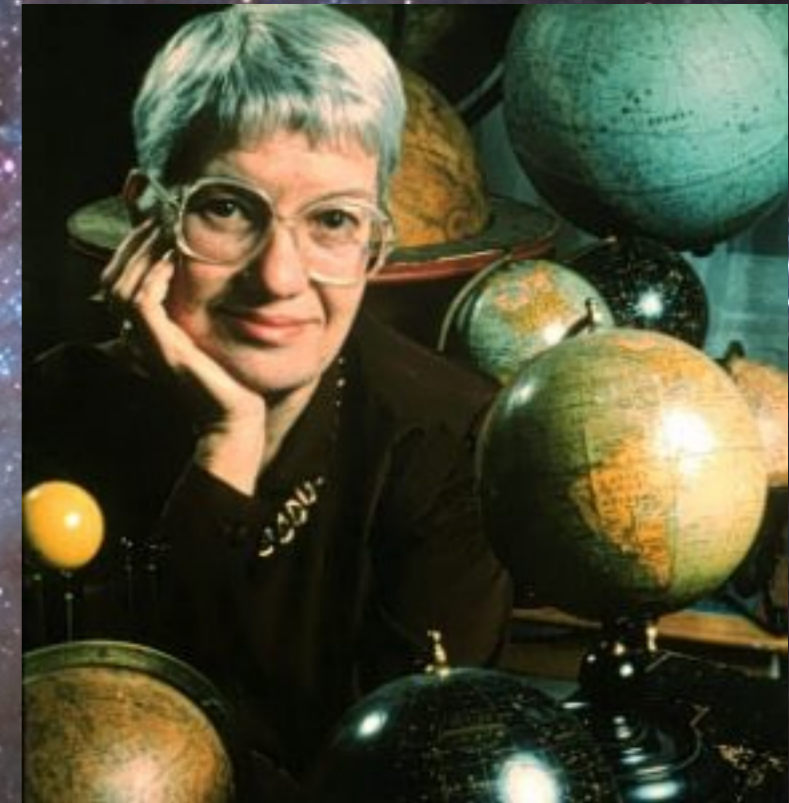
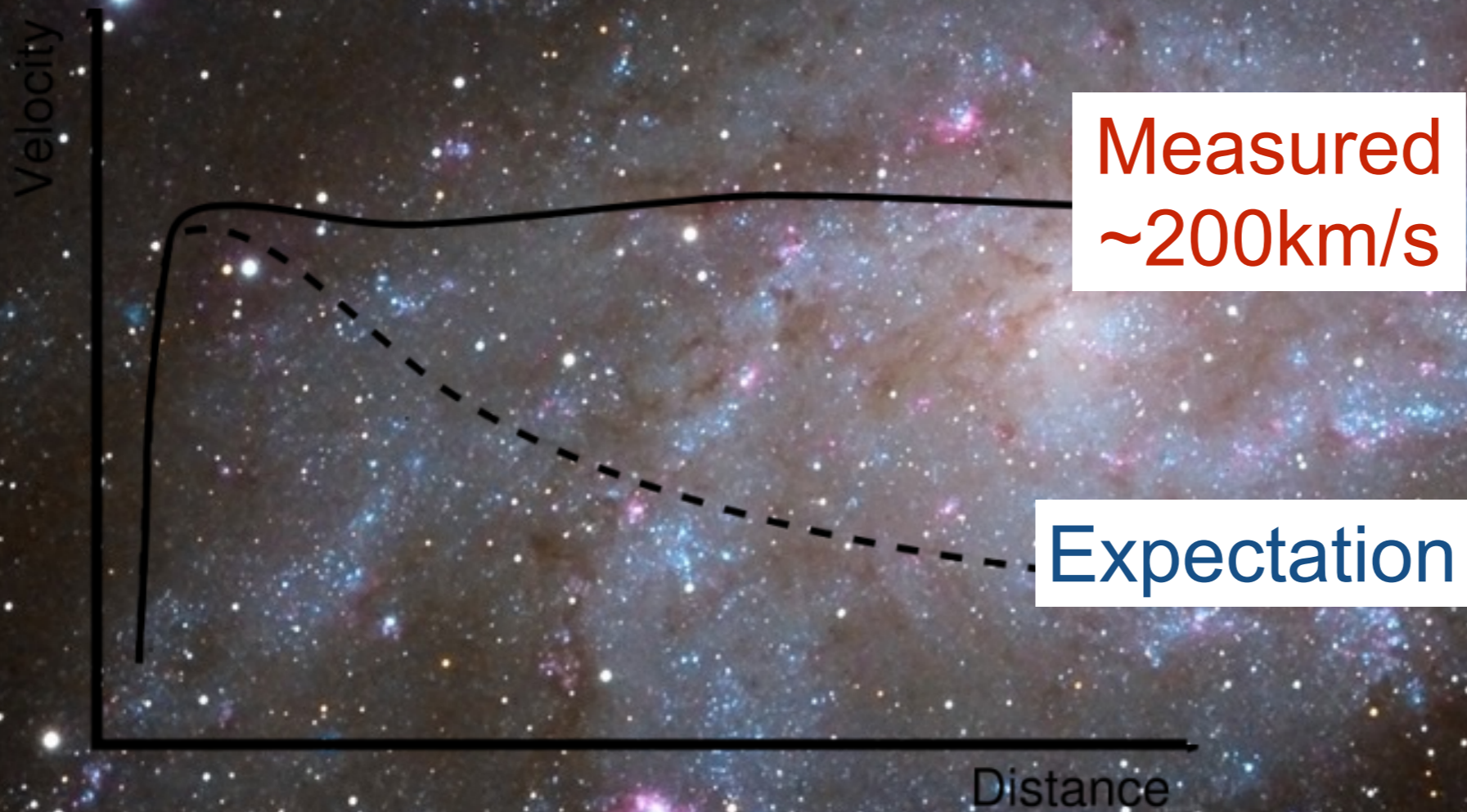


$$v \propto \sqrt{\frac{M(r)}{r}}$$

$$v(r < R) \propto r$$

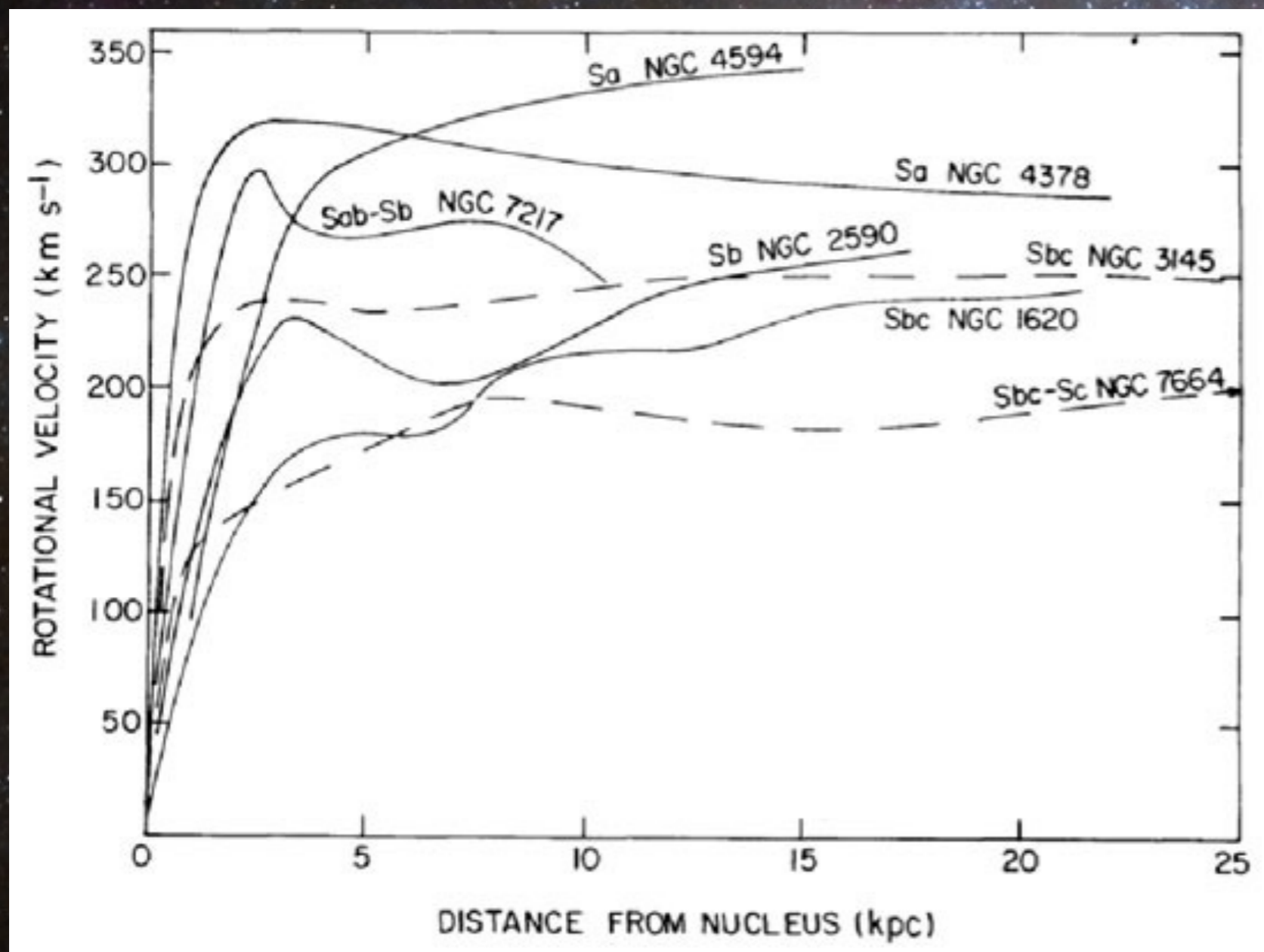
$$v(r > R) \propto \frac{1}{\sqrt{r}}$$

Rotational Curves



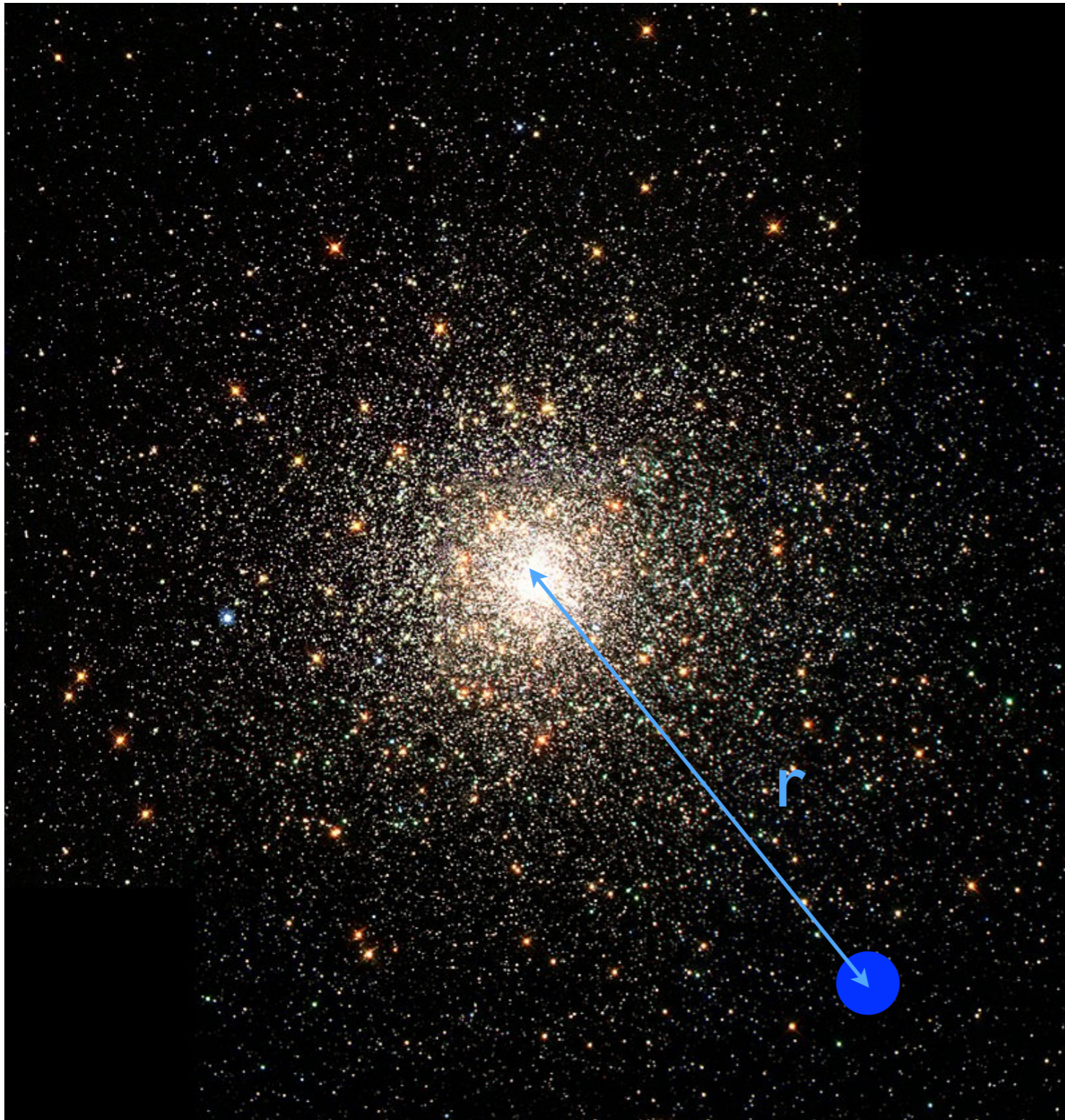
Vera Rubin
1928-2016

Rotational Curves



Vera Rubin
1928-2016

Extended to the Stars in a Galaxy

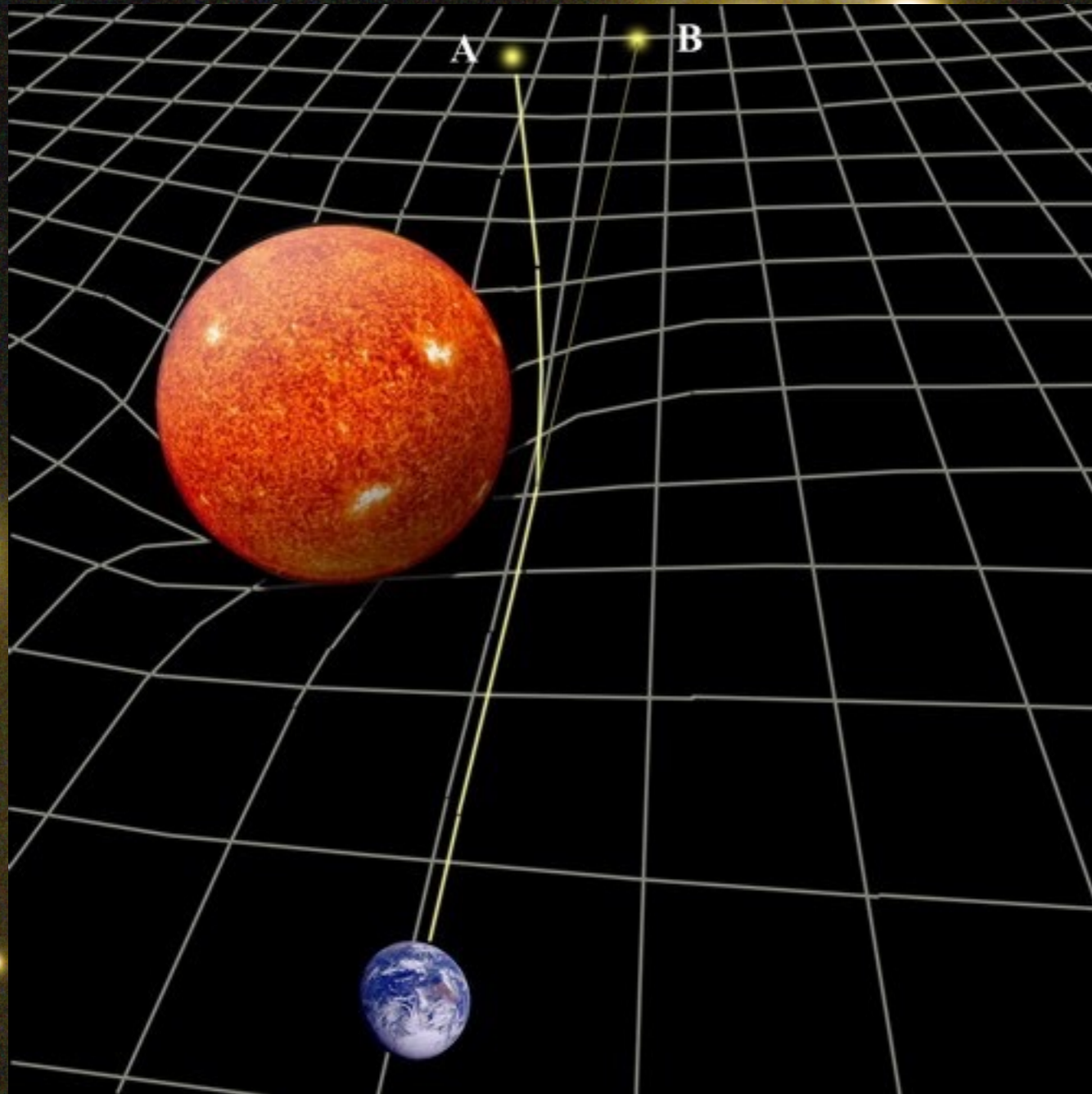


~~$$v_{\text{particle}}(r > R) \propto \sqrt{\frac{M_{\text{total}}}{r}}$$~~

$$v_{\text{particle}}(r > R) \propto \sqrt{\frac{M(r)}{r}}$$

$$\Rightarrow M(r) \propto r$$

Gravitation Lensing



The size of the Einstein ring is related to the mass of the lensing

$$\theta \propto \sqrt{M_{\text{lense}}}$$

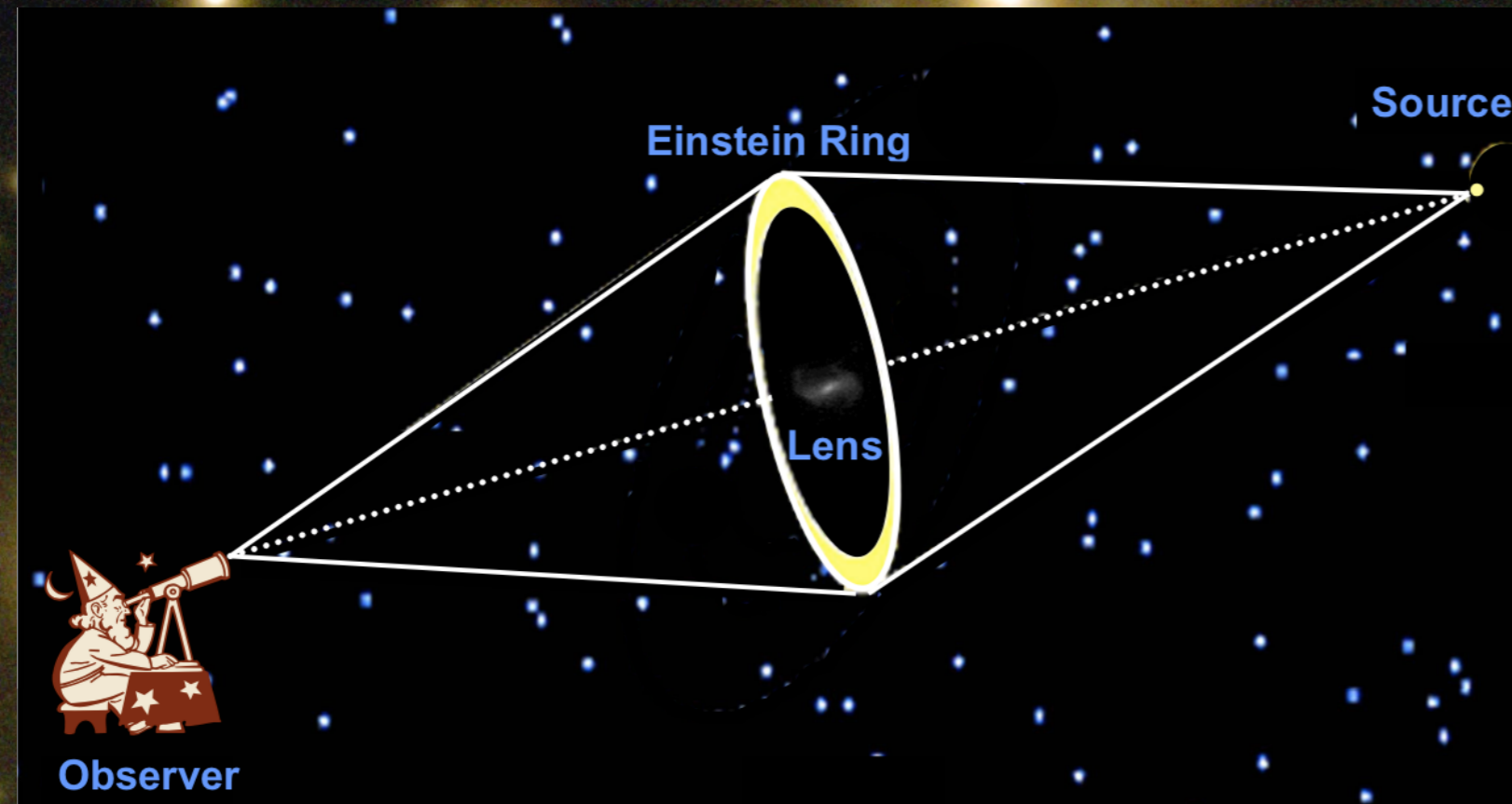
Abell 2218
Cluster

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Abell 2218
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Abell 2218
Cluster

Bullet Clusters



1E0657-558
2006 observed

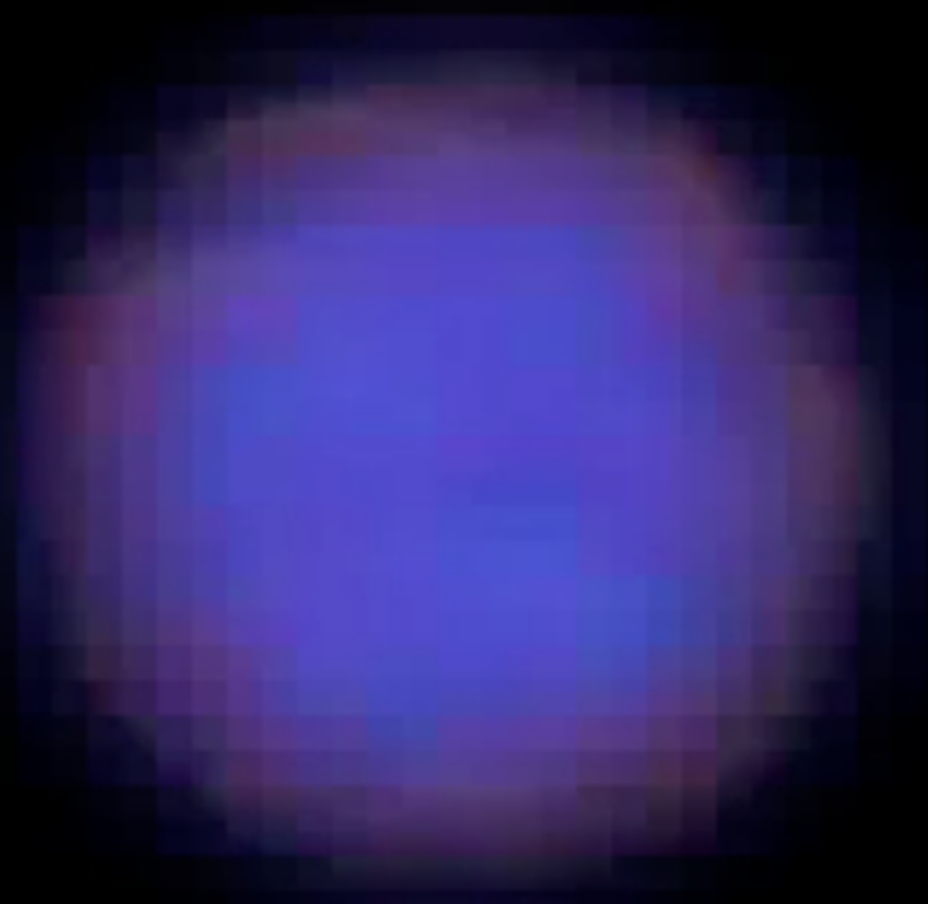
Bullet Clusters

Normal Matter
(X-ray image)

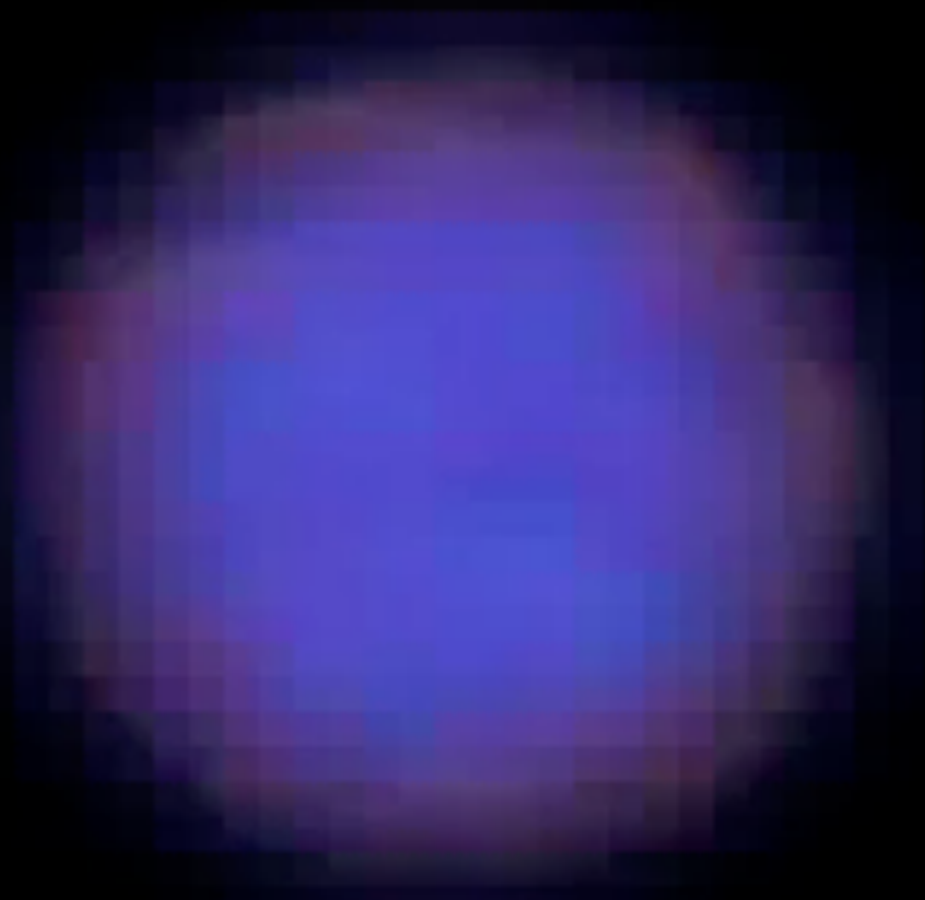
Dark Matter

1E0657-558
2006 observed

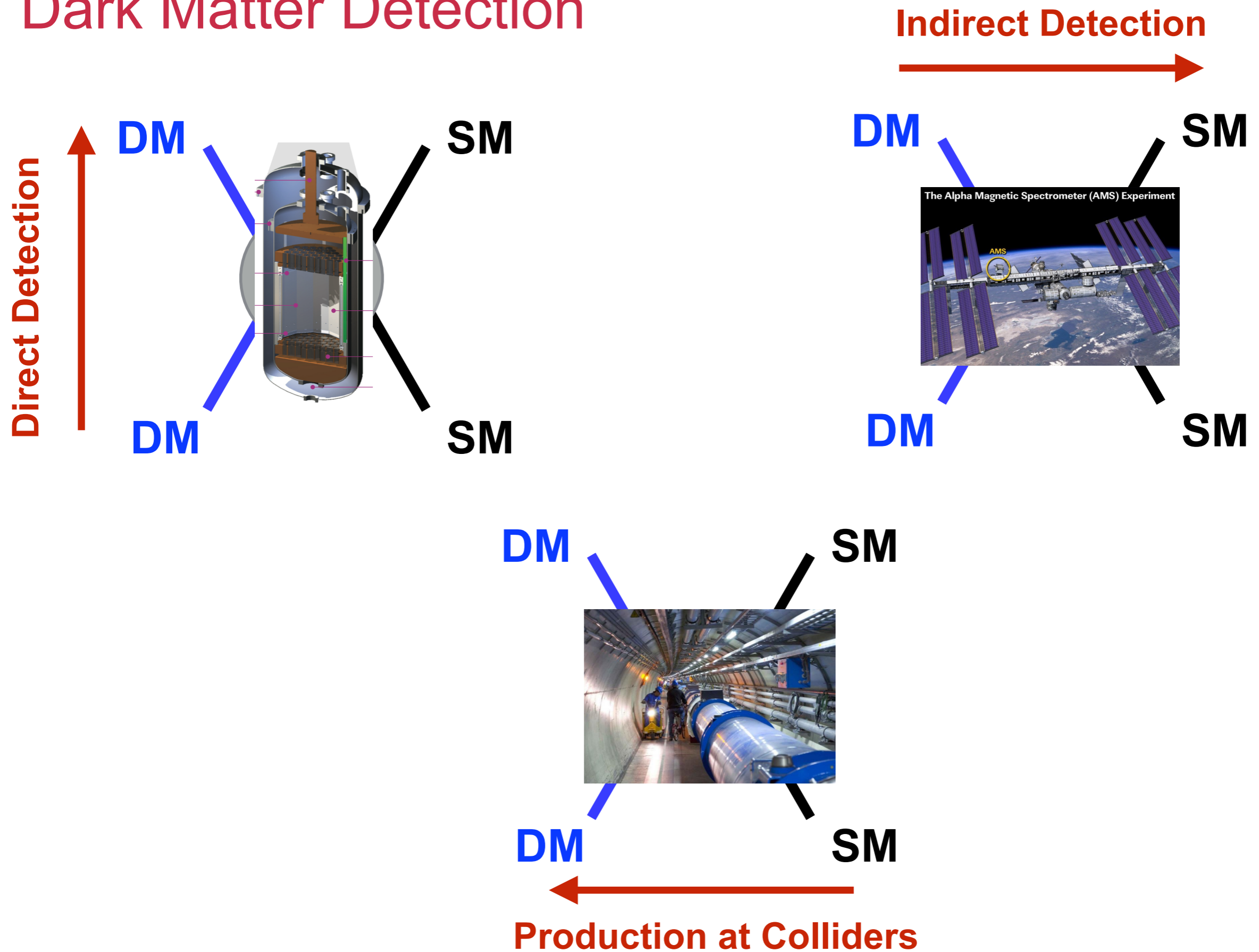
Bullet Clusters



Bullet Clusters



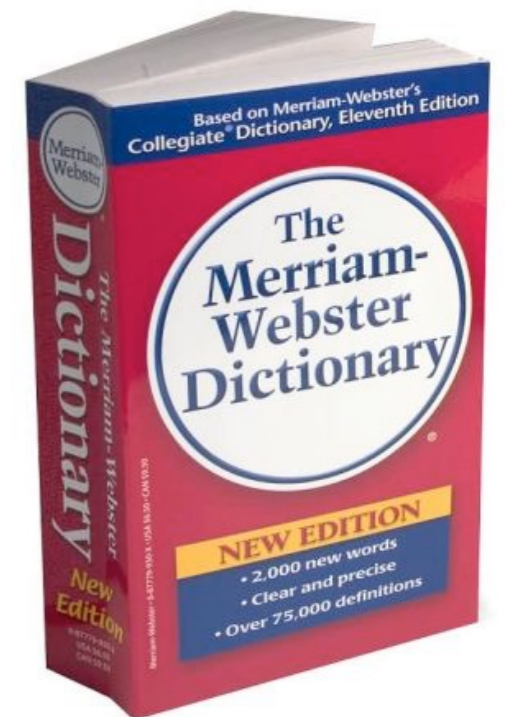
Dark Matter Detection



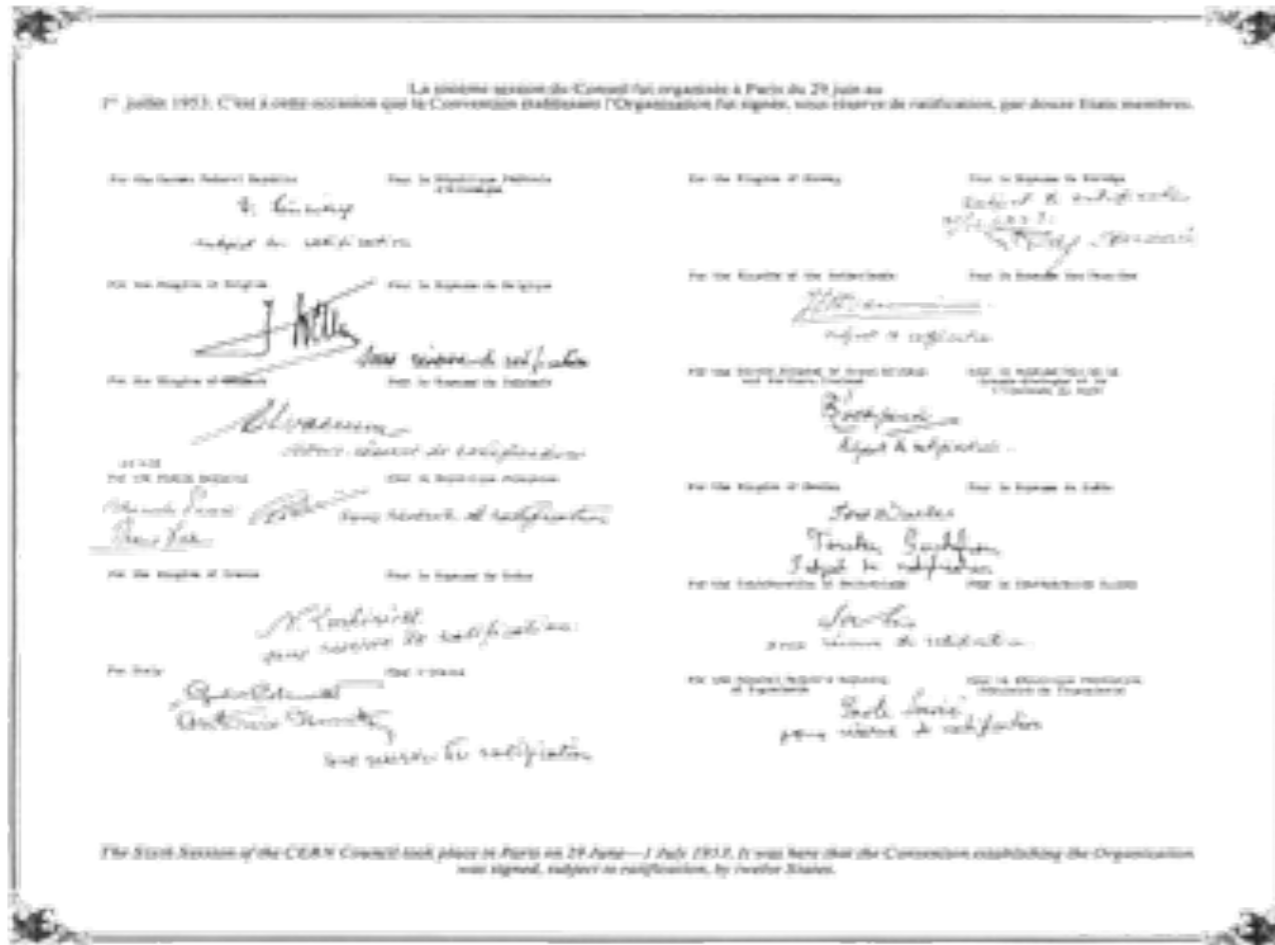
Introduction to LHC and CMS

- **CERN**

- Conseil Européen pour la Recherche Nucléaire
- European Council for Nuclear Research
- Location of LHC and the experiments



CERN



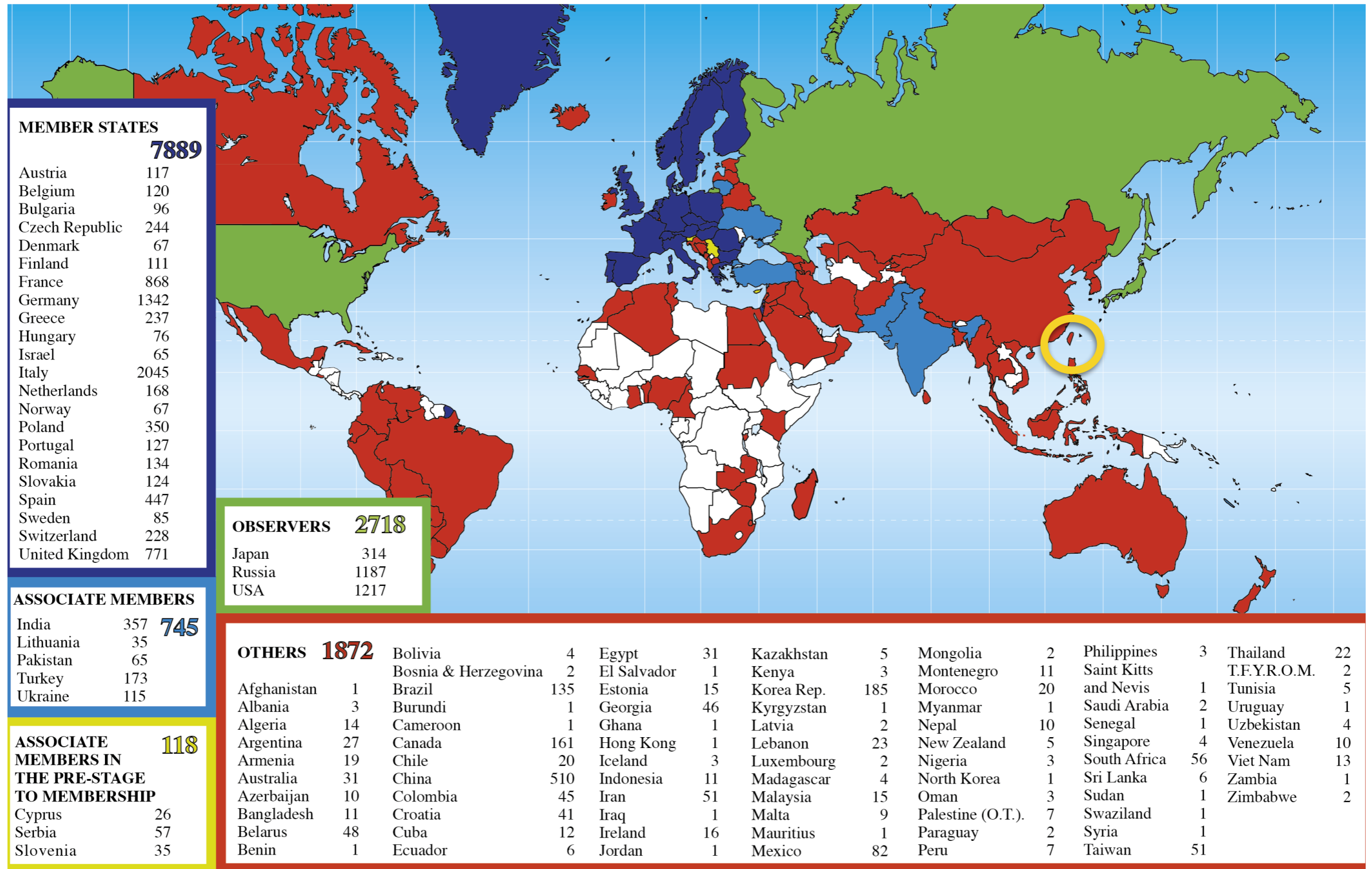
- Established by 12 European countries on 1954/09/29
- Origin of WWW
 - ▶ Tim Berners-Lee in 1989
- Director
 - ▶ Fabiola Gianotti

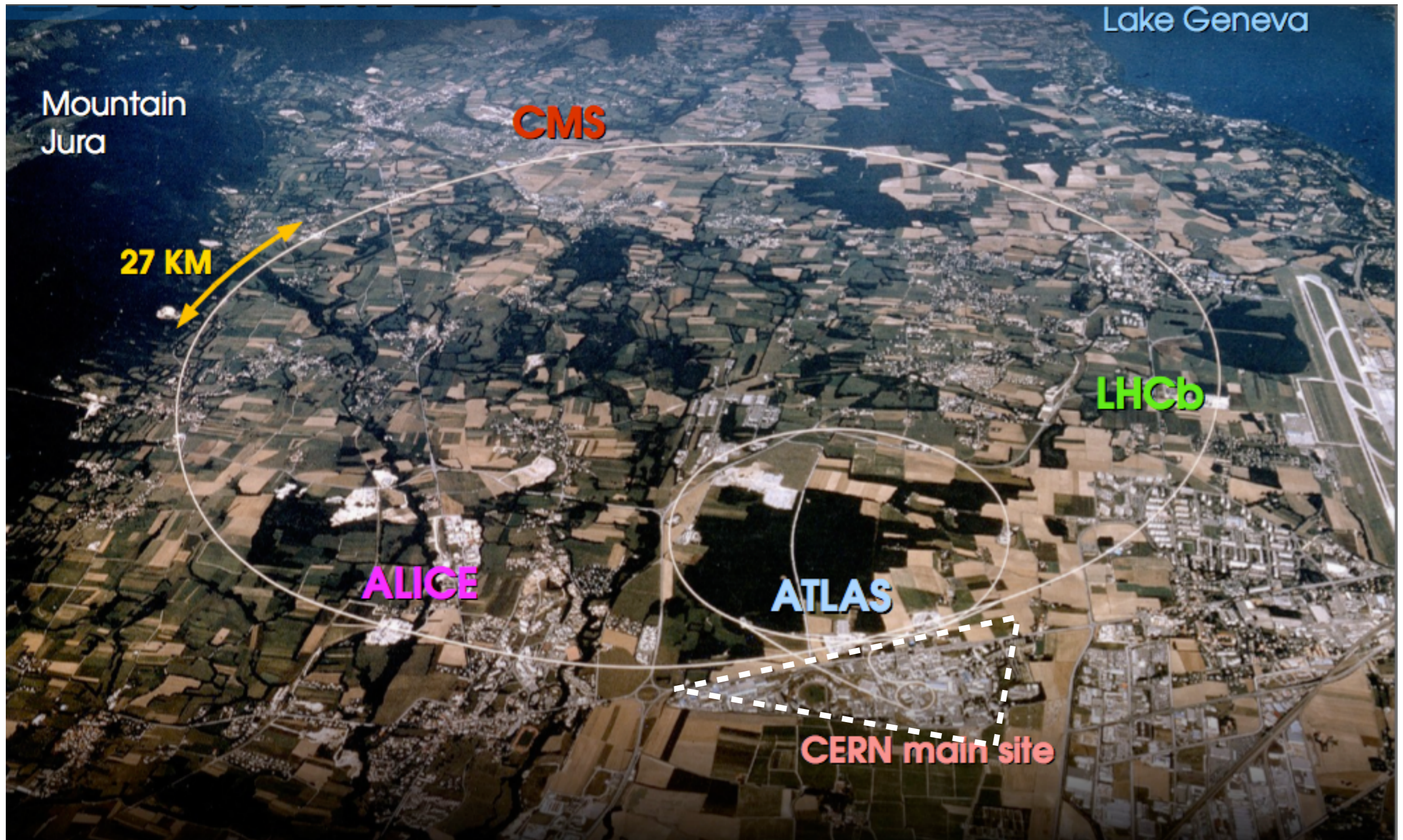
- 23 member states
- Yearly budget $\sim 10^9$ CHF (= 3.2×10^{10} TWD)
 - ▶ Germany, UK, France, Italy
 - ▶ LHC cost $\sim 4.3 \times 10^9$ CHF



Users Around the World

Distribution of All CERN Users by Nationality on 24 January 2018





LHC Birdview

Moun
Jura

eneva

ATLAS

CERN main site

LHC Birdview

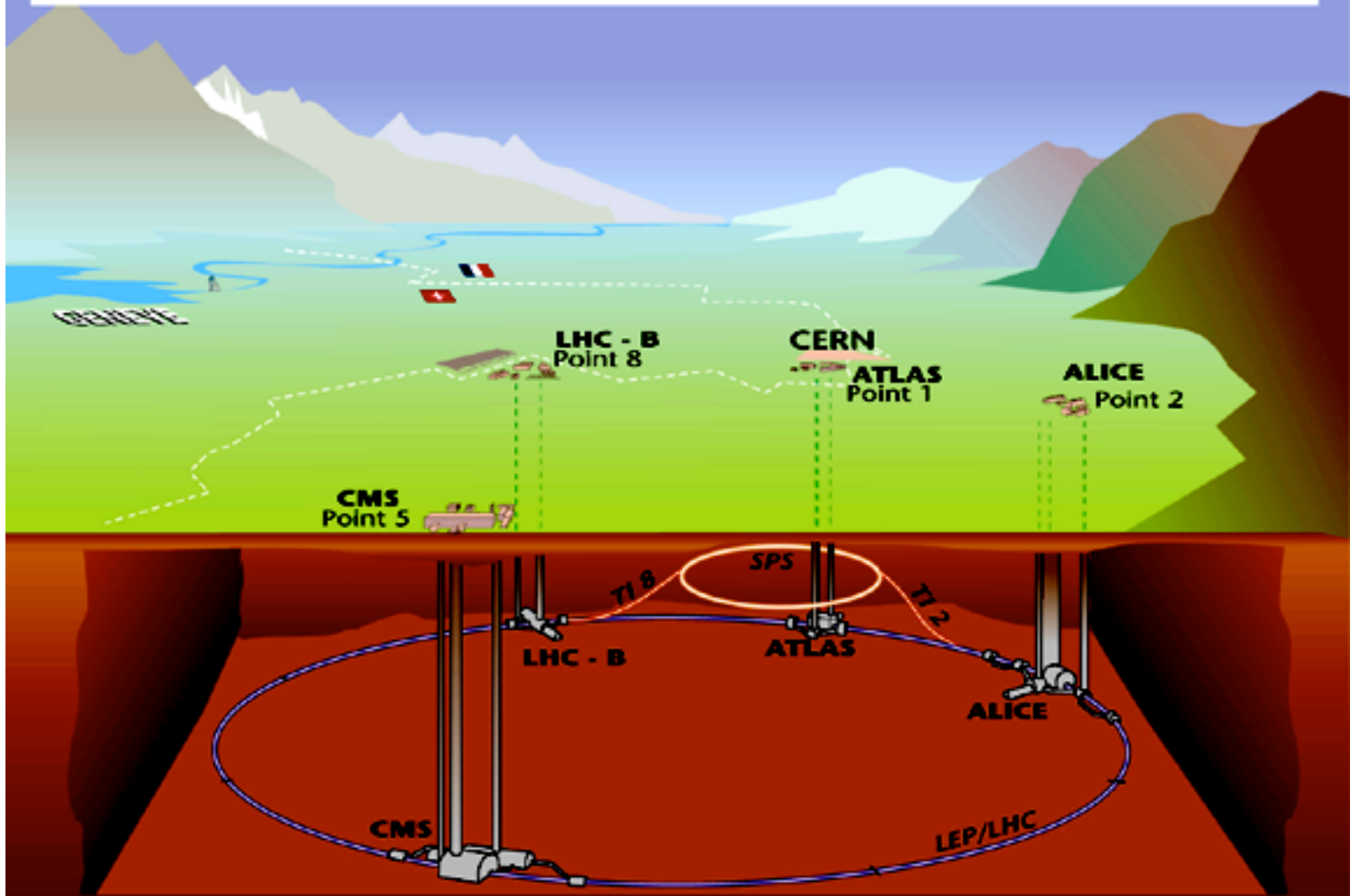


LHC Birdview



LHC Birdview

Overall view of the LHC experiments.



Tunnel circumference 26.7 km, tunnel diameter 3.8 m
Depth : ~ 50-175 m - tunnel is inclined by ~ 1.4%

Photothèque - E540 - V10/09/97

LHC Tunnel



Magnetic dipole field: 8.3 Tesla
Beam-pipe pressure: 10^{-13} atm

1232 superconducting dipoles
Operating temperature: 1.9 K

LHC Tunnel



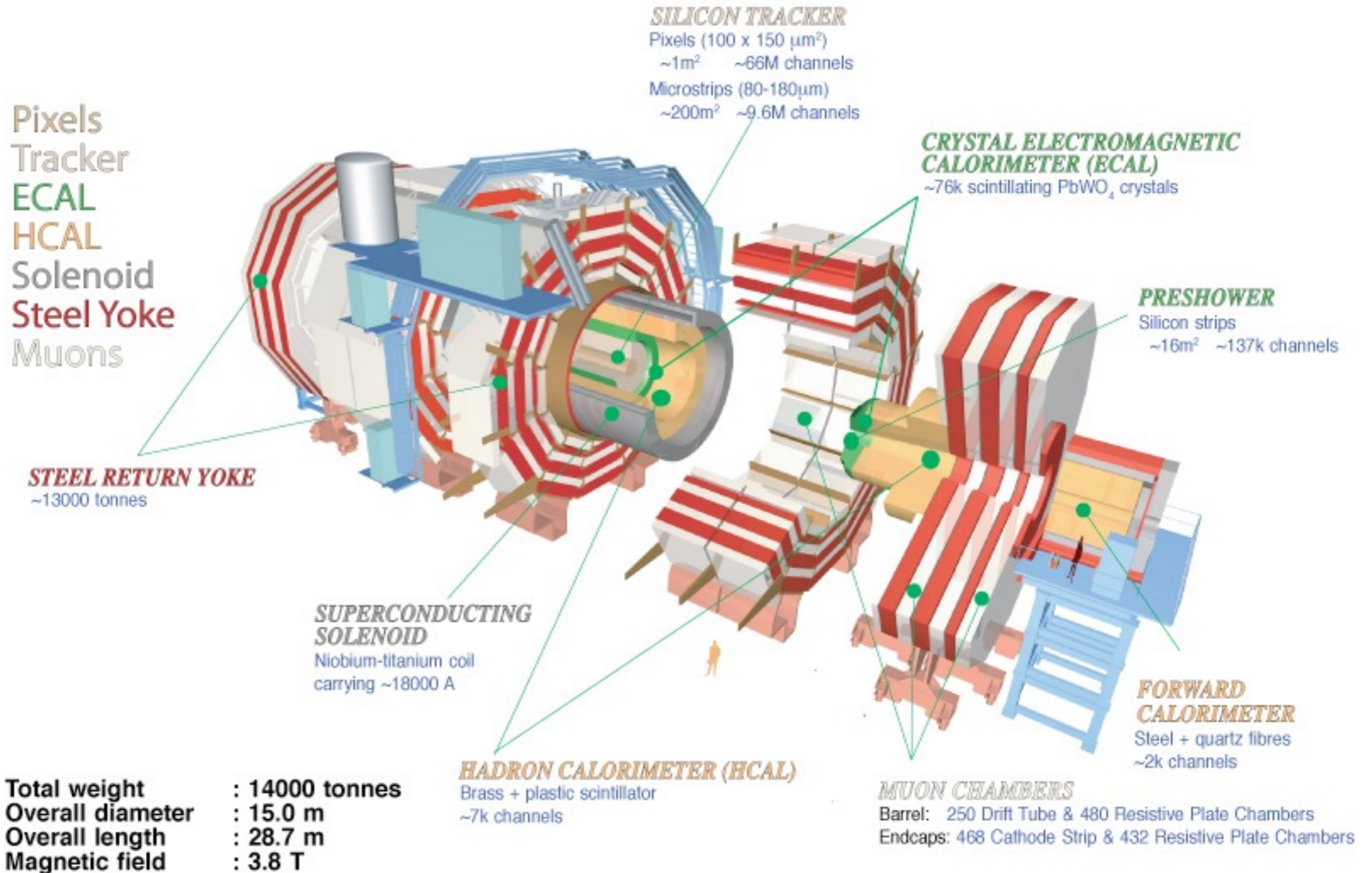
LHC Tunnel

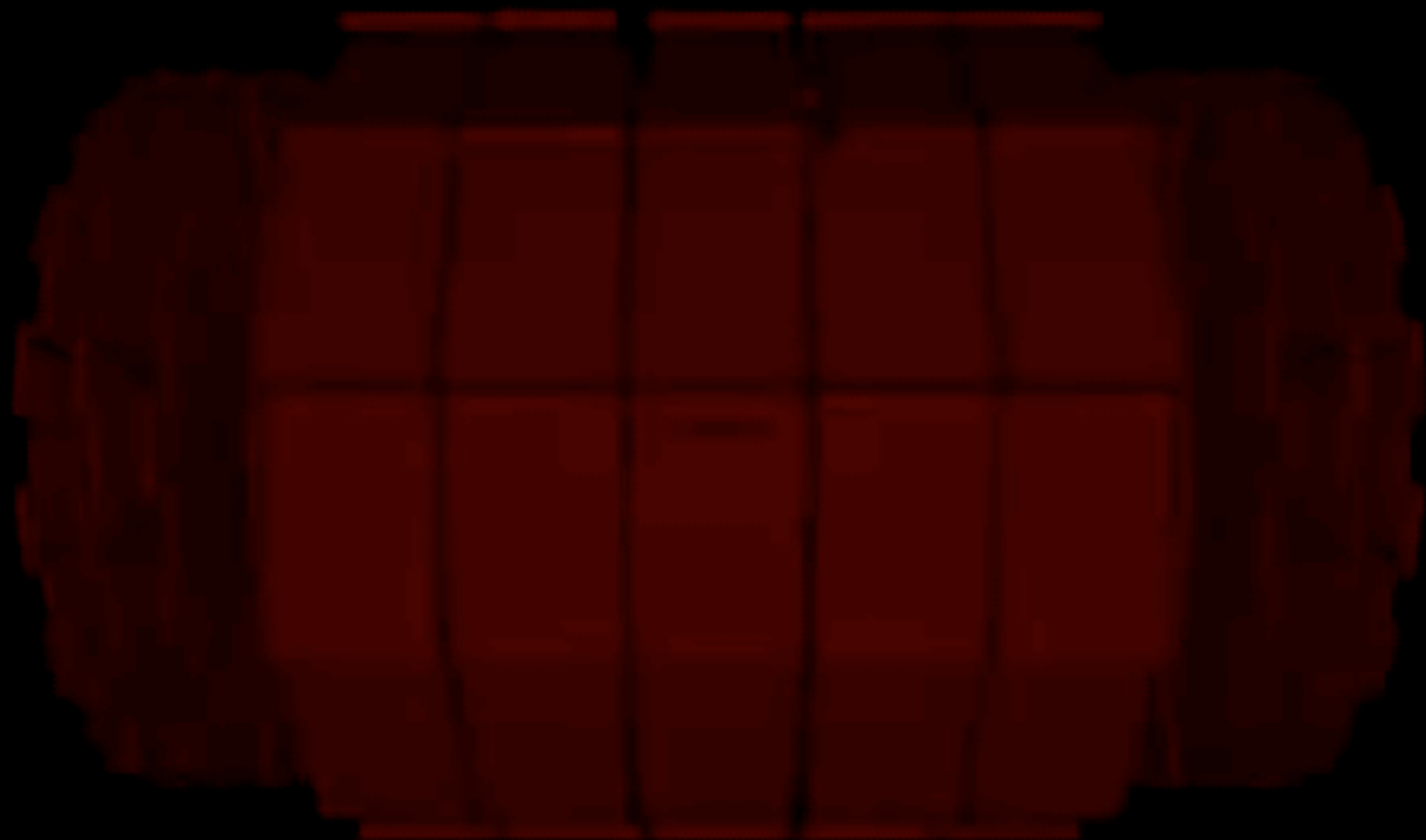


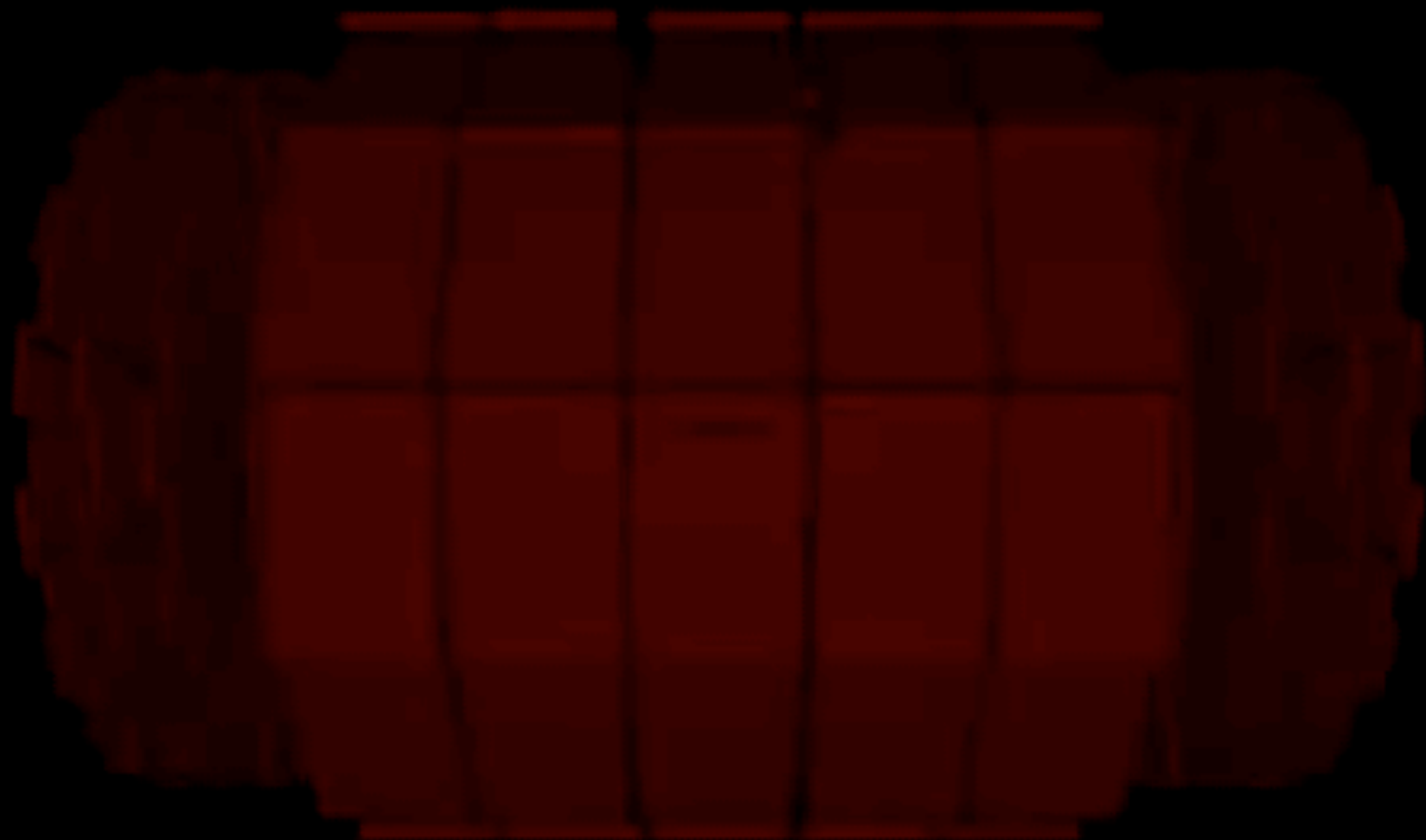
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1232 superconducting dipoles
Operating temperature: 1.9 K

CMS Detector Sketch







Muon Chamber

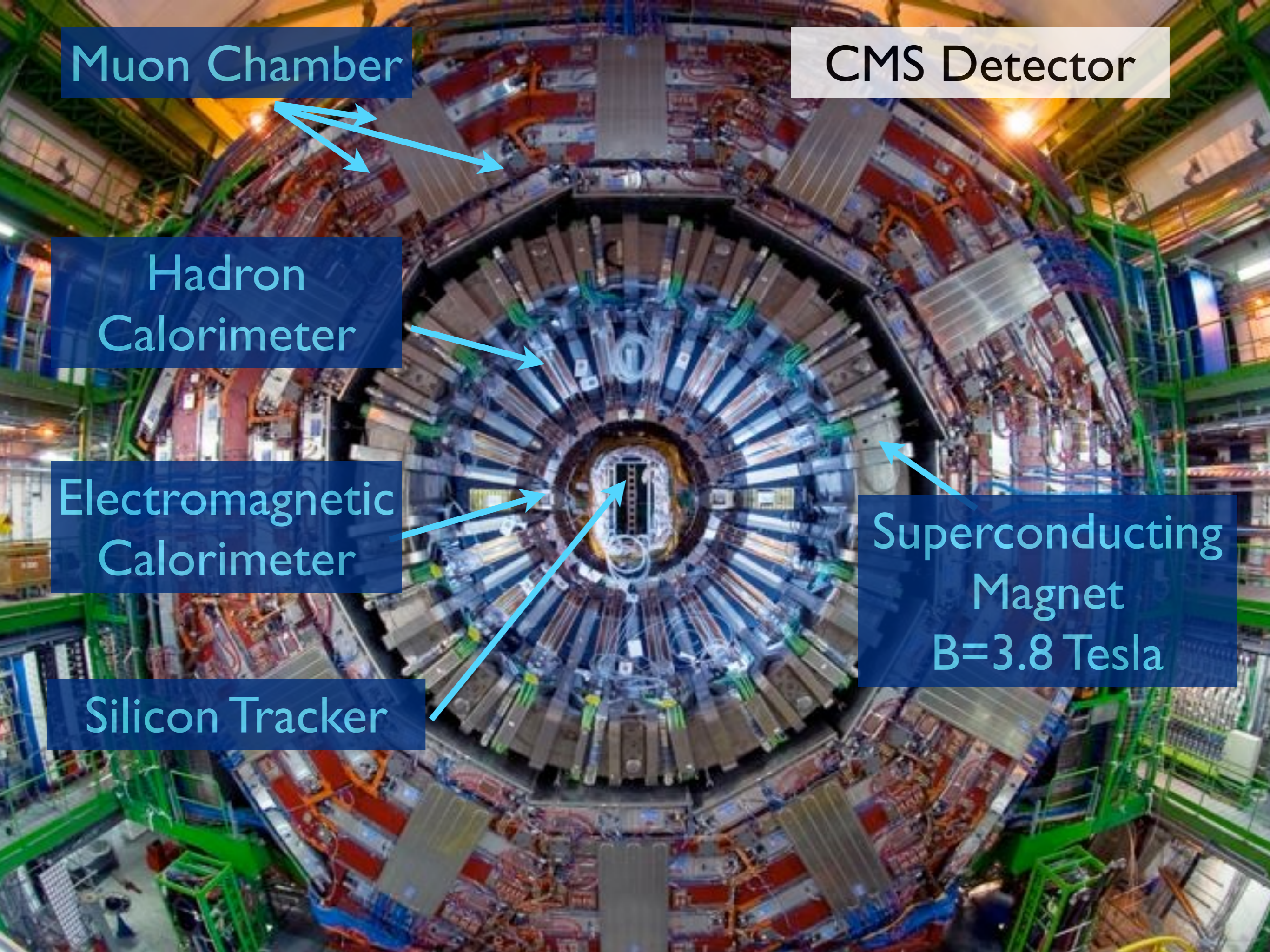
CMS Detector

Hadron
Calorimeter

Electromagnetic
Calorimeter

Superconducting
Magnet
 $B=3.8$ Tesla

Silicon Tracker



Muon Chamber

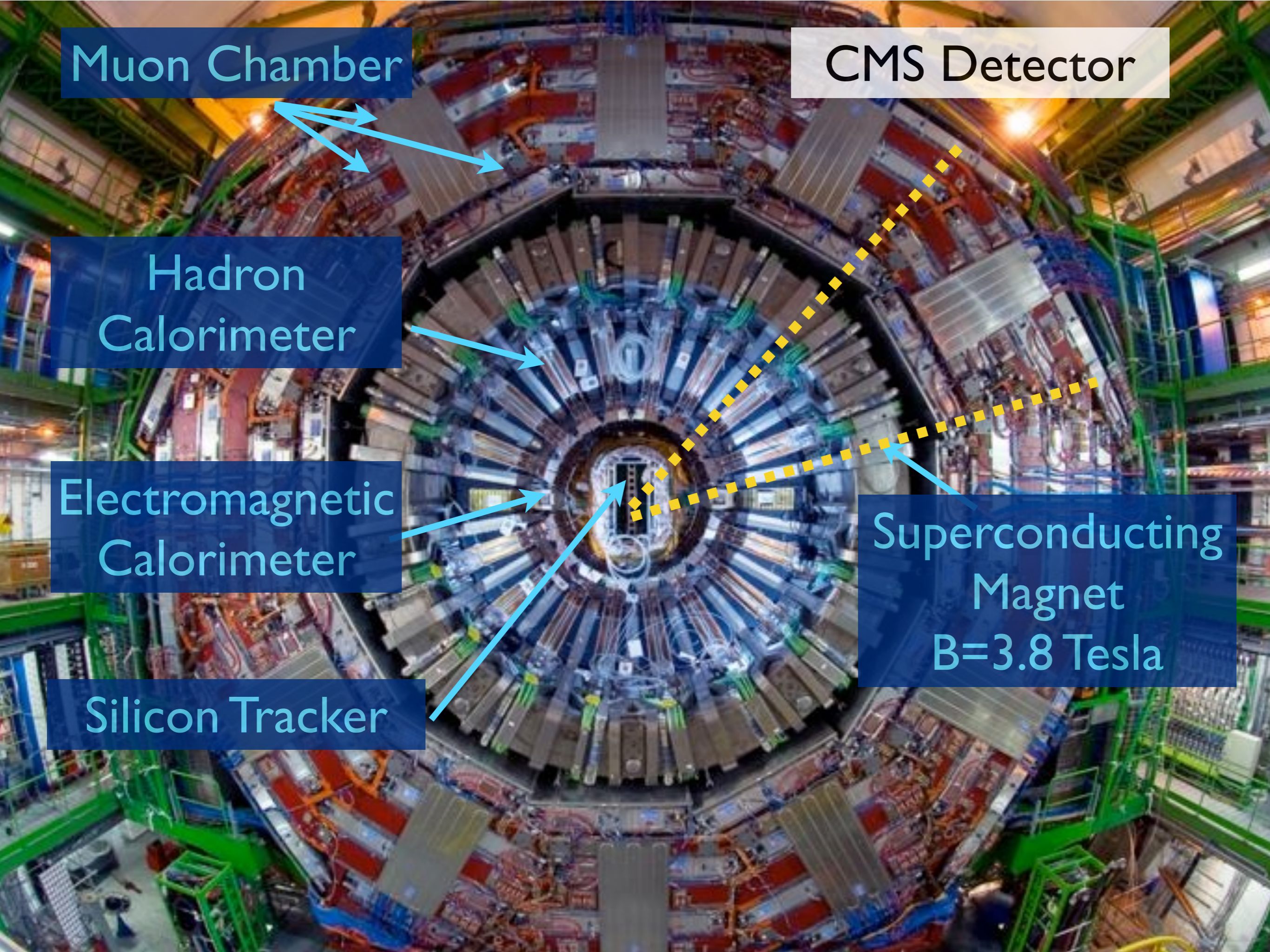
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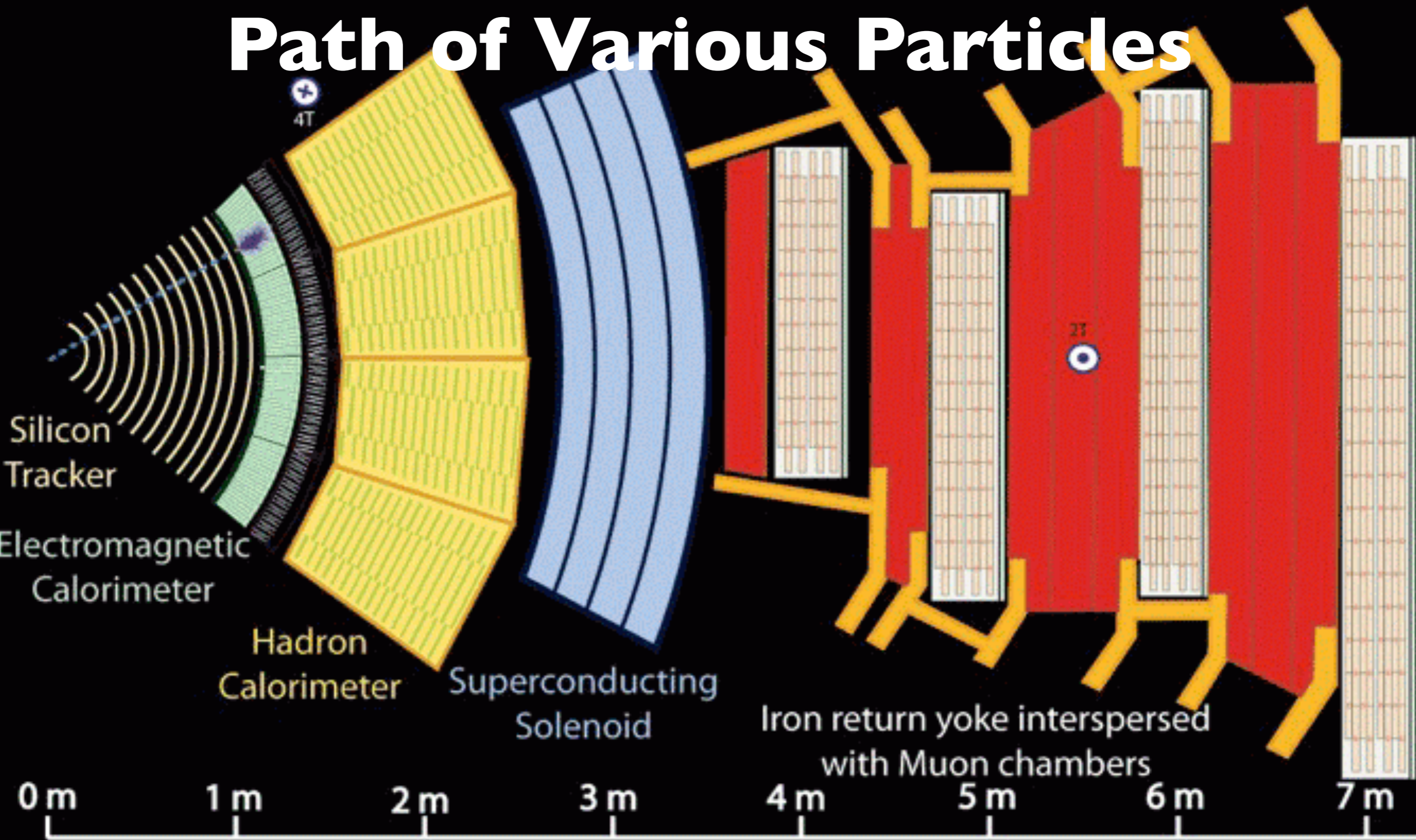
Electromagnetic
Calorimeter

Silicon Tracker

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Magnet
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Path of Various Particles



Key:

— Muon

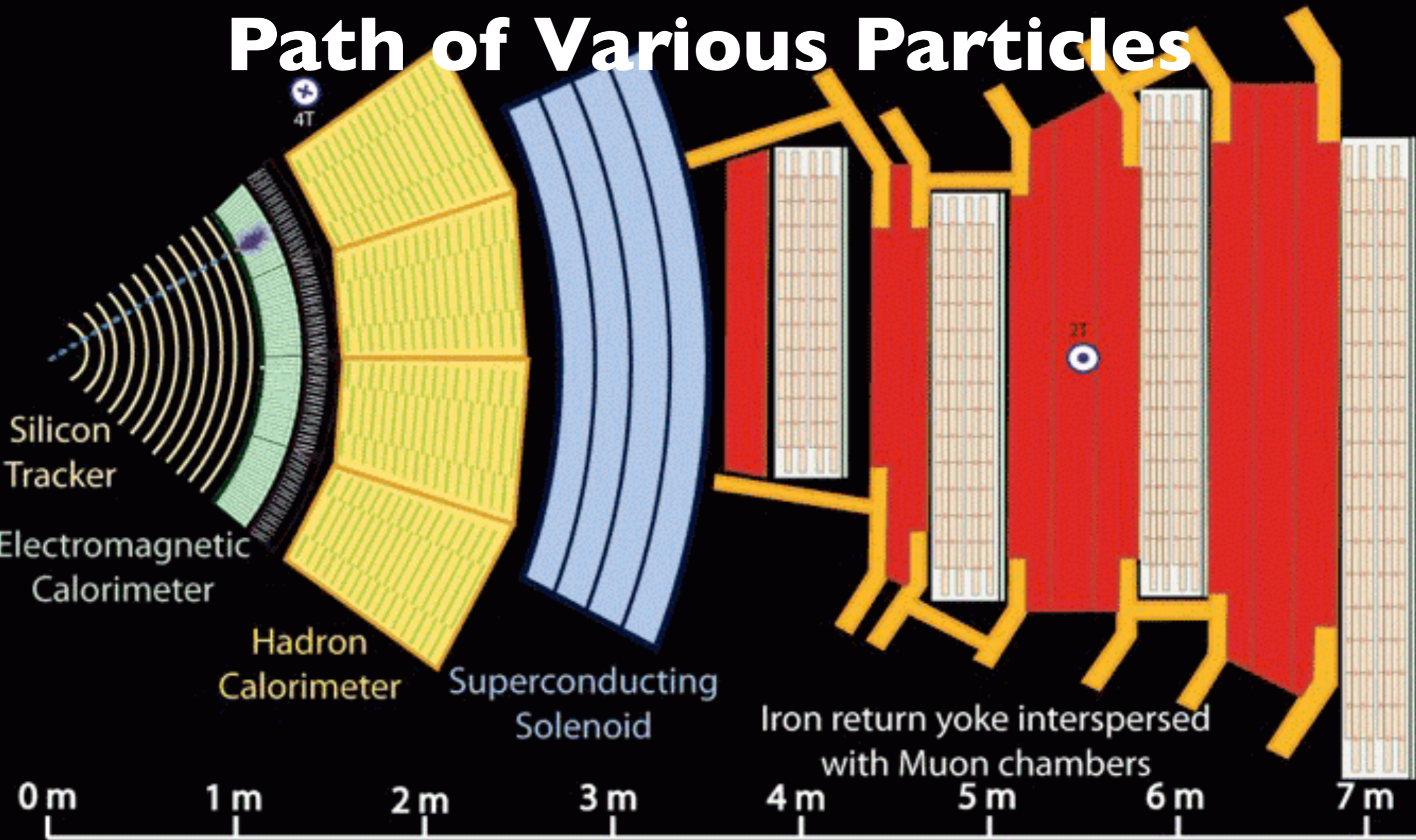
— Electron

— Charged Hadron (e.g. Pion)

- - - Neutral Hadron (e.g. Neutron)

- - - Photon

Path of Various Particles



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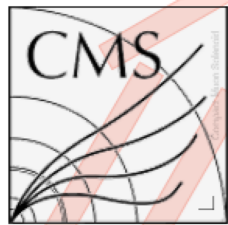
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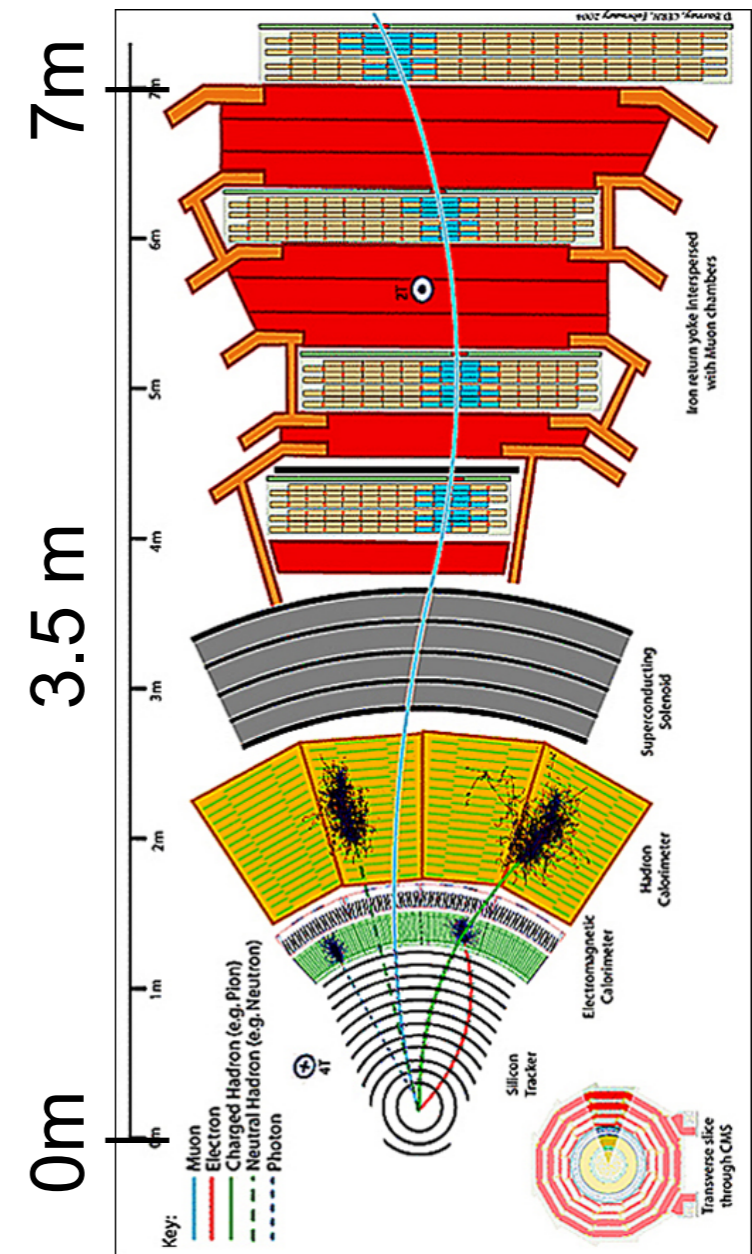
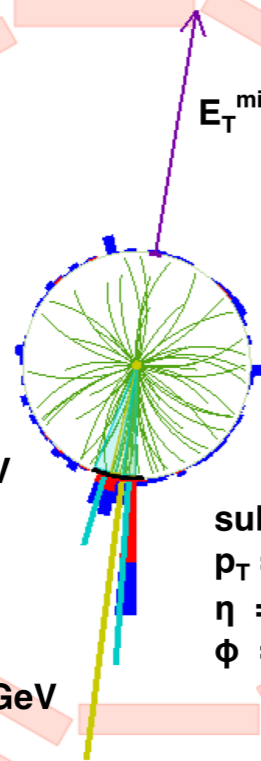
- - - Photon

What Is Dark Matter at Colliders?

- Neutral, weakly-interactive, massive, and stable on the distance-scales of tens of meters
- Dark matter appears as missing transverse momentum in collider detectors

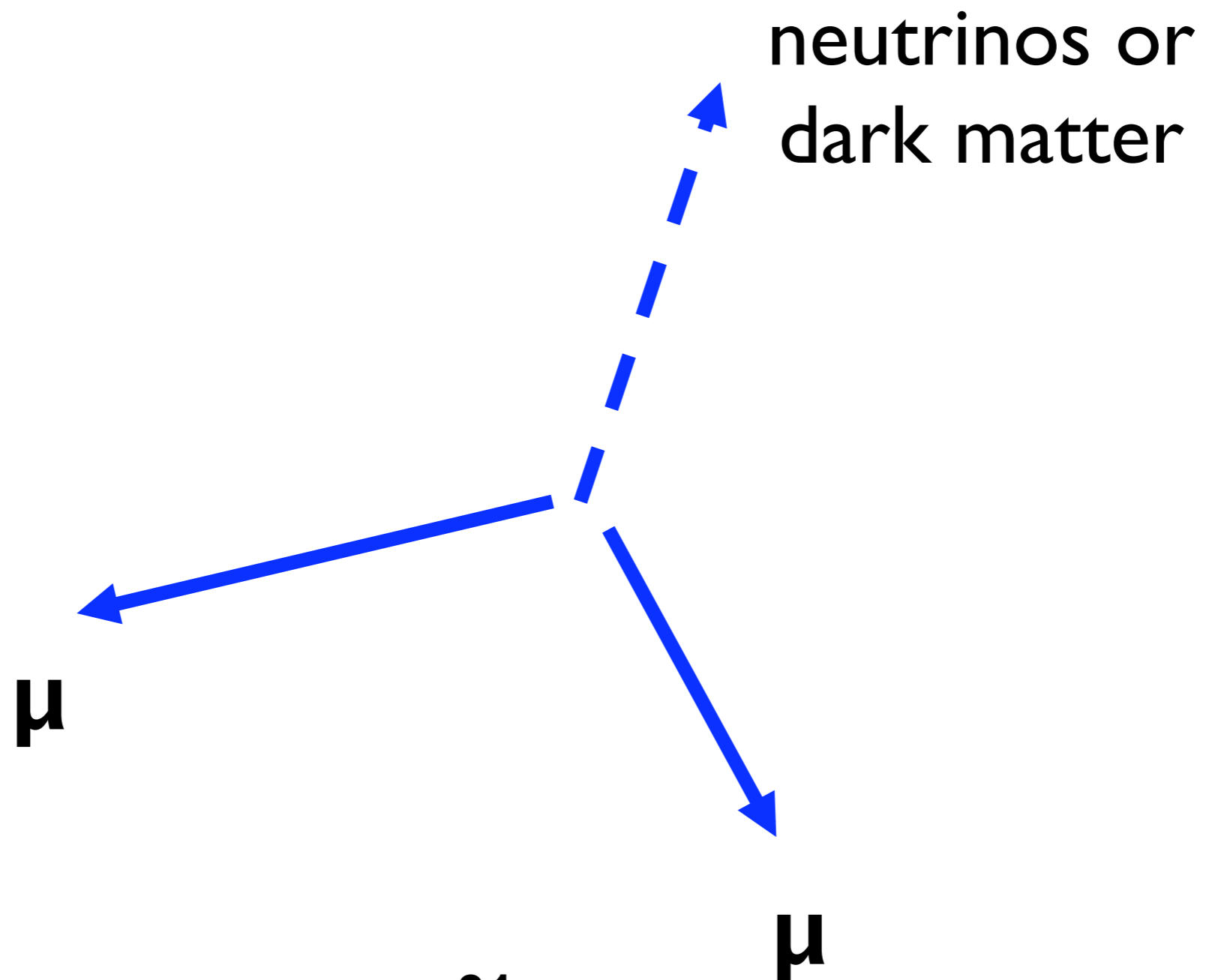


CMS experiment at LHC, CERN
Data recorded: Mon Sep 28 03:40:40 2015 CEST
Run/Event: 257645/597084610



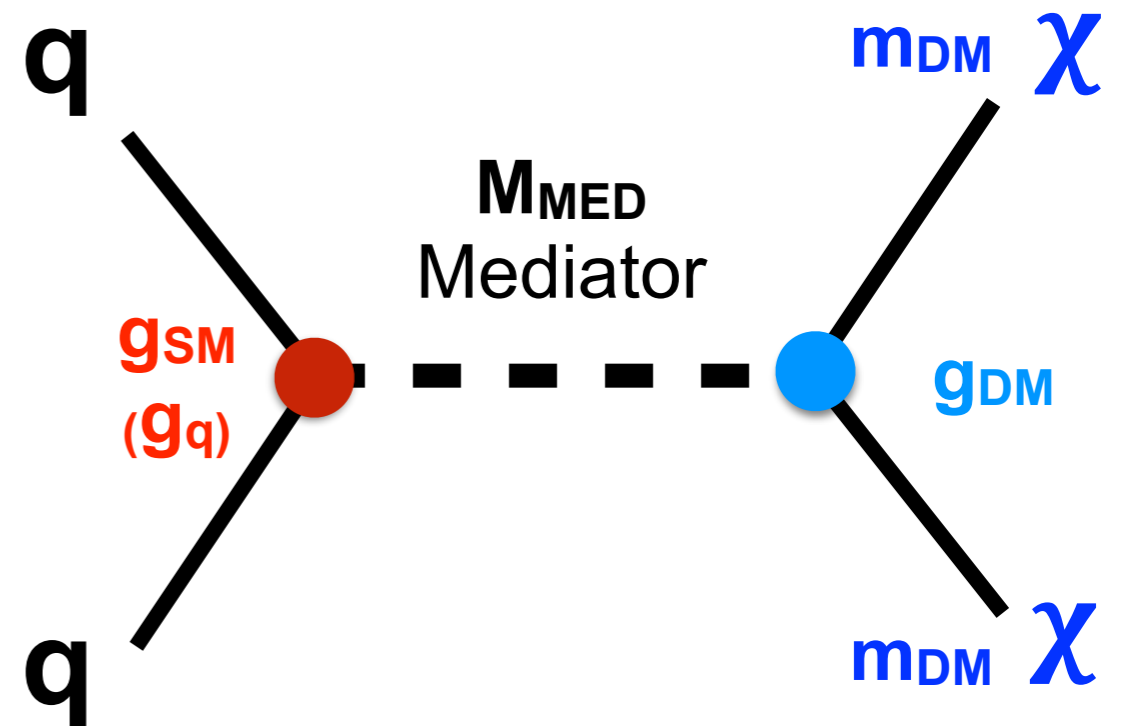
Missing Transverse Momentum

- The negative of the total transverse momentum of all observed particles in the detector



Simplified Models for Direct DM Production

- Mediator has minimal decay width
- Minimal set of parameters
 - coupling structure, M_{MED} , m_{DM} , g_{SM} (g_q), g_{DM}

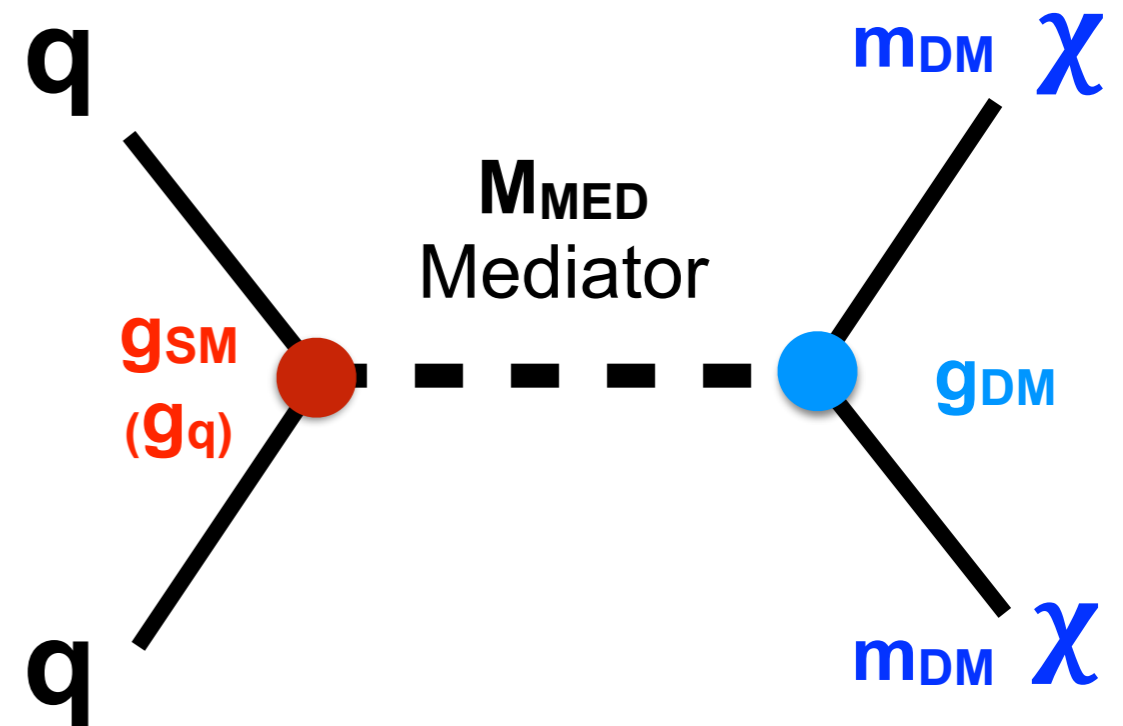


Simplified Models for Direct DM Production

Features of Mediators

	spin 0	spin 1
Charge Q	$Q_{\text{med}} = 0$ for s-channel	
Mass m	unknown	
Dark sector bosons similar to	H [1609.09079]	γ, Z, Z'
Lorentz structure	scalar 1 pseudosc. γ_5	vector γ^μ axial v. $\gamma^\mu \gamma_5$
Coupling "g"	\propto mass	\propto charge
Consequences	$m_b \gg m_d$	$Q_b = Q_d$

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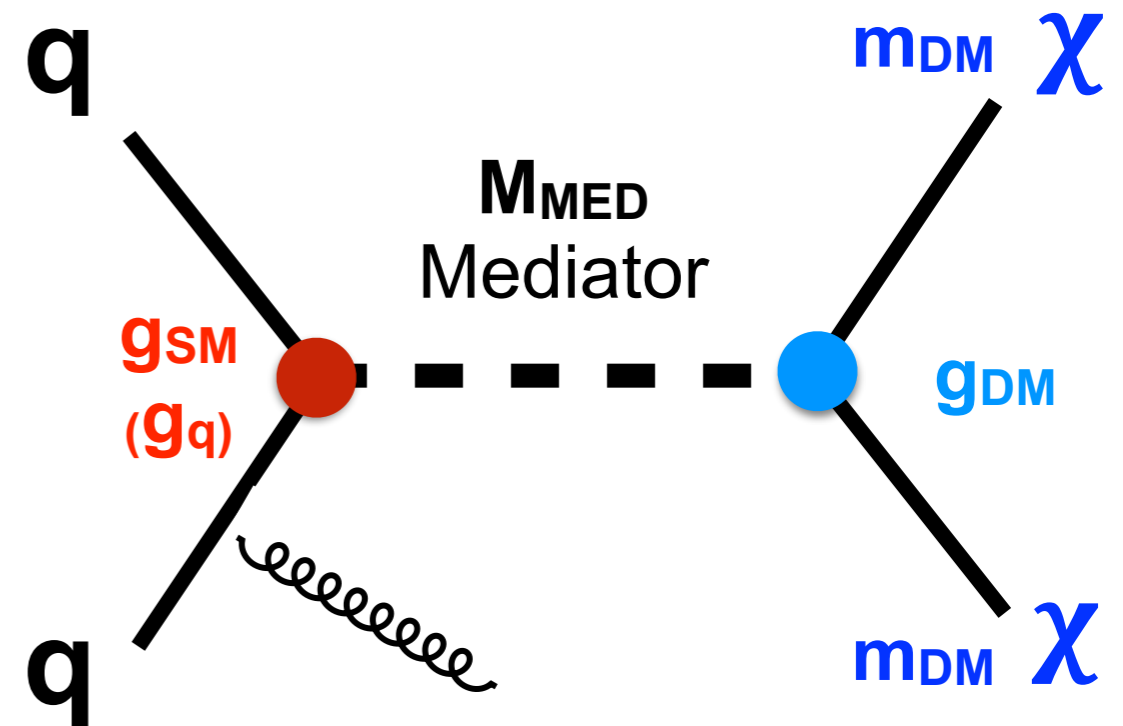
Tae Min Hong, LHCP 2017

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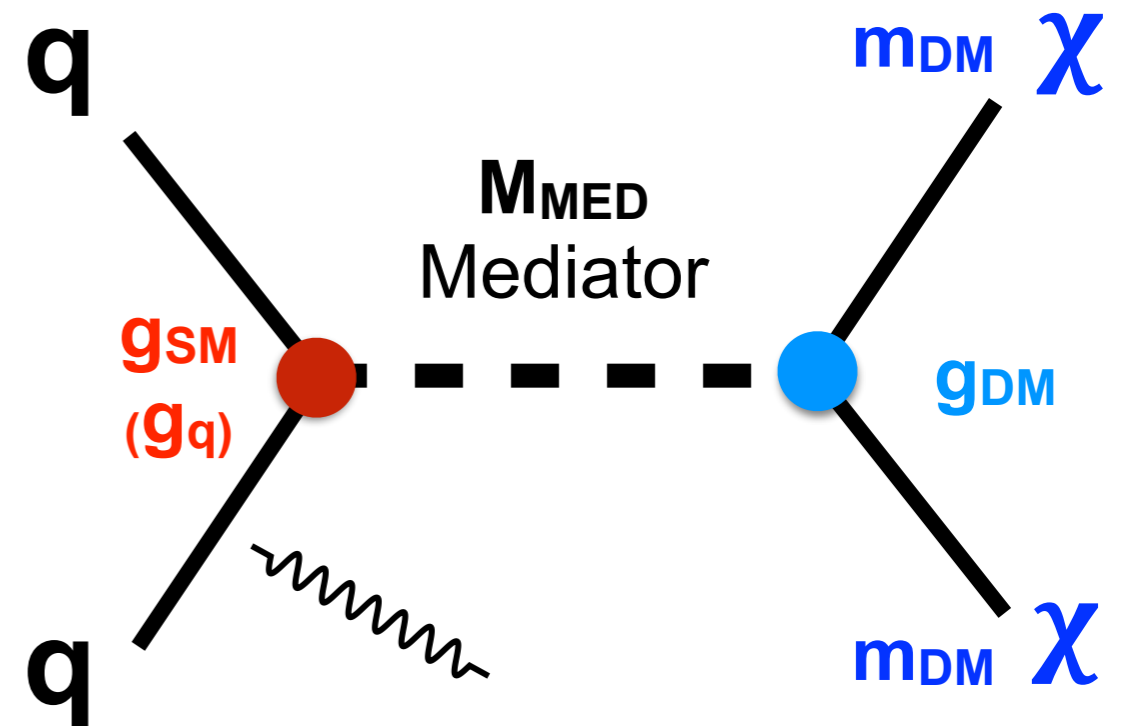
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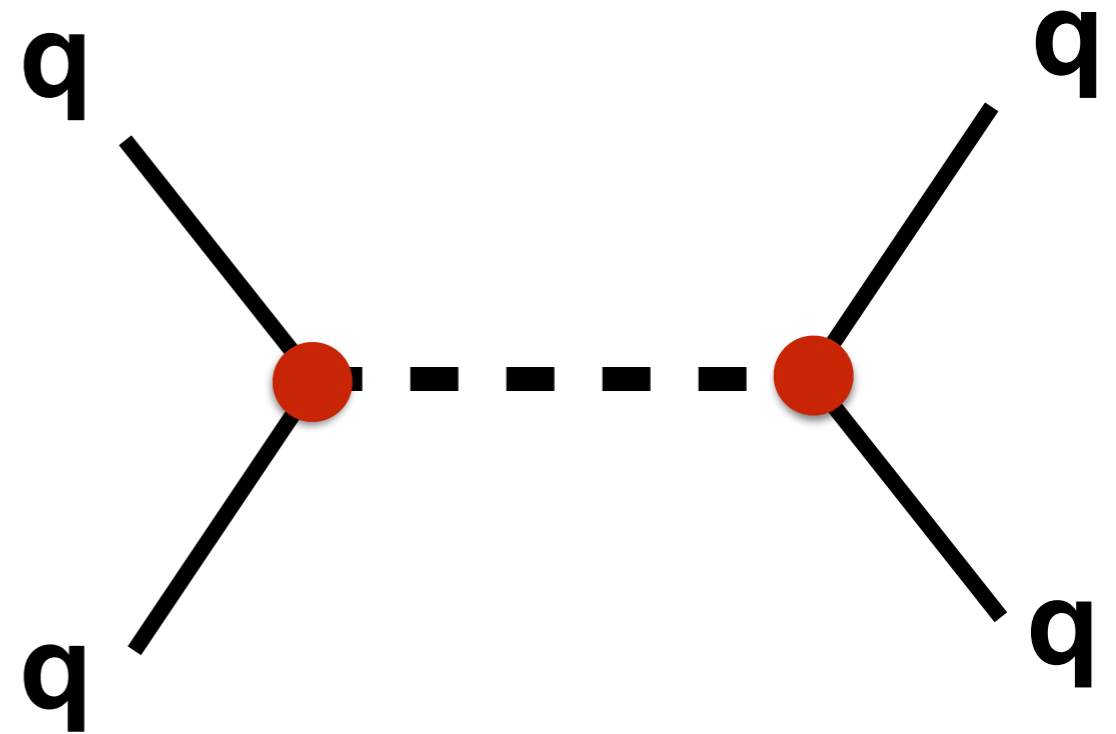
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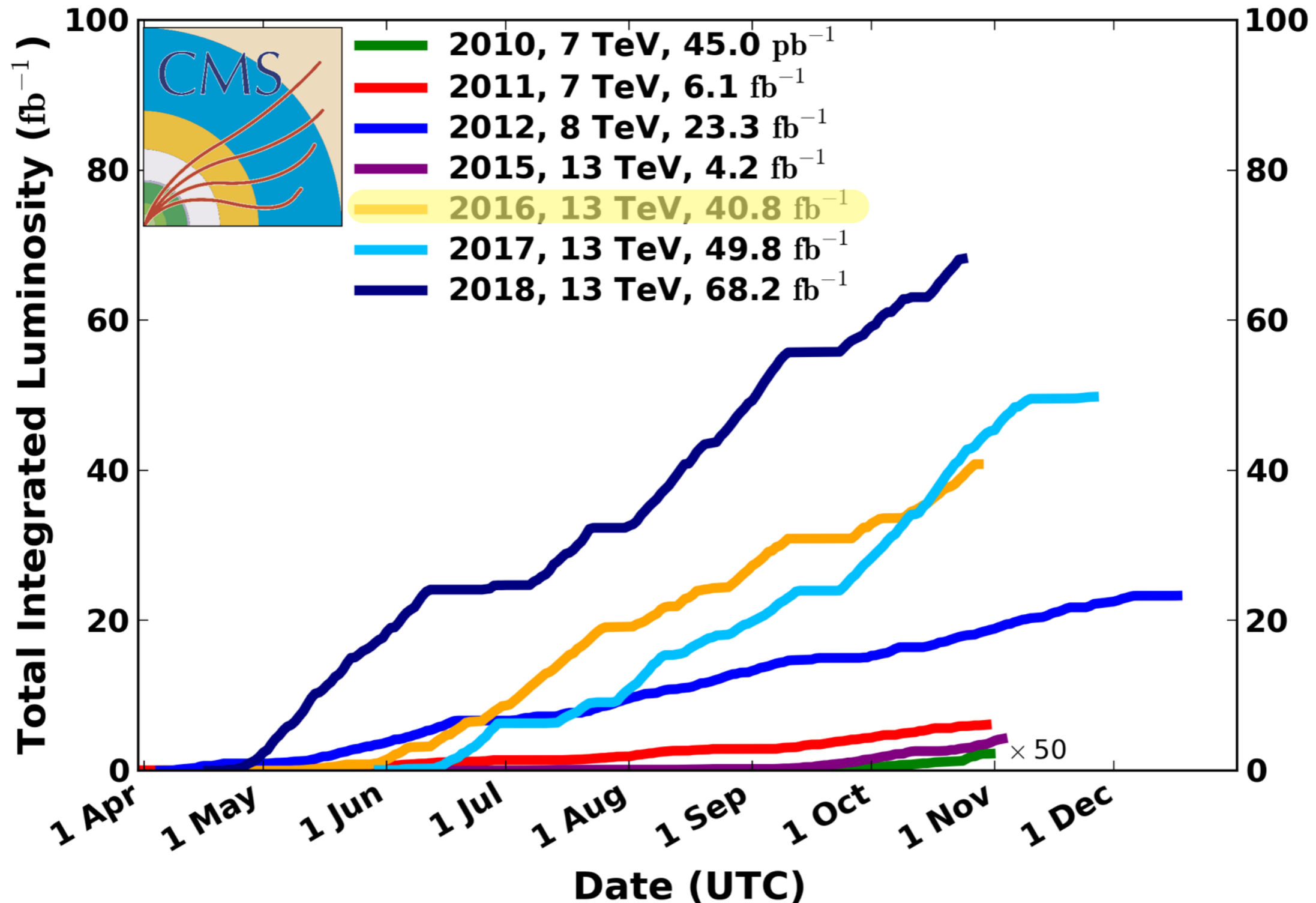


Tae Min Hong, LHCP 2017

Amount of Data We Use

CMS Integrated Luminosity, pp

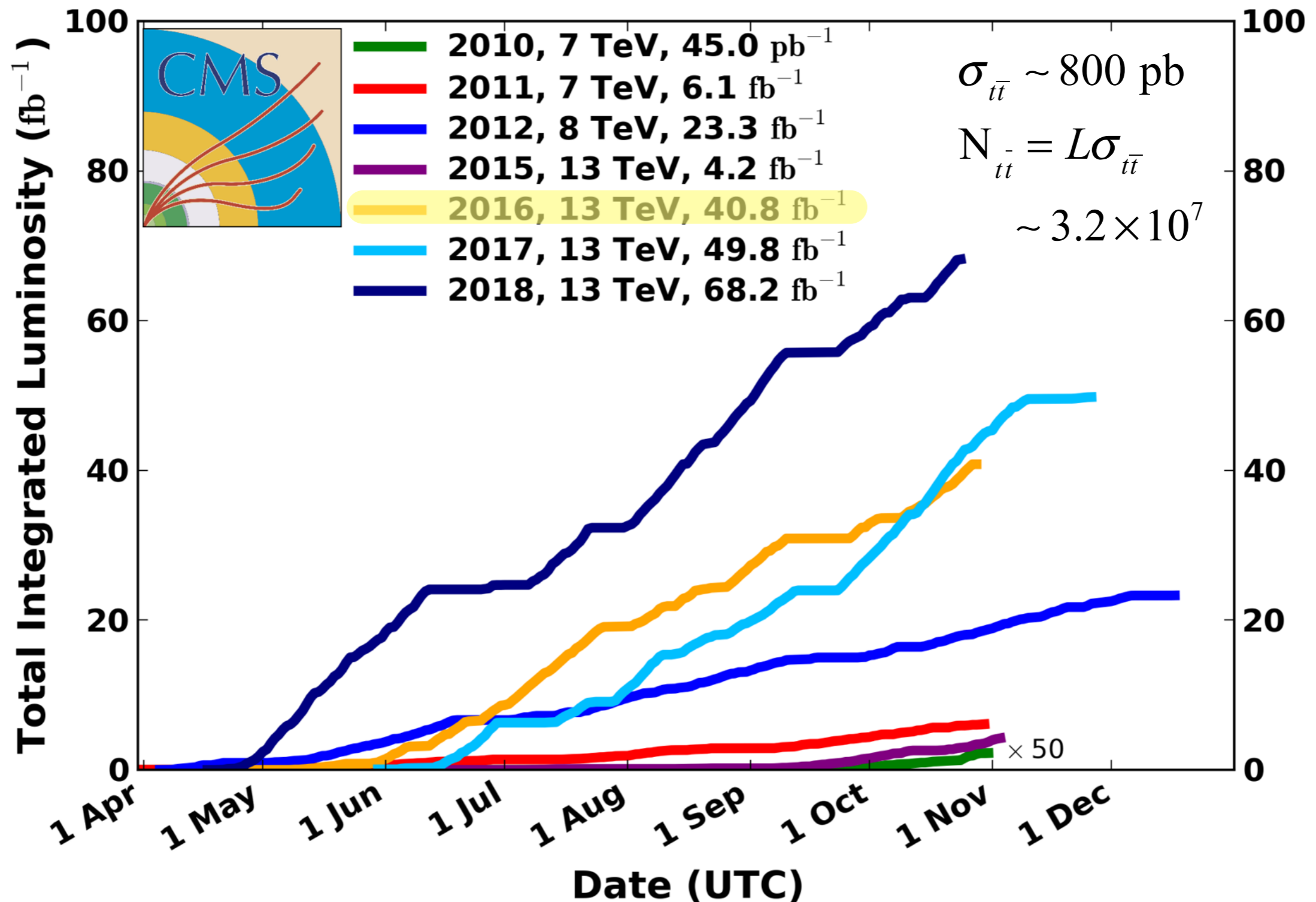
Data included from 2010-03-30 11:22 to 2018-10-24 04:00 UTC



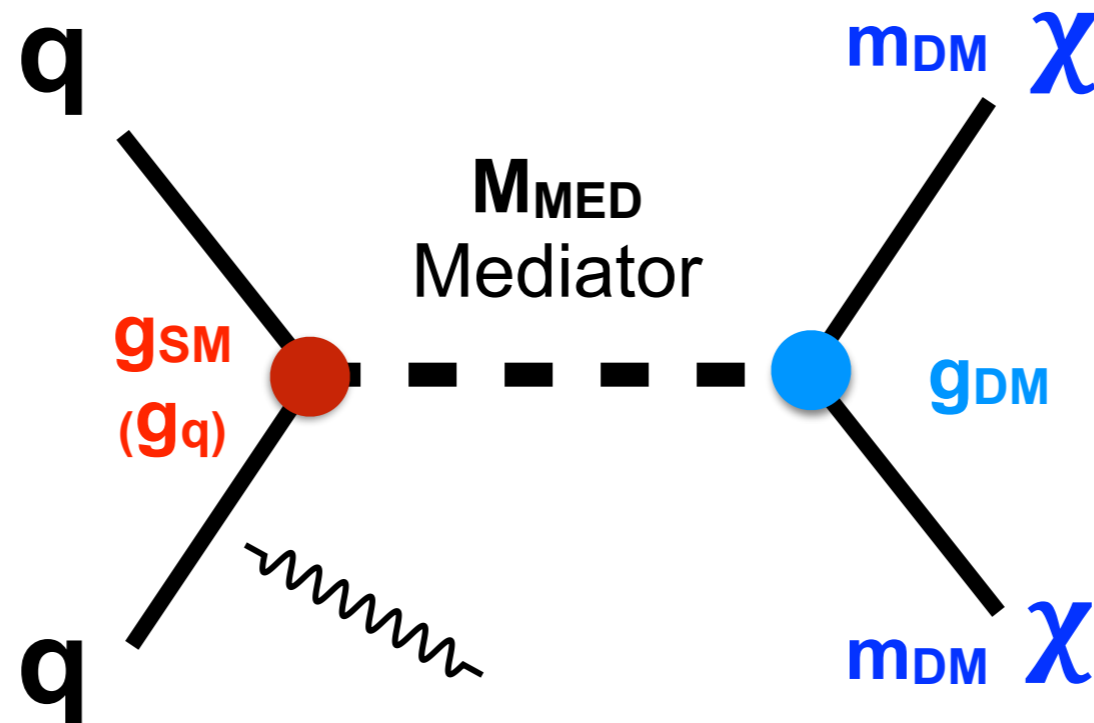
Amount of Data We Use

CMS Integrated Luminosity, pp

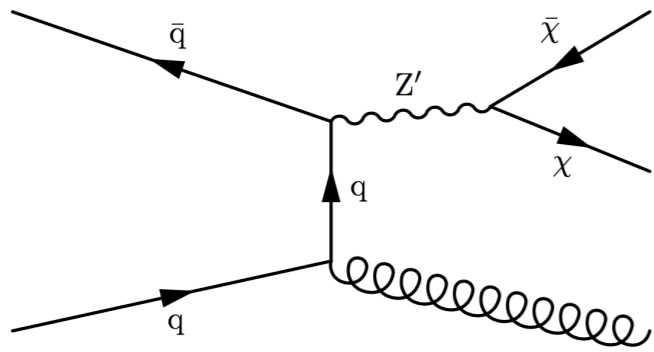
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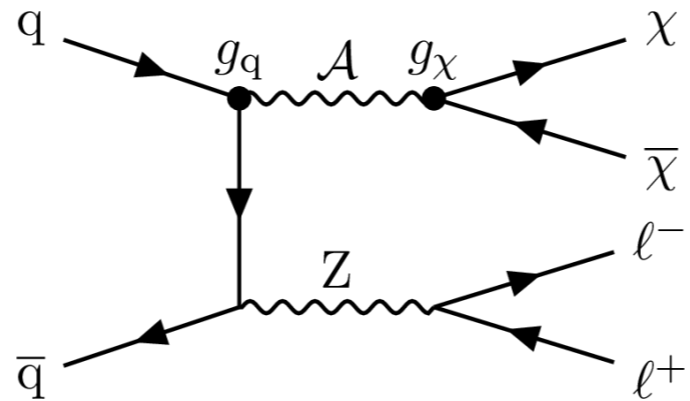
DM Searches with Missing Transverse Momentum Signatures



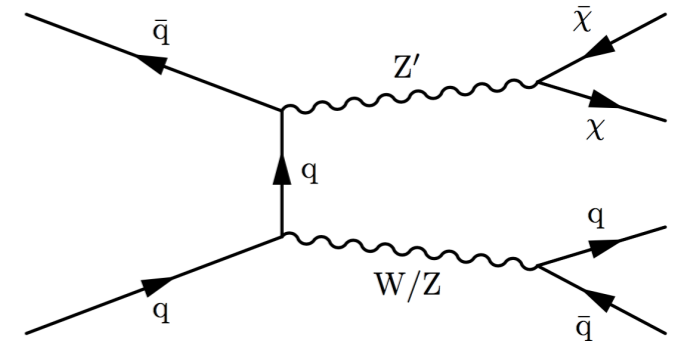
Mono-X Diagrams of Direct DM Production



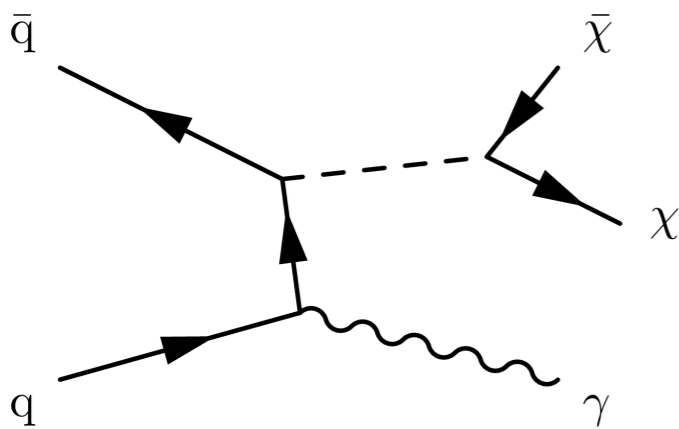
Mono-jet



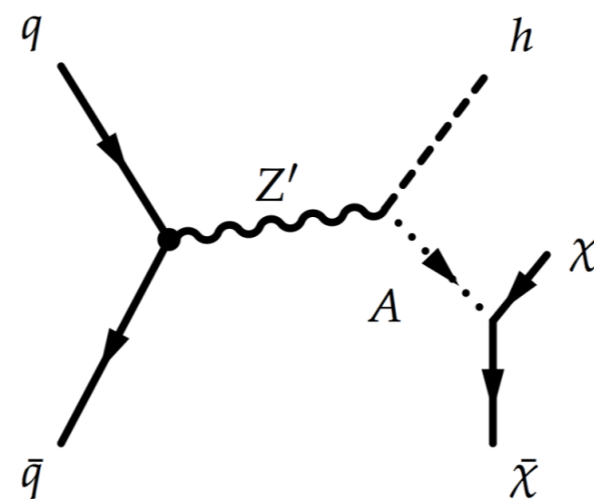
Mono-Z(leptonic)



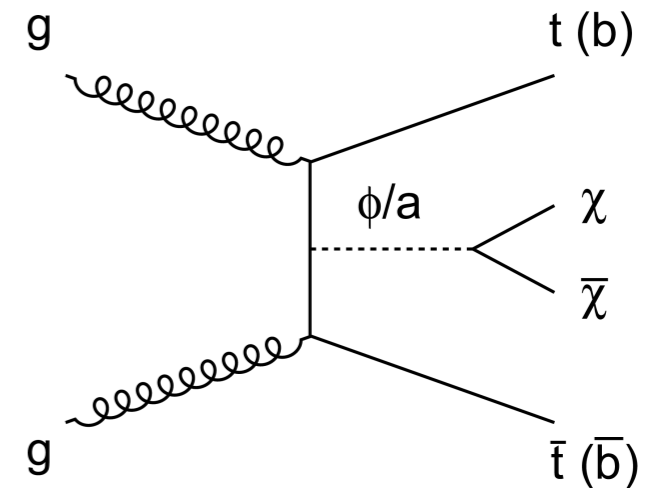
Mono-W/Z(hadronic)



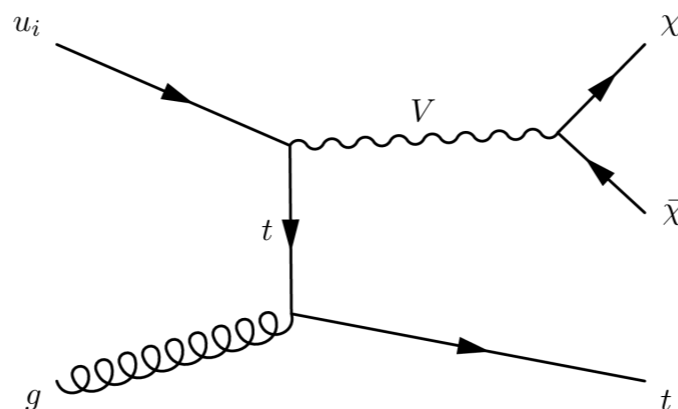
Mono-photon



Mono-h (bb, $\gamma\gamma$)



Mono-tt/bb



Mono-top

Challenges of Missing Transverse Momentum

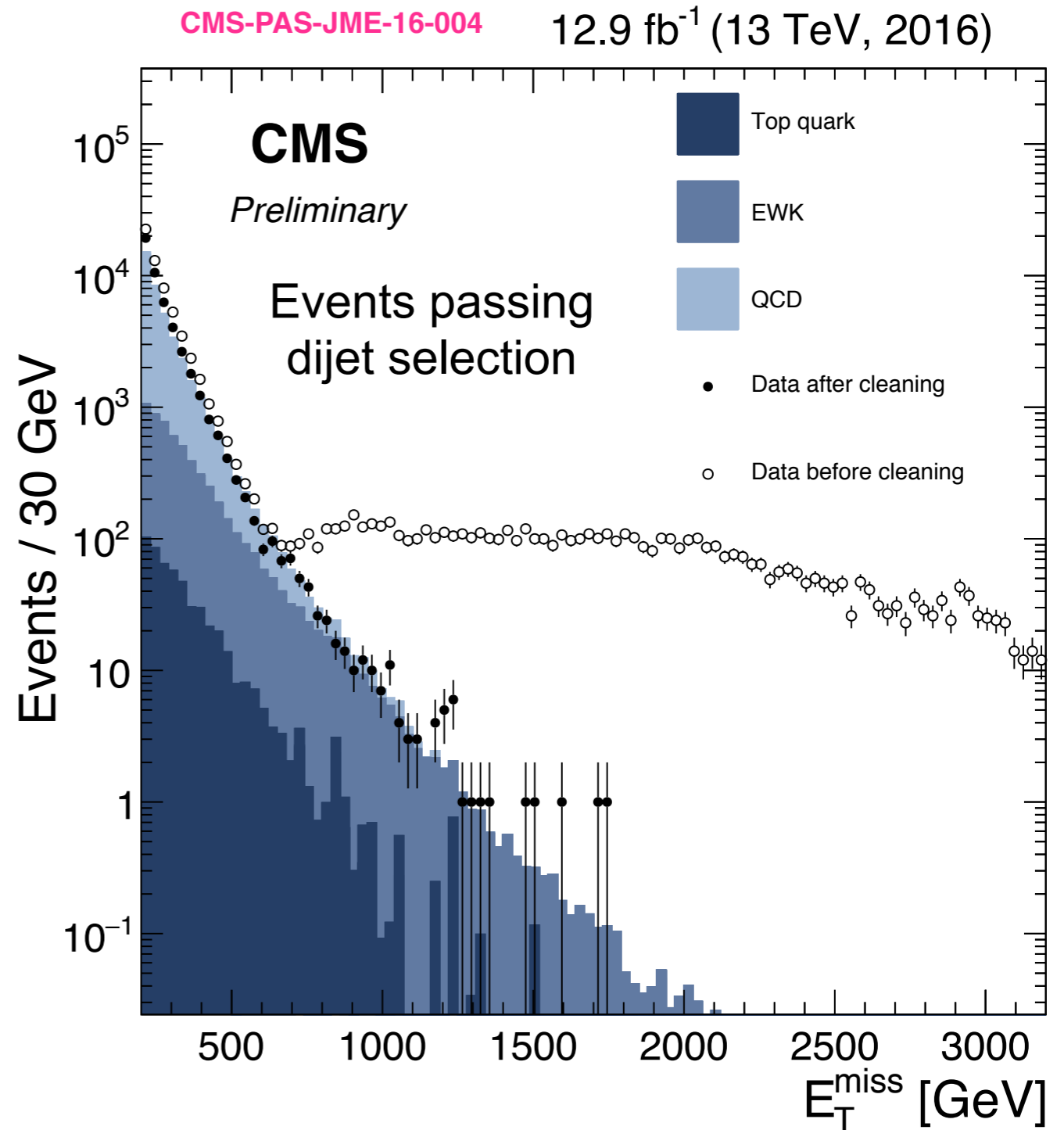
- Anomalous high MET can be due to:
 - Particles striking sensors in the ECAL photodetectors
 - Beam halo
 - Dead cells in ECAL or HCAL
 - Noise in ECAL or HCAL



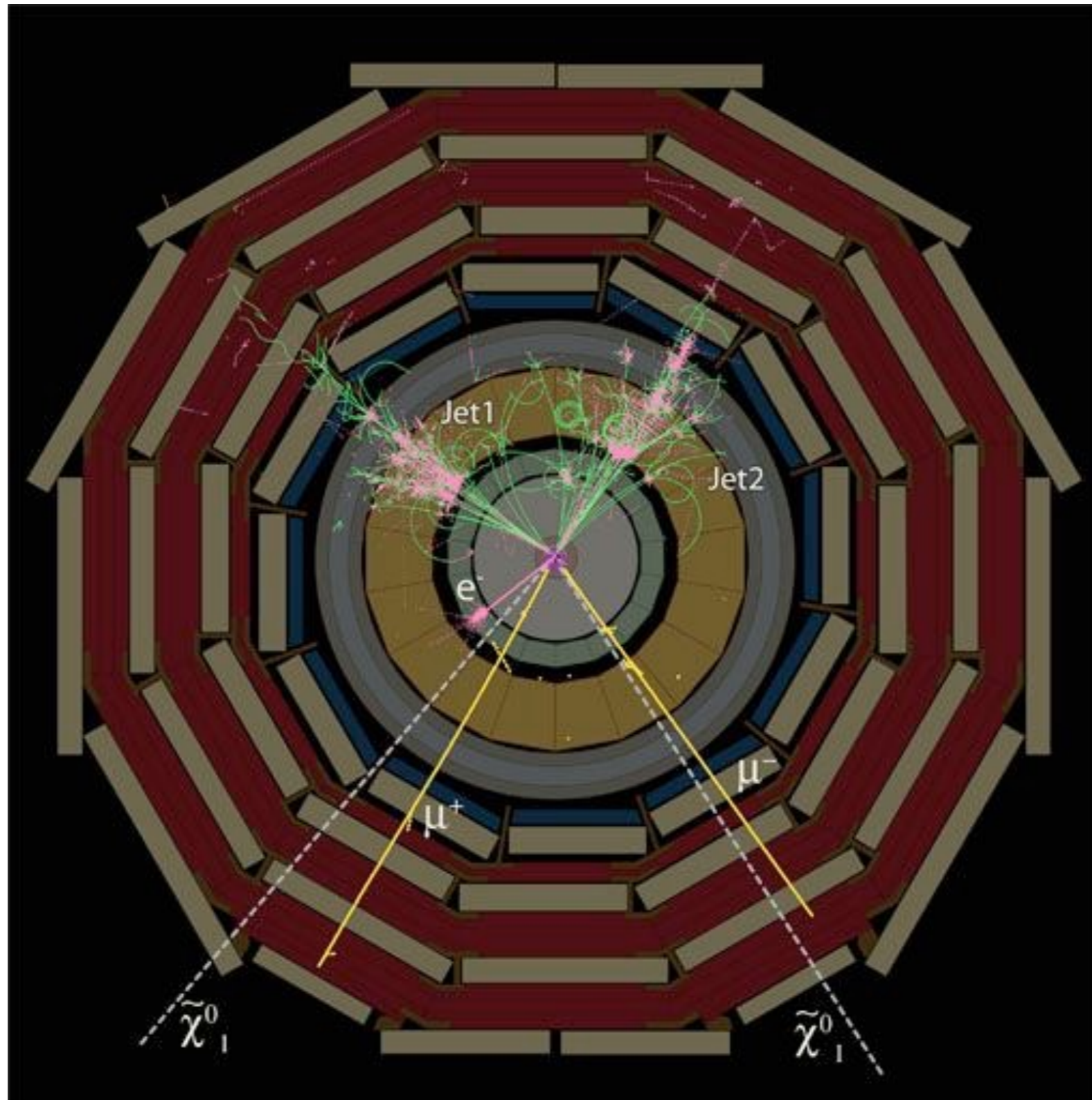
Raman Khurana



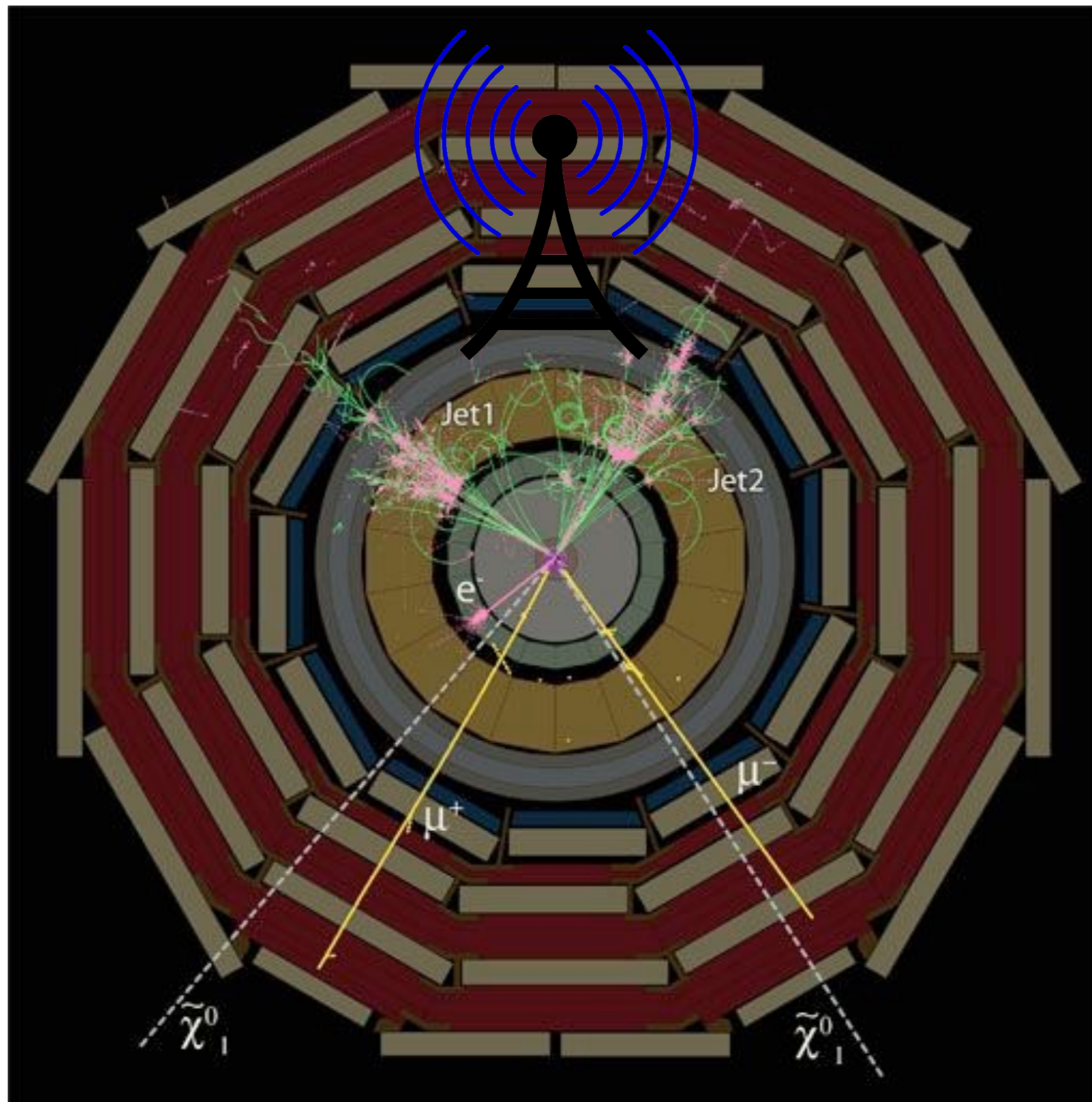
Ching-Wei Chen



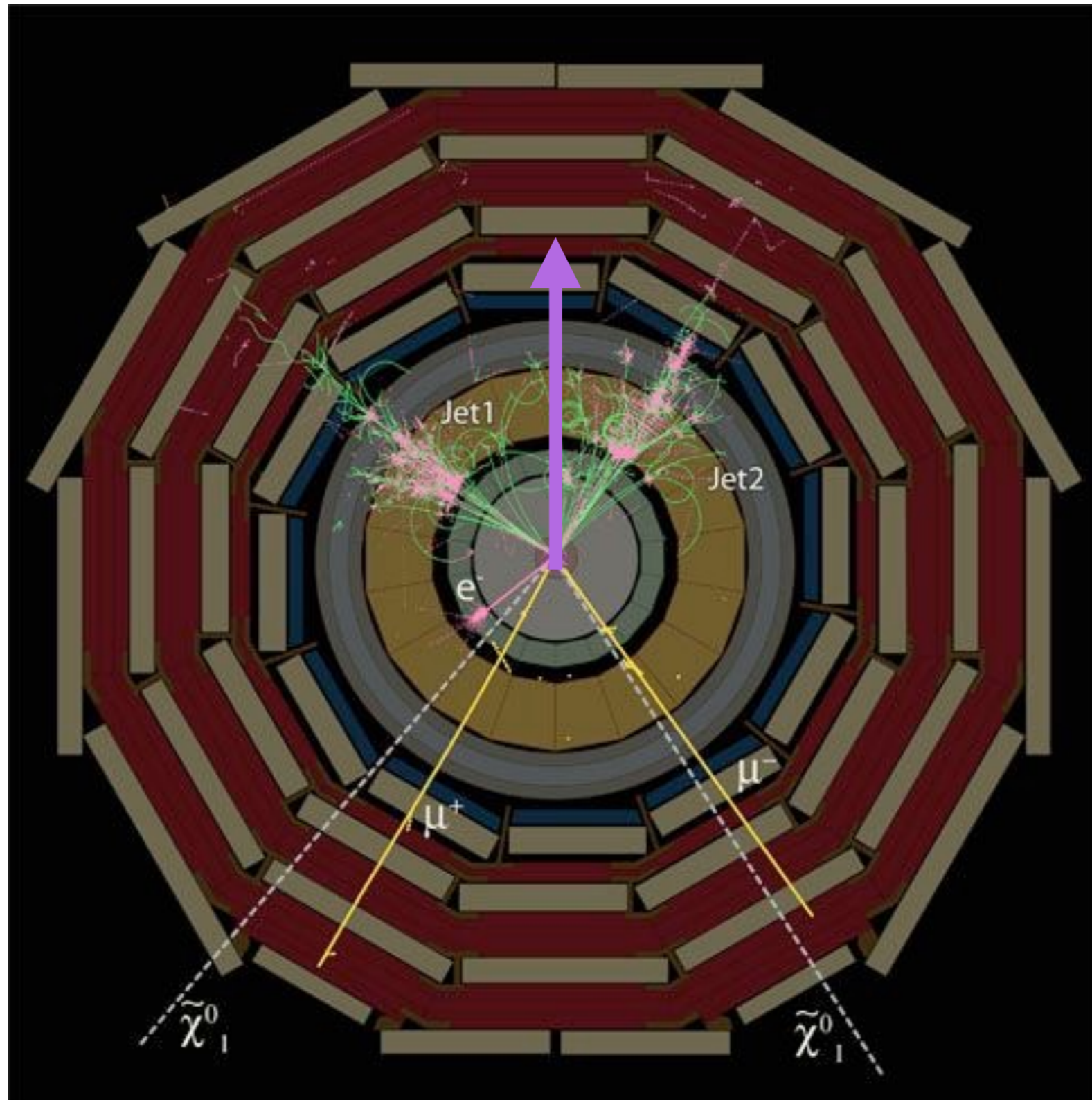
Fake Missing Transverse Momentum: Noise



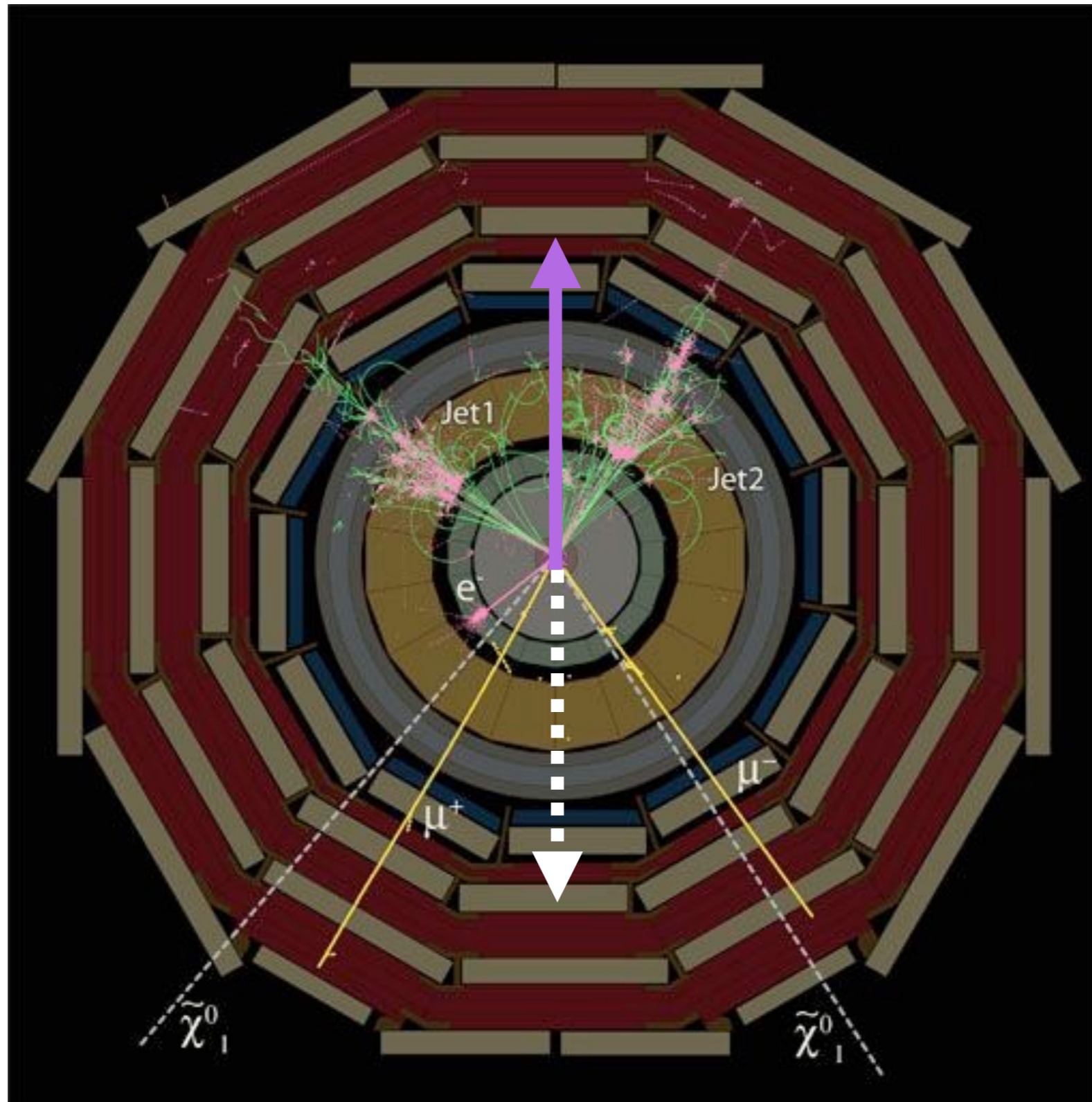
Fake Missing Transverse Momentum: Noise



Fake Missing Transverse Momentum: Noise



Fake Missing Transverse Momentum: Noise

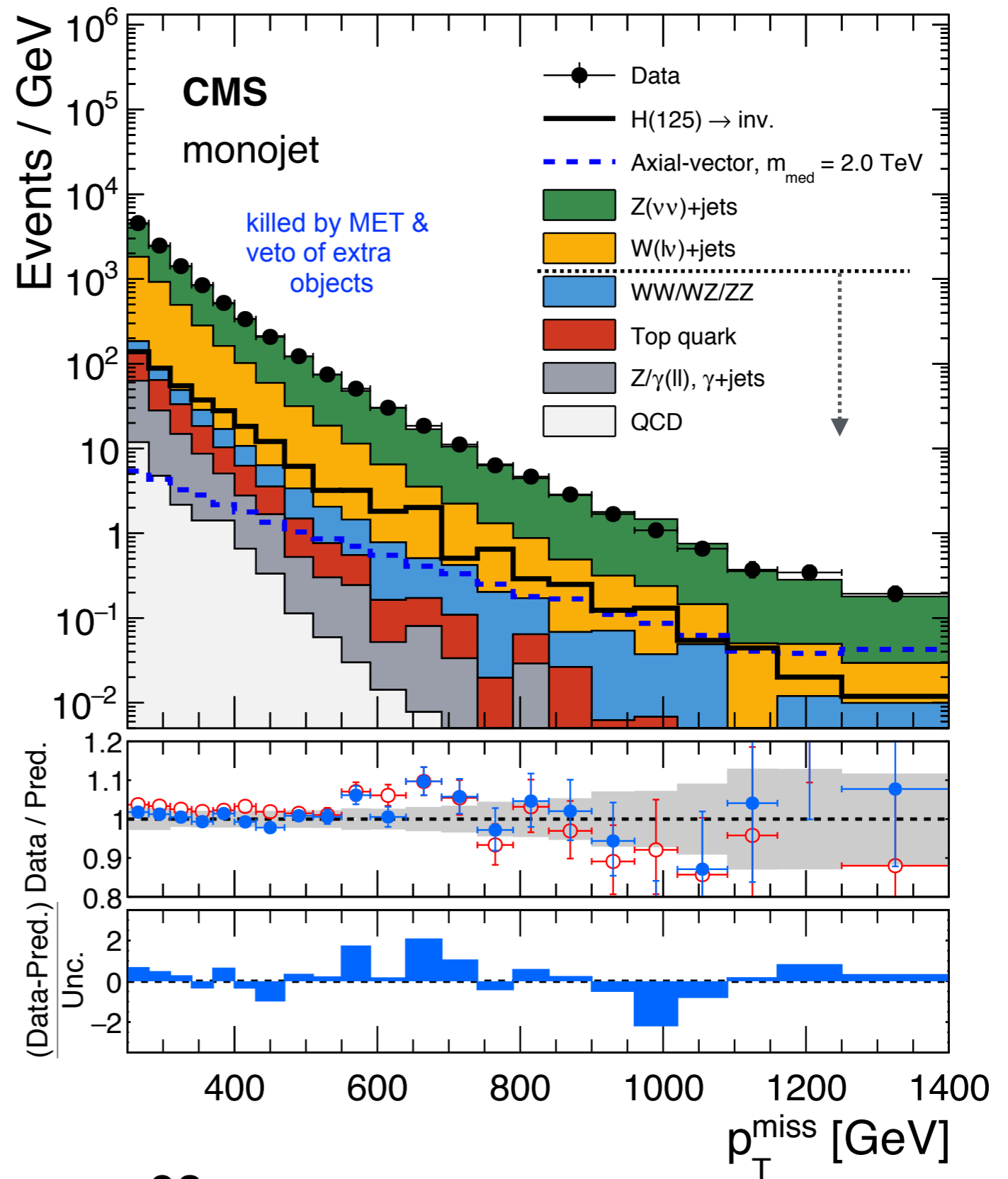


Mono-X Searches in Hadronic Final State

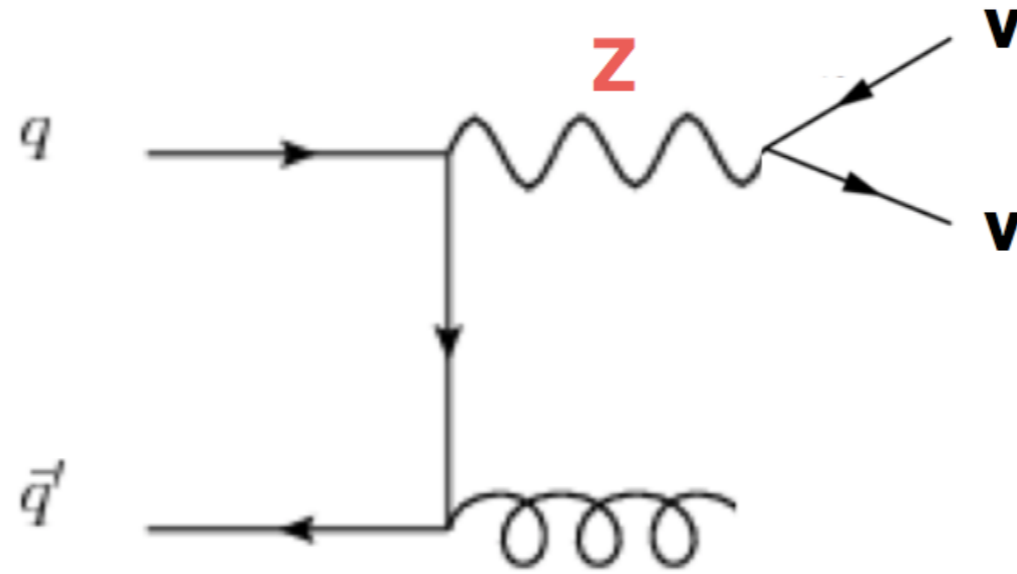
PRD 97, 092005 (2018)

35.9 fb⁻¹ (13 TeV)

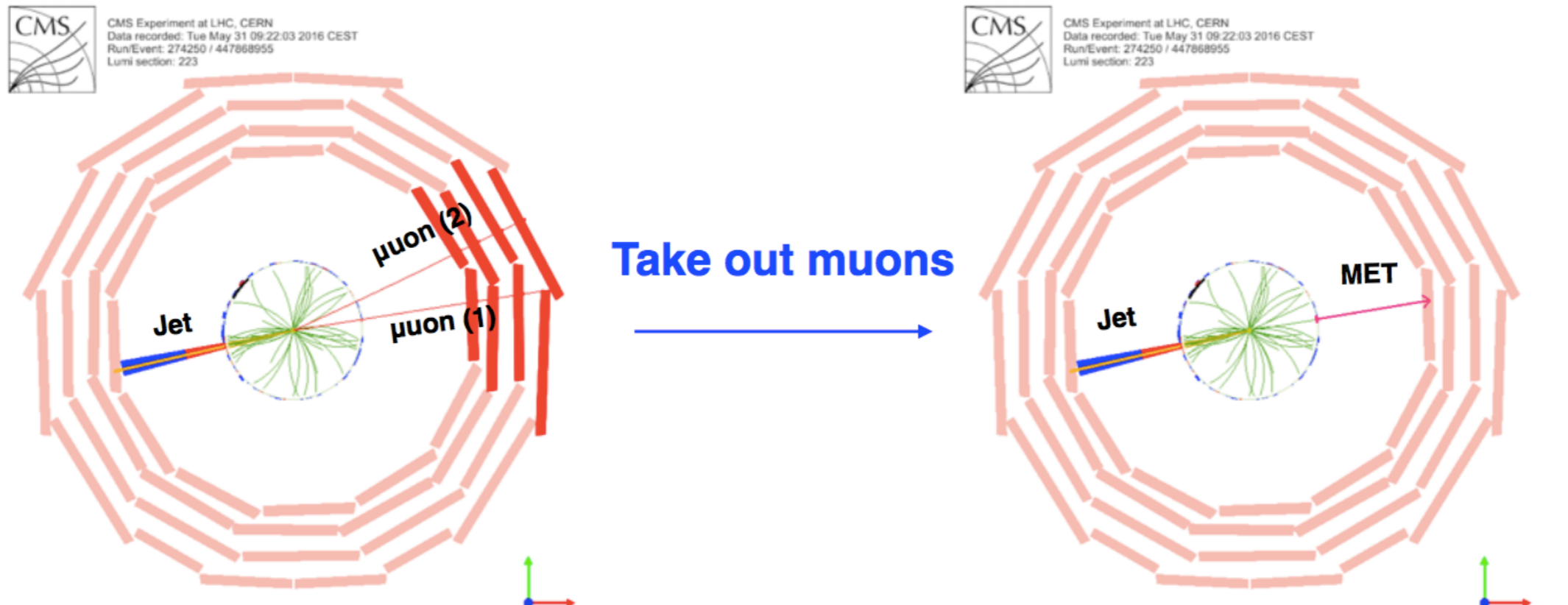
- Rely on MET triggers (offline MET cut ≥ 200 GeV)
- Major background from $Z(\rightarrow \nu\nu)+\text{jets}$, $W(\rightarrow l\nu)+\text{jets}$



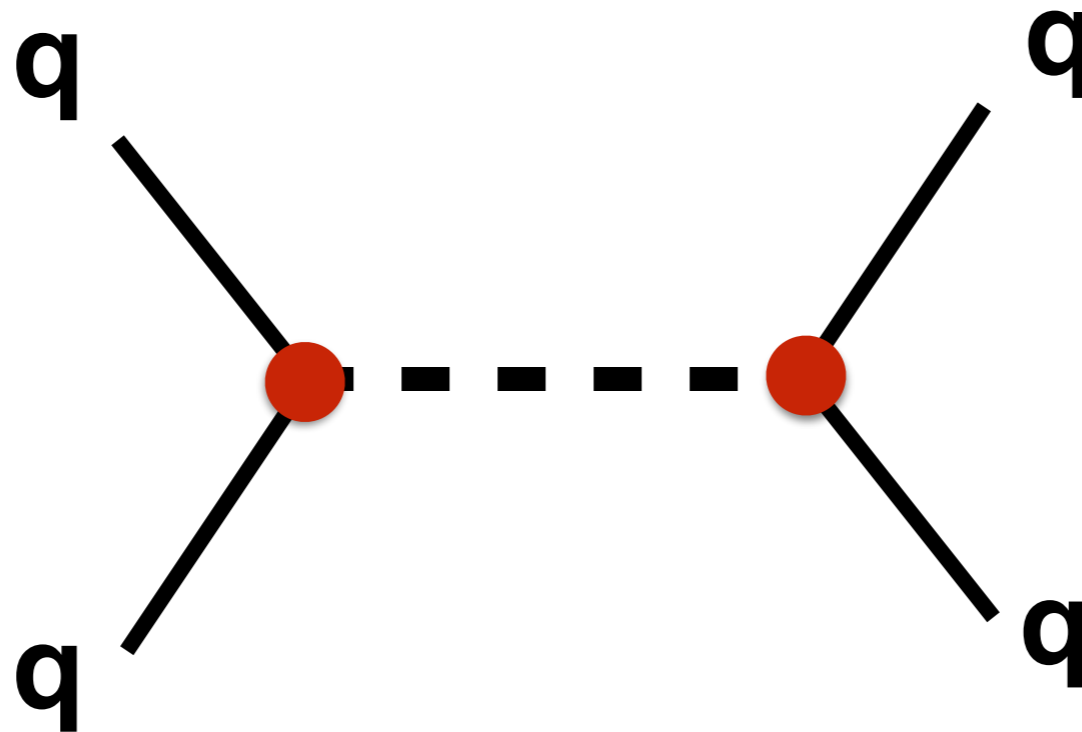
Estimation of Z+Jets Background



If we remove the muons from a $Z \rightarrow \mu\mu$ event, it mimics a $Z \rightarrow \nu\nu$ event

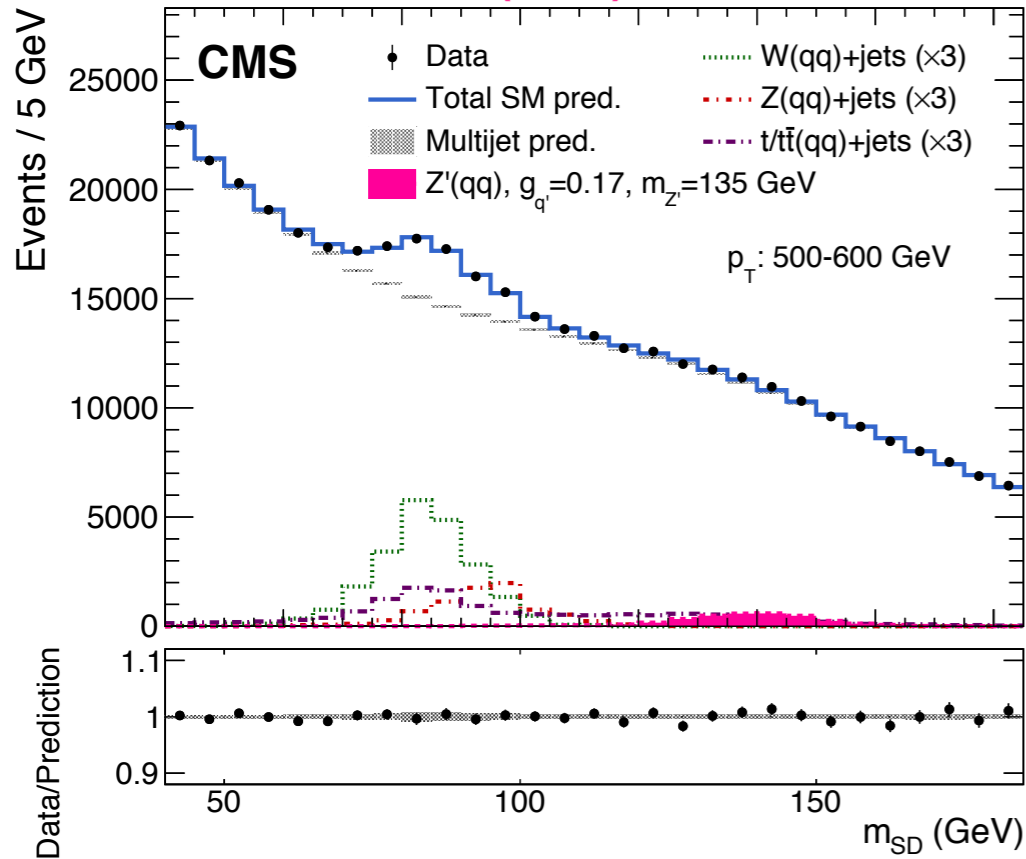


Searches for Visible Mediator Decays

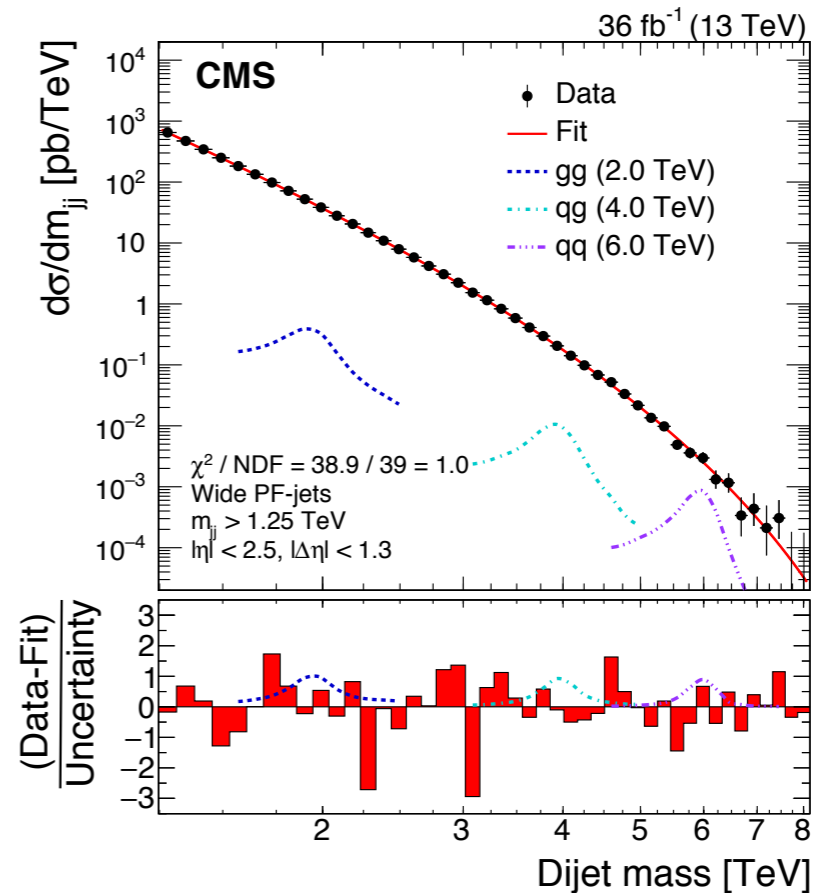
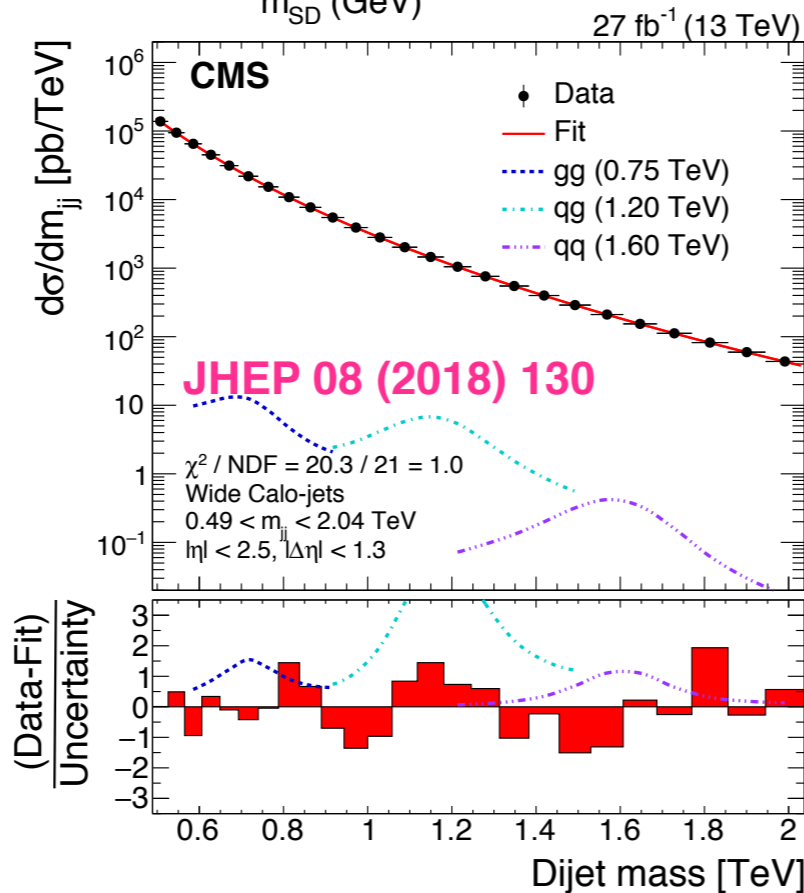
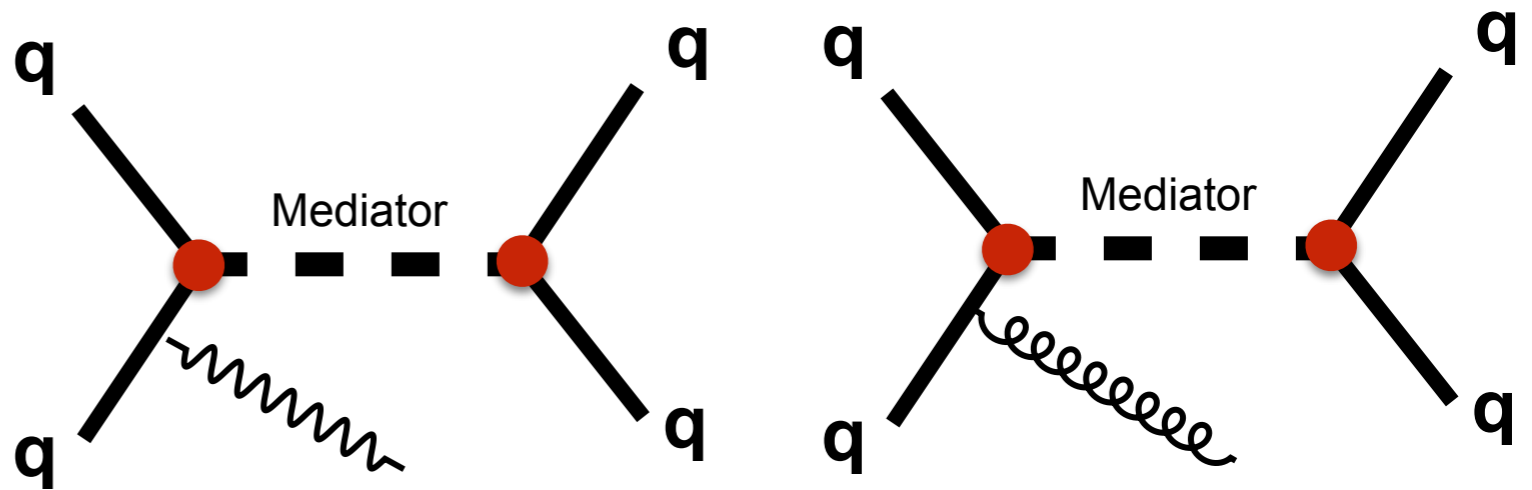


Visible Mediator Searches

JHEP 01 (2018) 097 35.9 fb⁻¹ (13 TeV)

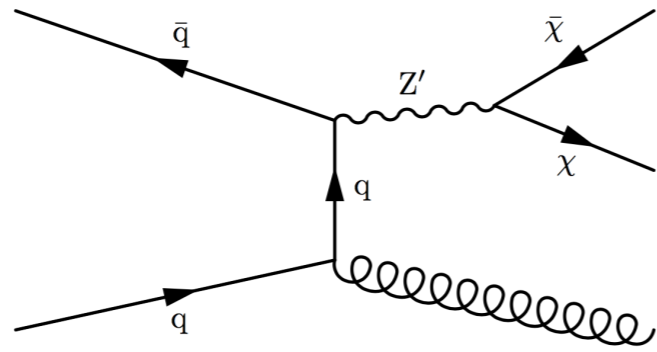


- high- p_T/H_T trigger for large- M_{jj} , ISR γ /jet tag or data with only trigger-level objects (data scouting) for small- M_{jj}

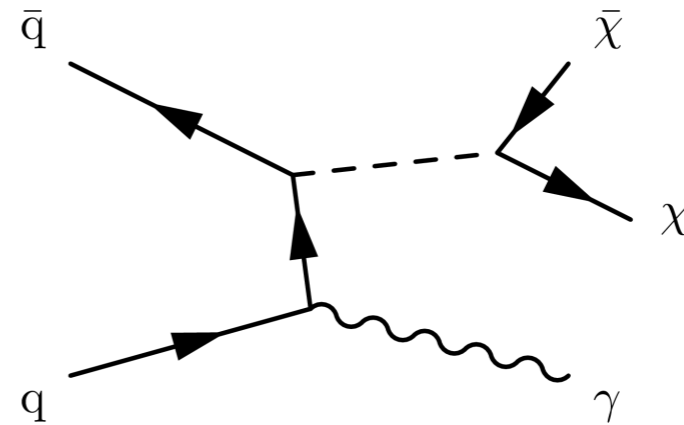


Result Interpretation

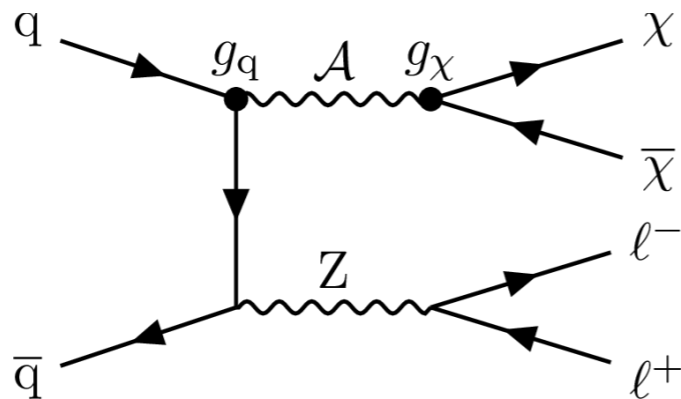
Mono-X With Vector/Axial Mediators



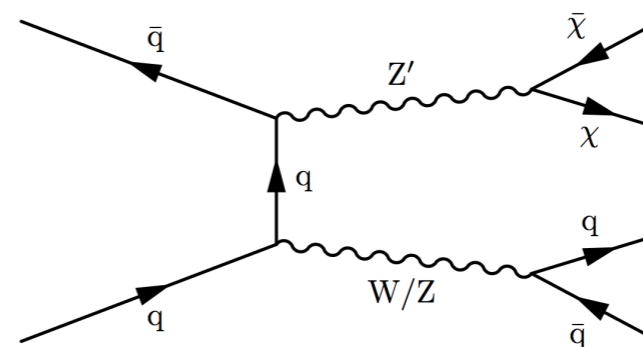
Mono-jet



Mono-photon

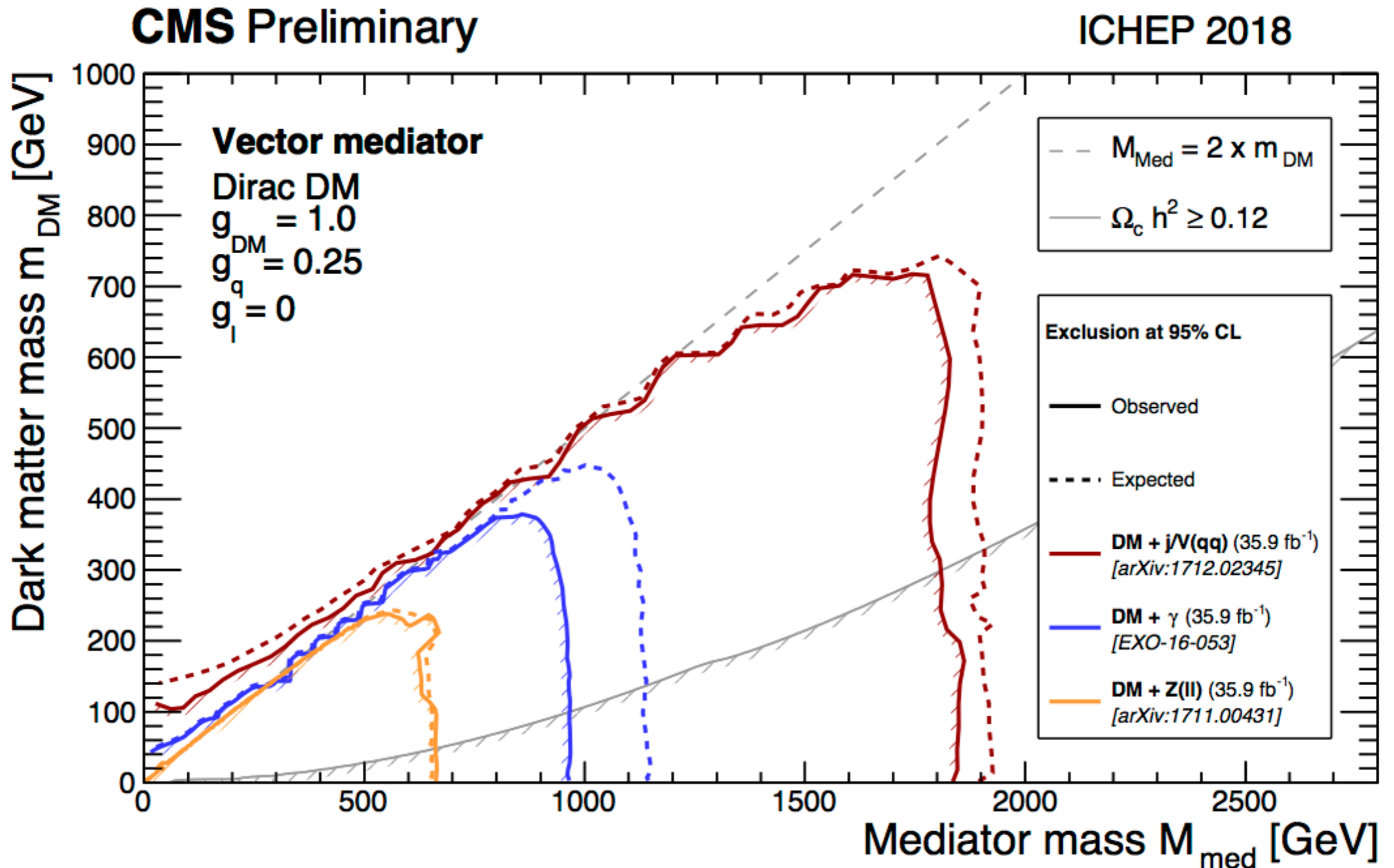


Mono-Z (leptonic)

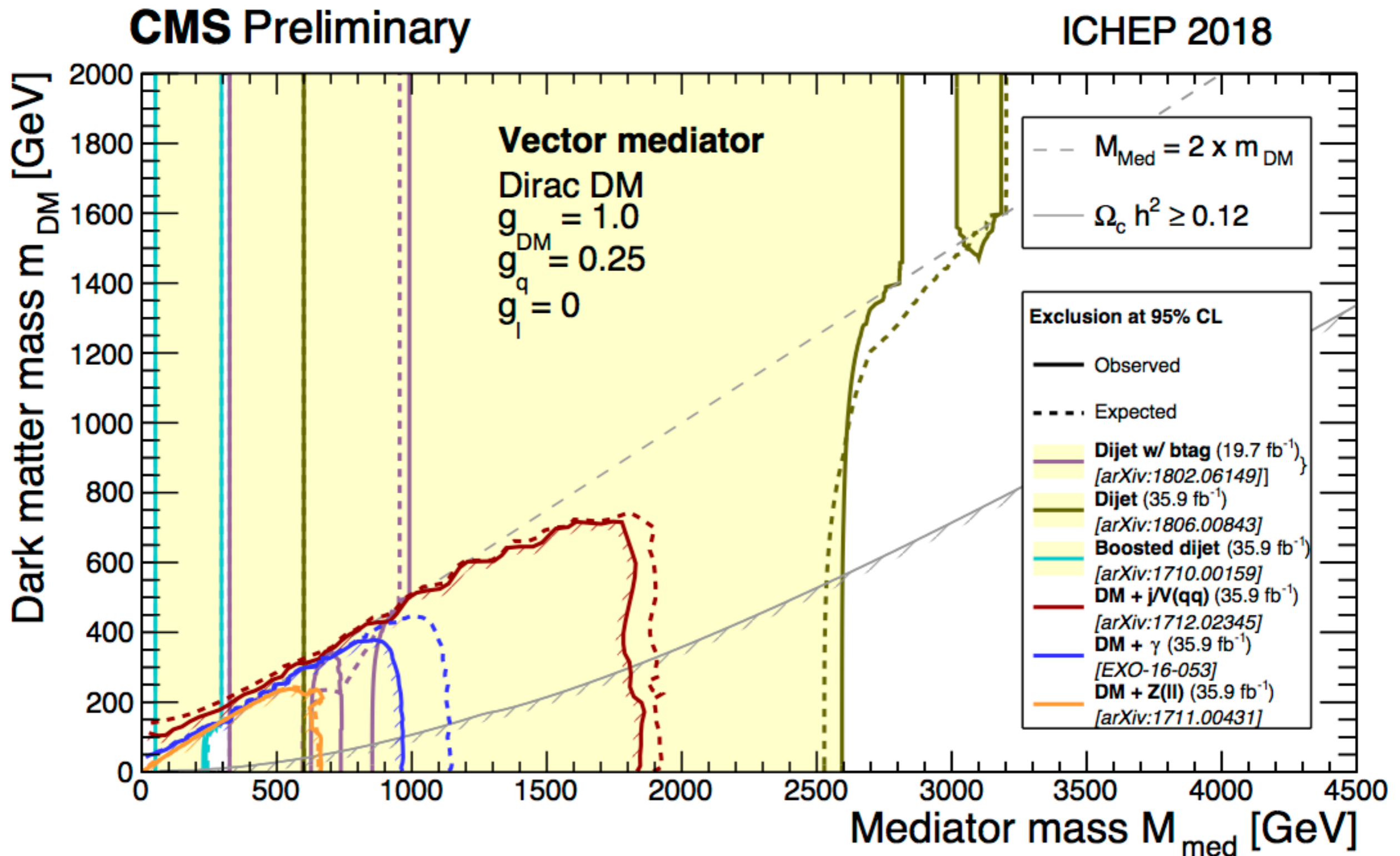


Mono-W/Z (hadronic)

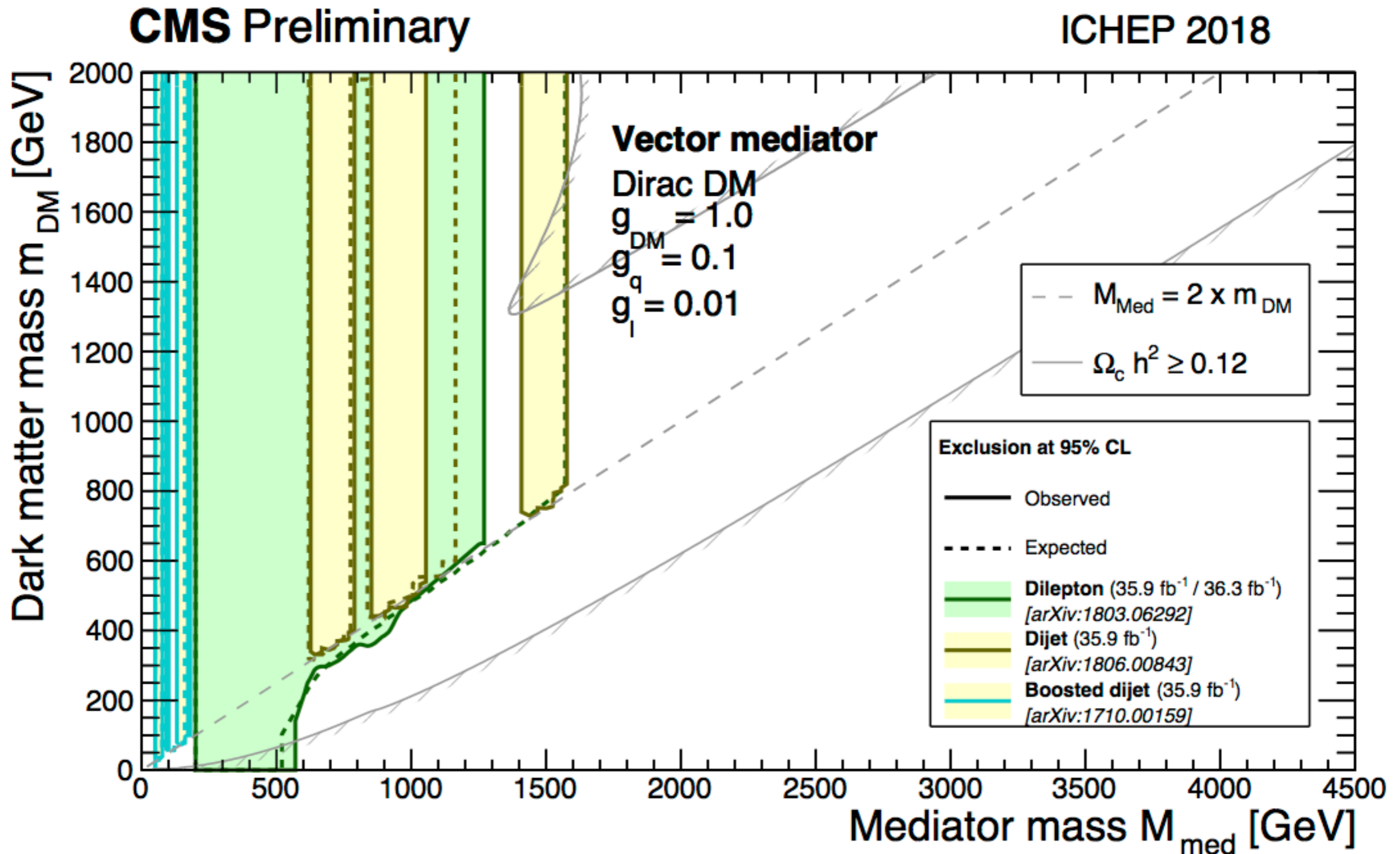
Collider Results Only (Vector Mediator)-Mono-X



Collider Results Only (Vector Mediator)

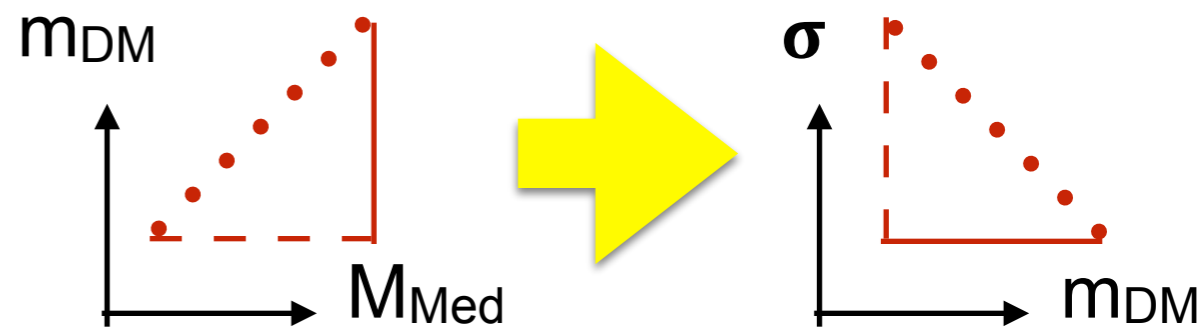
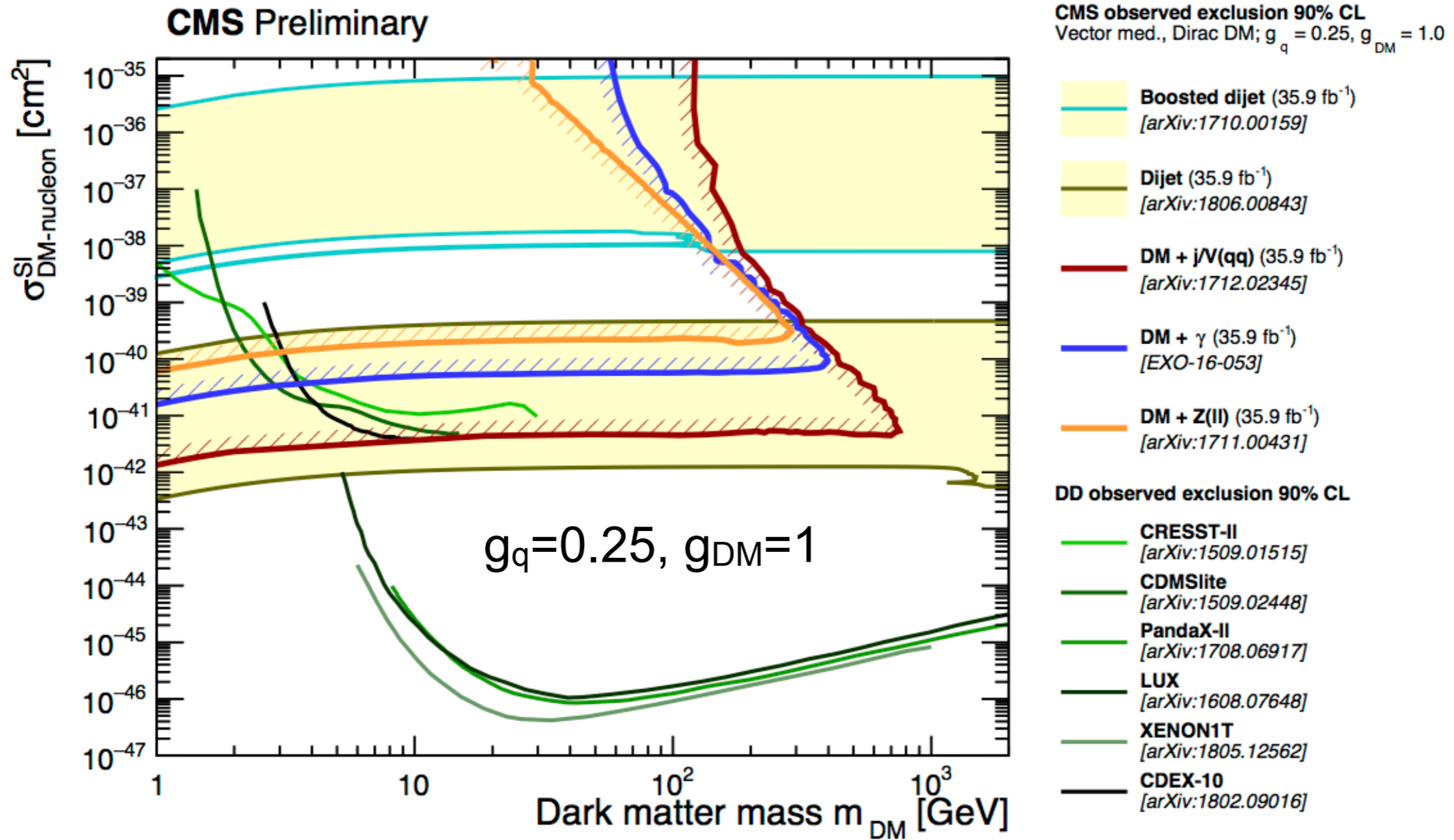


If We Use Different Parameter Values



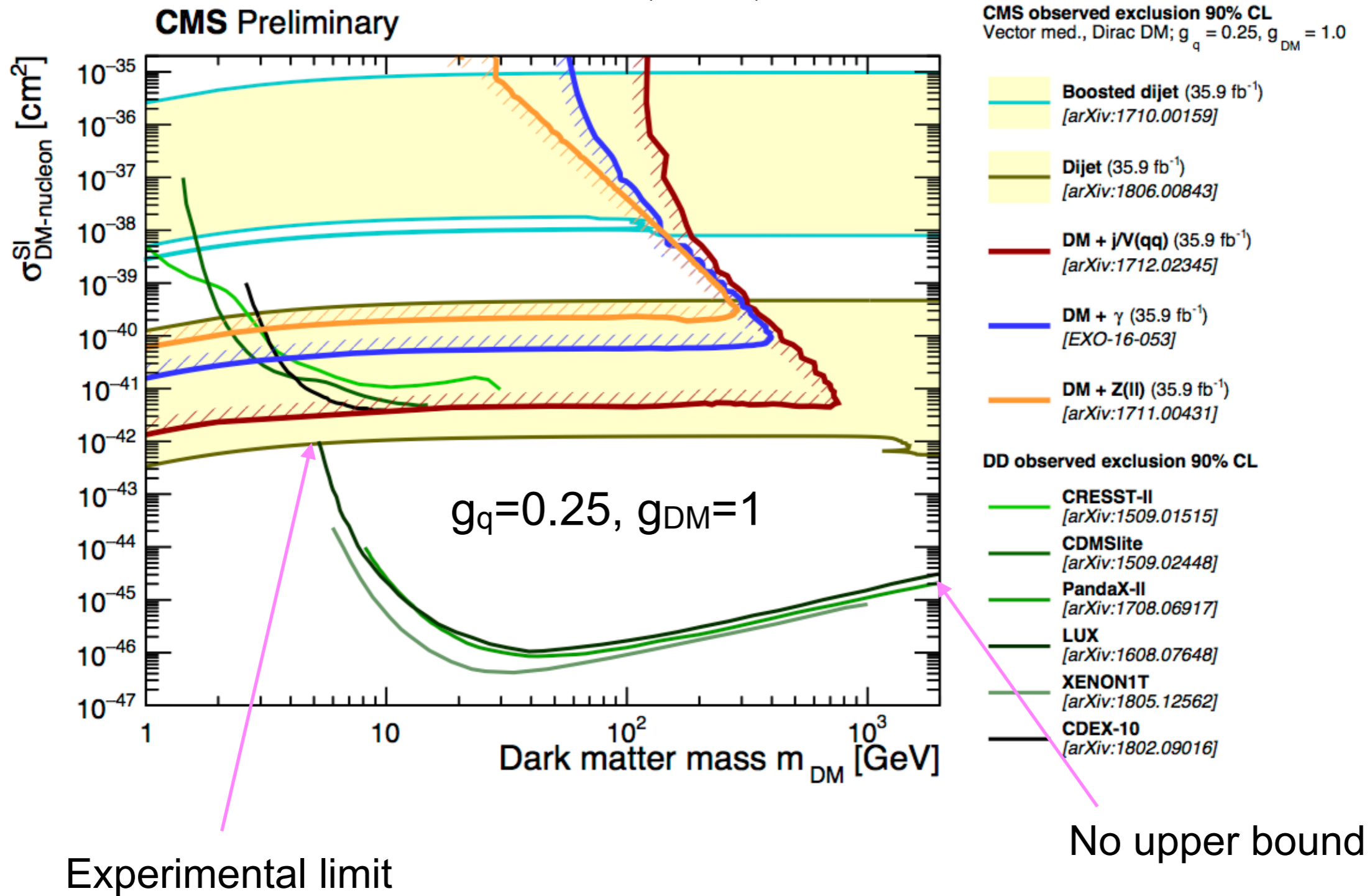
Collider v.s. Non-Collider Experiments (SI)

$$\sigma_{\text{SI}}^{\text{vector}} \approx 6.9 \times 10^{-41} \text{ cm}^2 \left(\frac{g_q g_{\text{DM}}}{0.25} \right)^2 \left(\frac{1 \text{ TeV}}{M_{\text{med}}} \right)^4 \left(\frac{\mu_{n\chi}}{1 \text{ GeV}} \right)^2$$



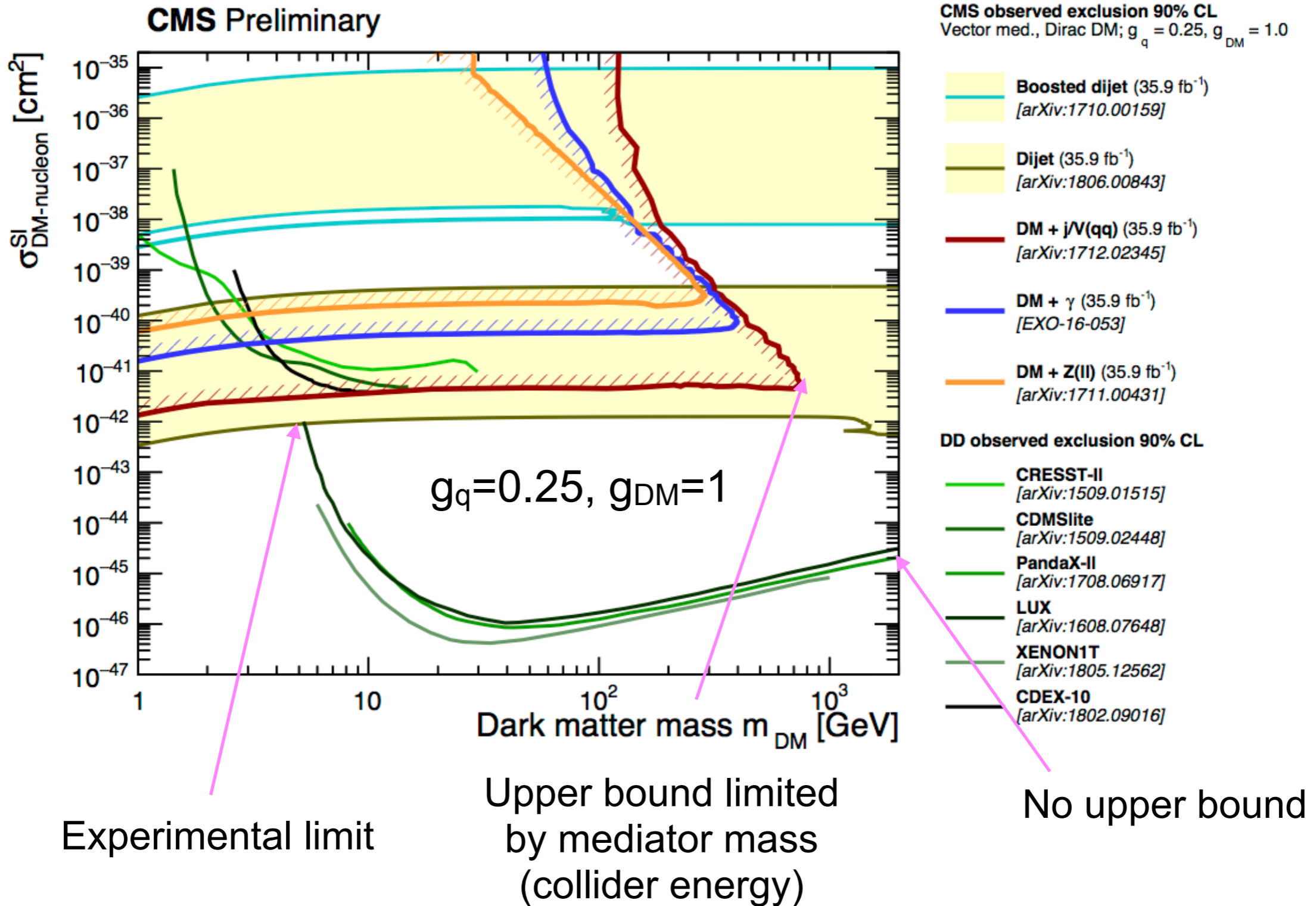
Collider v.s. Non-Collider Experiments (SI)

$$\sigma_{\text{SI}}^{\text{vector}} \approx 6.9 \times 10^{-41} \text{ cm}^2 \left(\frac{g_q g_{\text{DM}}}{0.25} \right)^2 \left(\frac{1 \text{ TeV}}{M_{\text{med}}} \right)^4 \left(\frac{\mu_{n\chi}}{1 \text{ GeV}} \right)^2$$



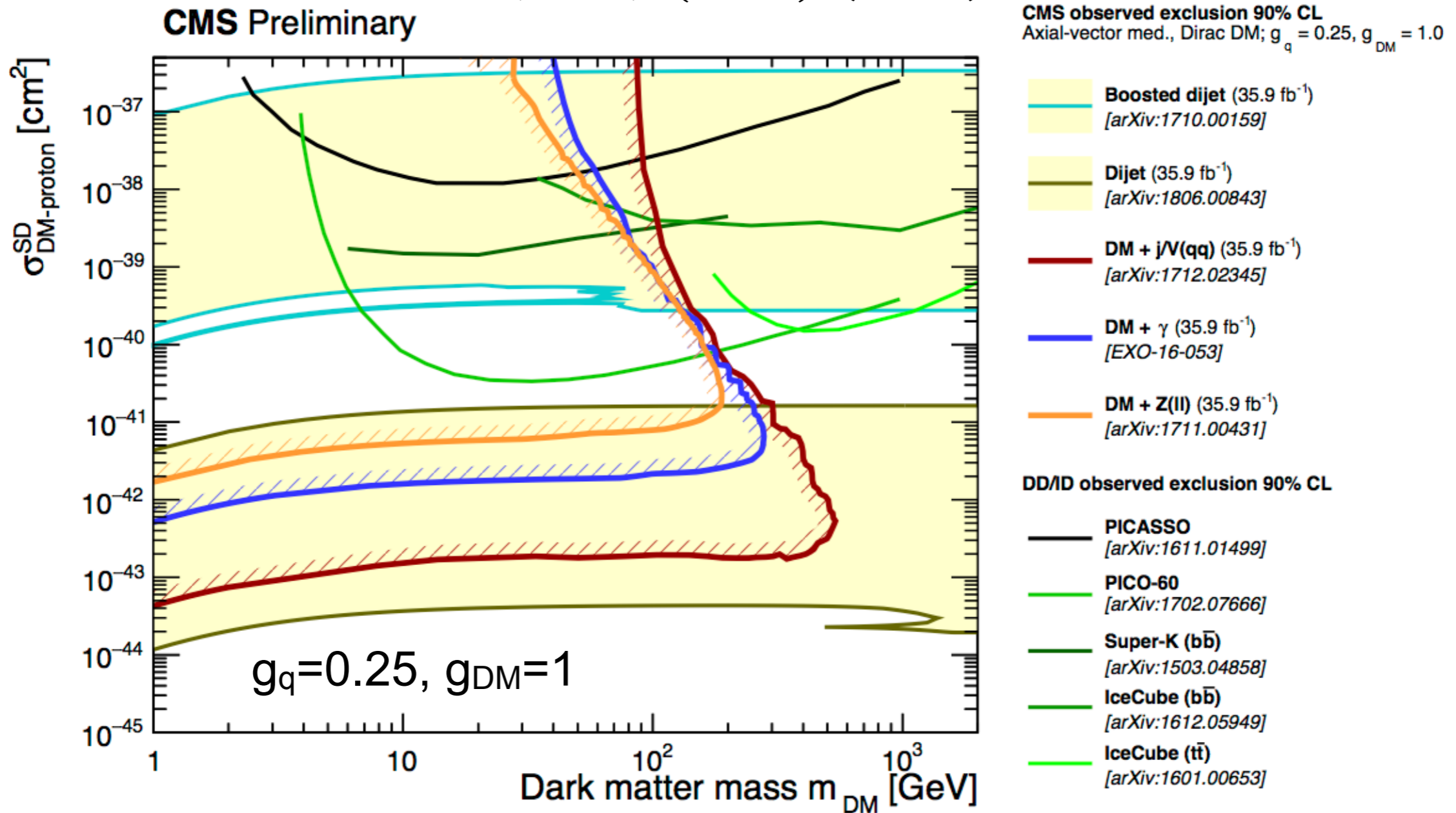
Collider v.s. Non-Collider Experiments (SI)

$$\sigma_{\text{SI}}^{\text{vector}} \approx 6.9 \times 10^{-41} \text{ cm}^2 \left(\frac{g_q g_{\text{DM}}}{0.25} \right)^2 \left(\frac{1 \text{ TeV}}{M_{\text{med}}} \right)^4 \left(\frac{\mu_{n\chi}}{1 \text{ GeV}} \right)^2$$



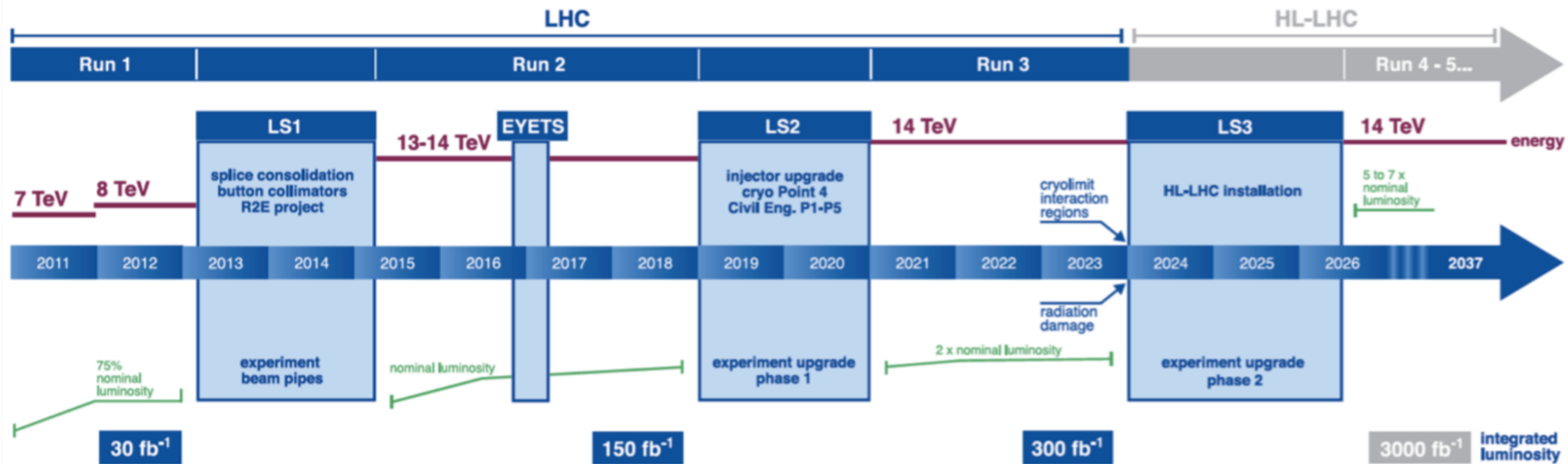
Collider v.s. Non-Collider Experiments (SD)

$$\sigma_{\text{SD}}^{\text{axial}} \approx 2.4 \times 10^{-42} \text{ cm}^2 \left(\frac{g_q g_{\text{DM}}}{0.25} \right)^2 \left(\frac{1 \text{ TeV}}{M_{\text{med}}} \right)^4 \left(\frac{\mu_{n\chi}}{1 \text{ GeV}} \right)^2$$

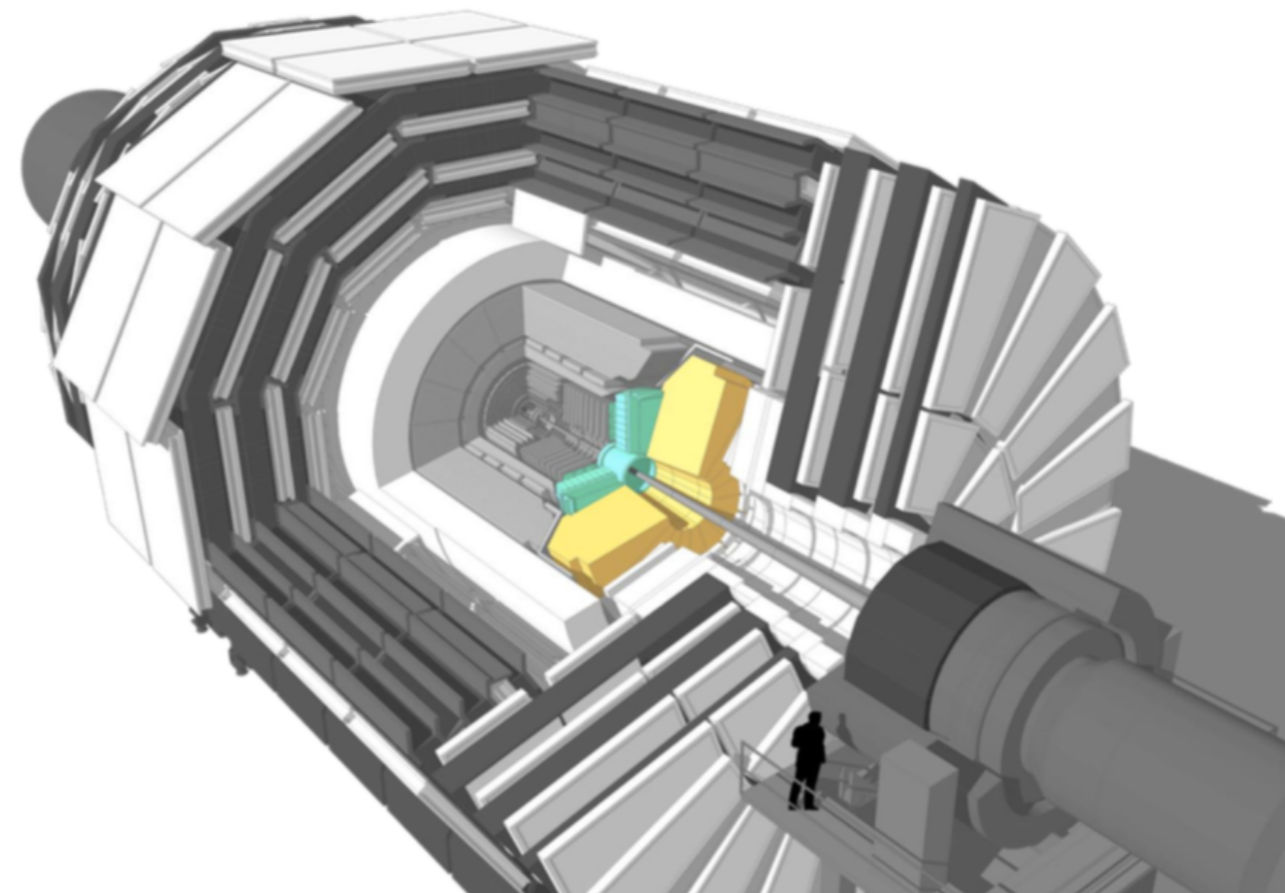
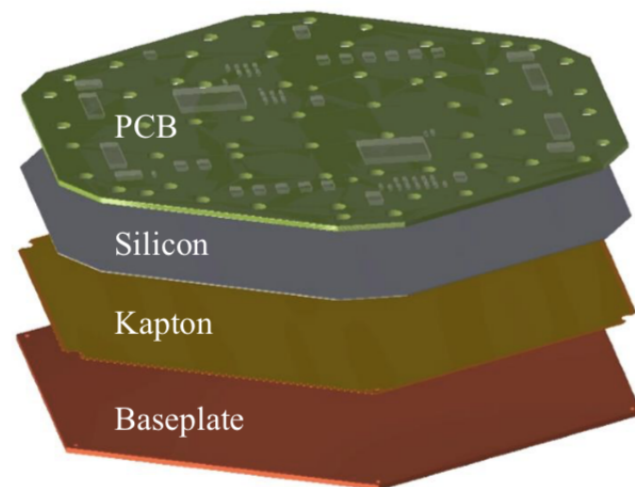
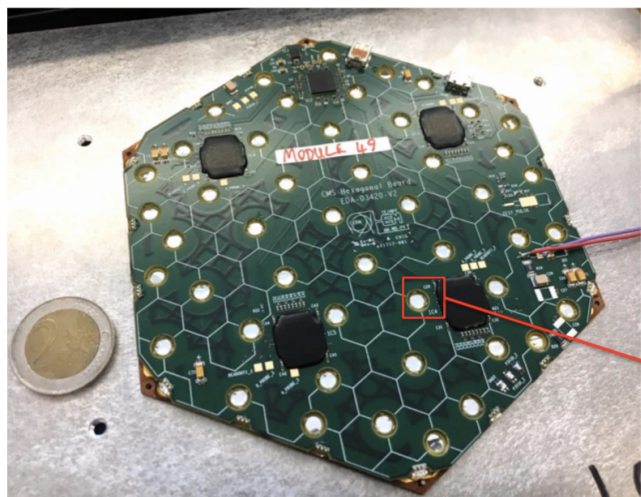
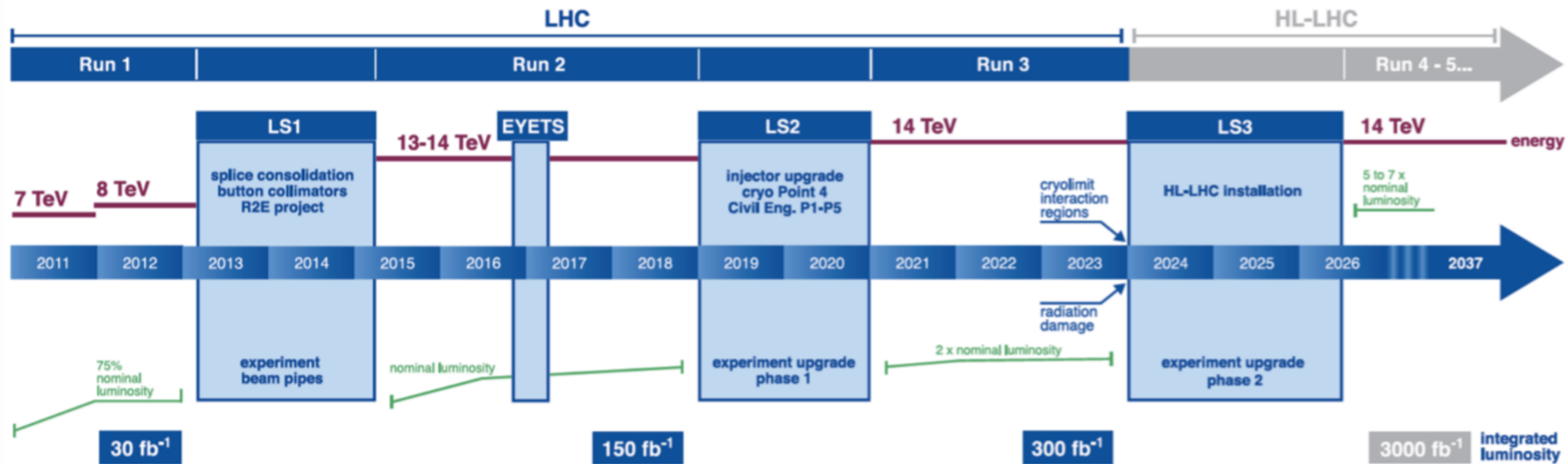


For the model parameters considered here, collider experiments can probe SD cross sections 2-3 orders of magnitude smaller than the non-collider experiments.

CMS Phase-2 Upgrade



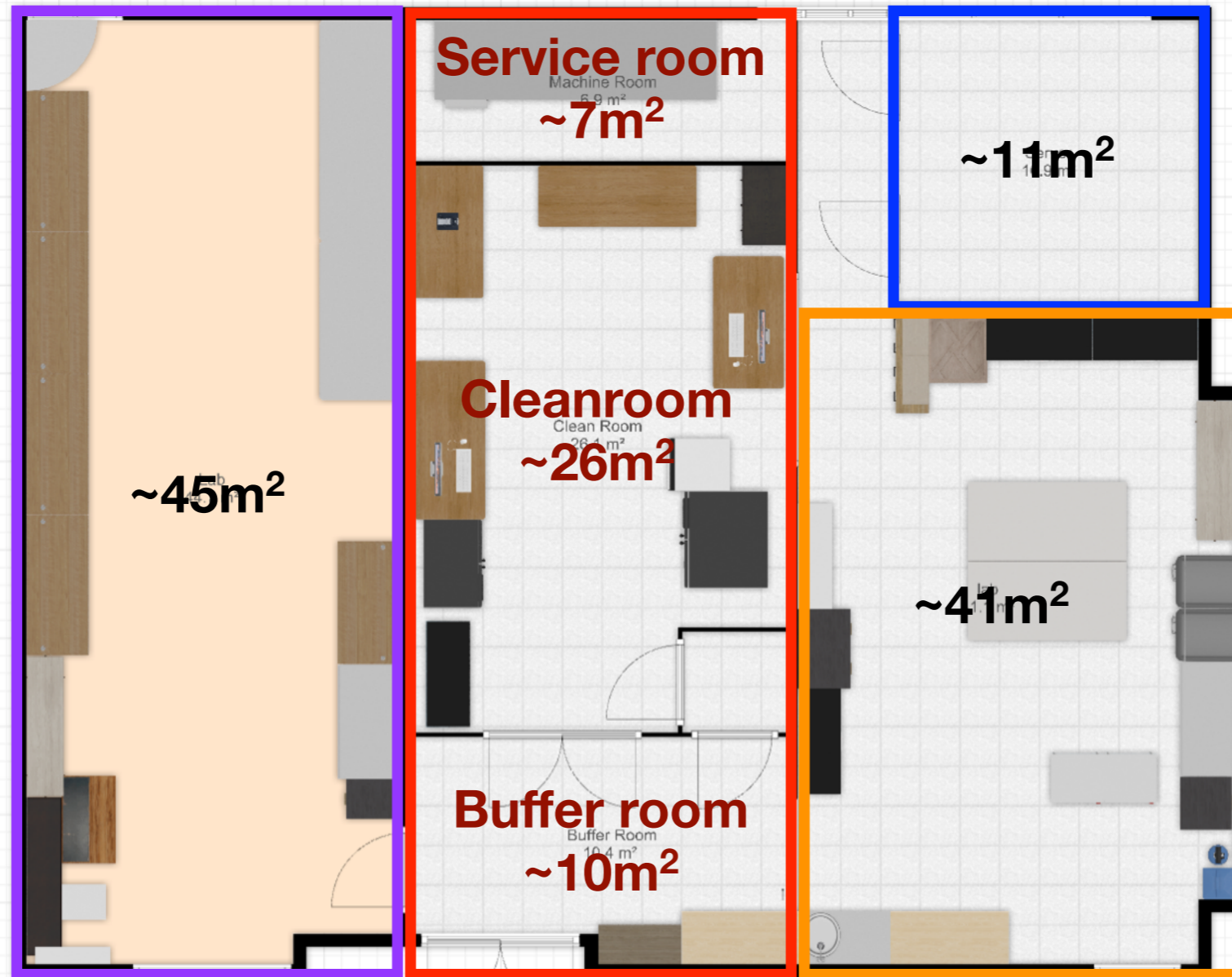
CMS Phase-2 Upgrade



The Detector Lab @ NCU

Grid computing room
for AMS, CMS, KAGRA

~2000 cores
~500 TB



Space for testing,
inspection,
repair and students

Cleanroom for Silicon
detector for CMS/sPHENIX

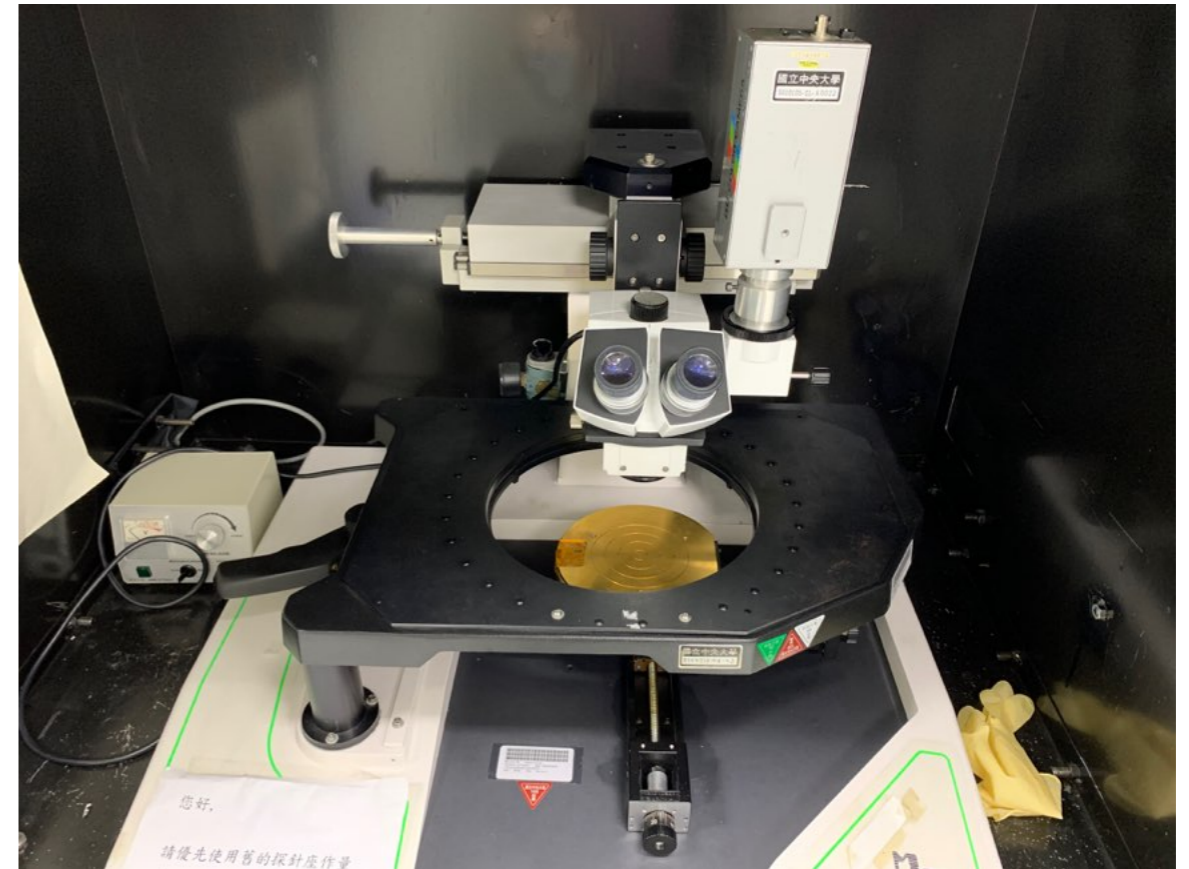
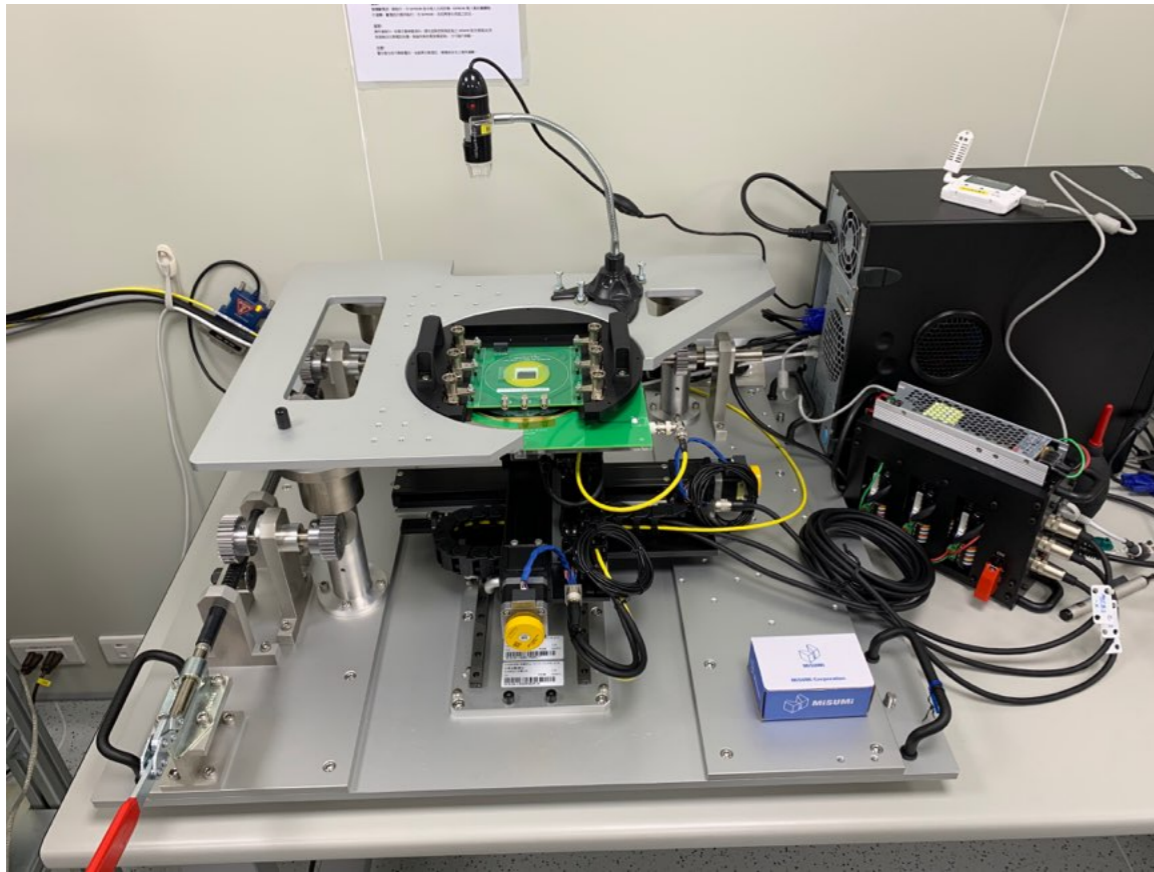
Space for Scintillator+SiPM
detector for muography

Cleanroom



- cleanroom: $\sim 26\text{m}^2$
- service + buffer room: $\sim 17\text{m}^2$
- **class 1000** with temperature and humidity controlled at 22°C and relative humidity (RH) 55% all year round
- fully operation with pressured dry-air service

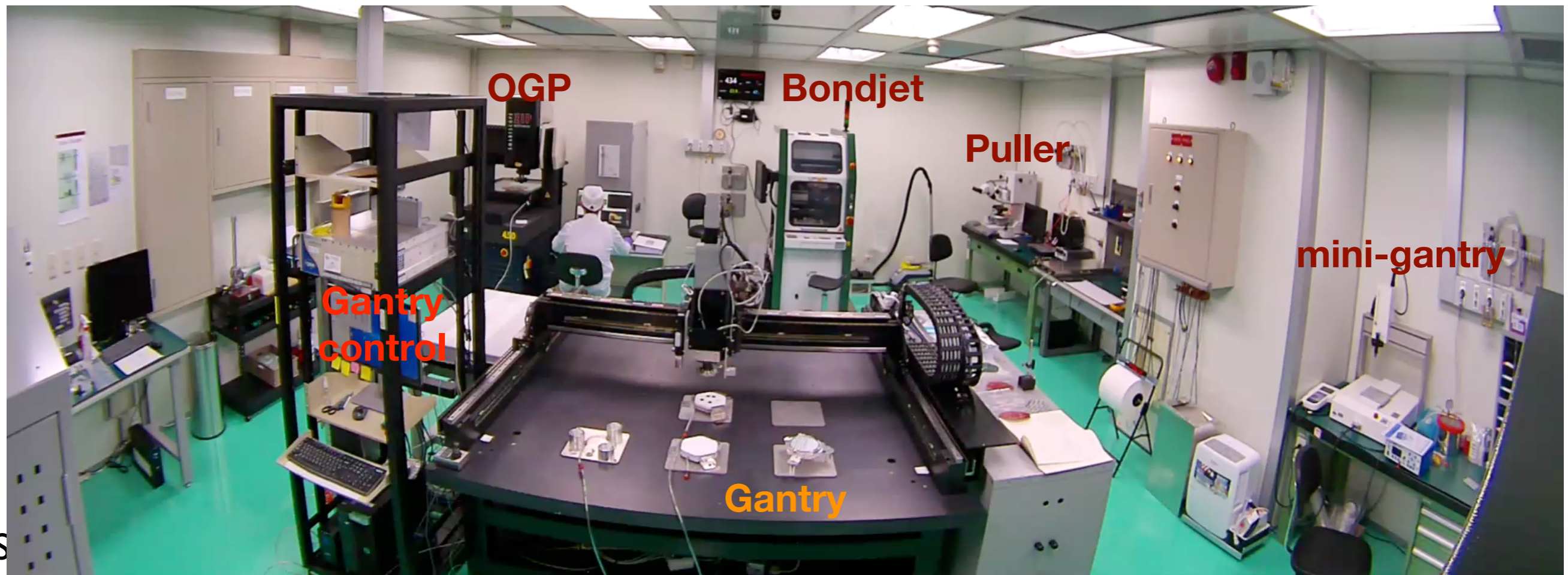
Probe Stations



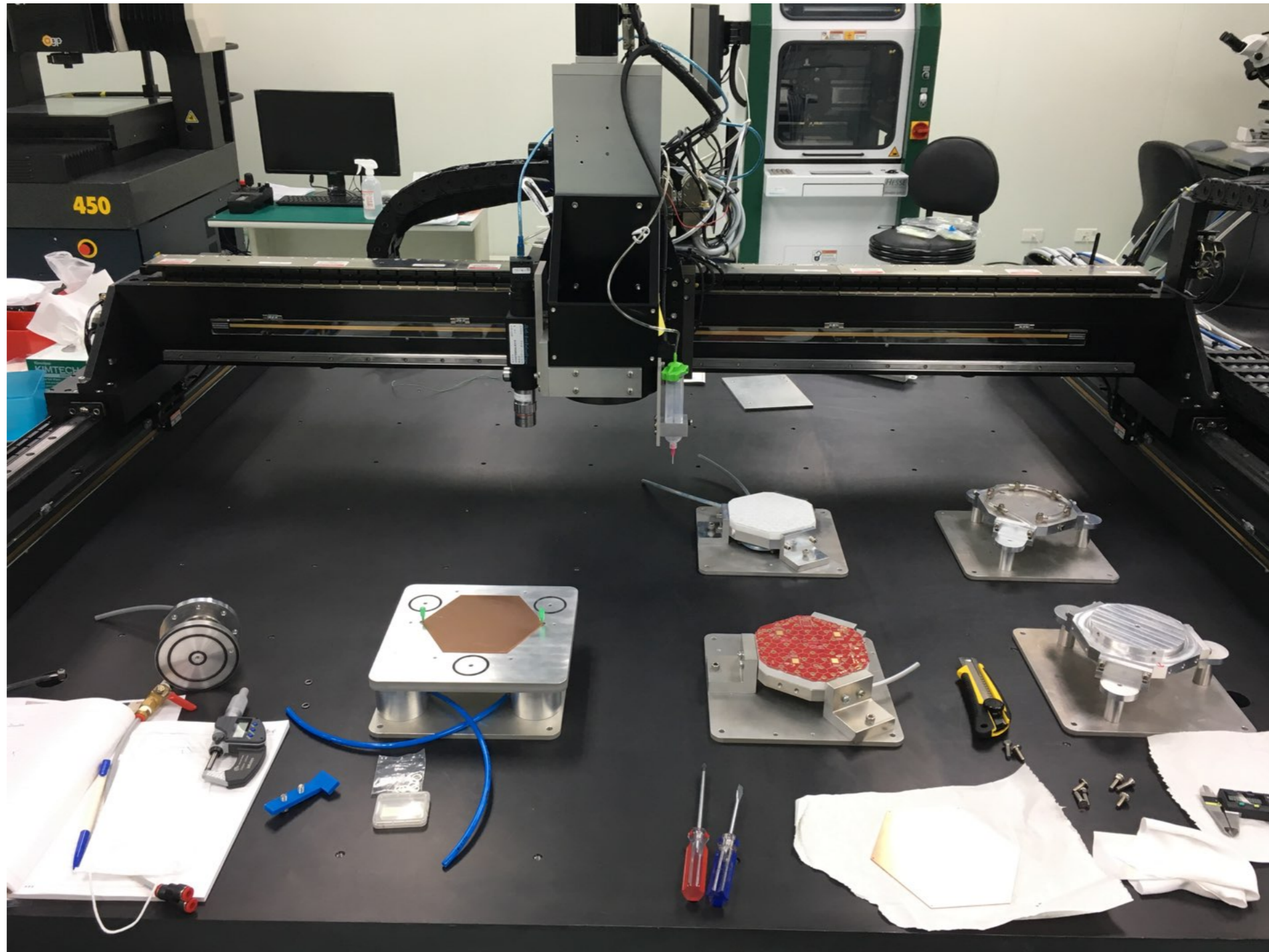
- left: self-designed 8-inch probe station used for the large pad silicon sensors
- right: 4-inch probe station used for PHOBOS and CMS Preshower (being upgraded for sPHENIX)
- A new 8-inch MPI probe station was installed in mid-November for CMS HGCal and sPHENIX

Cleanroom Equipments at NTU

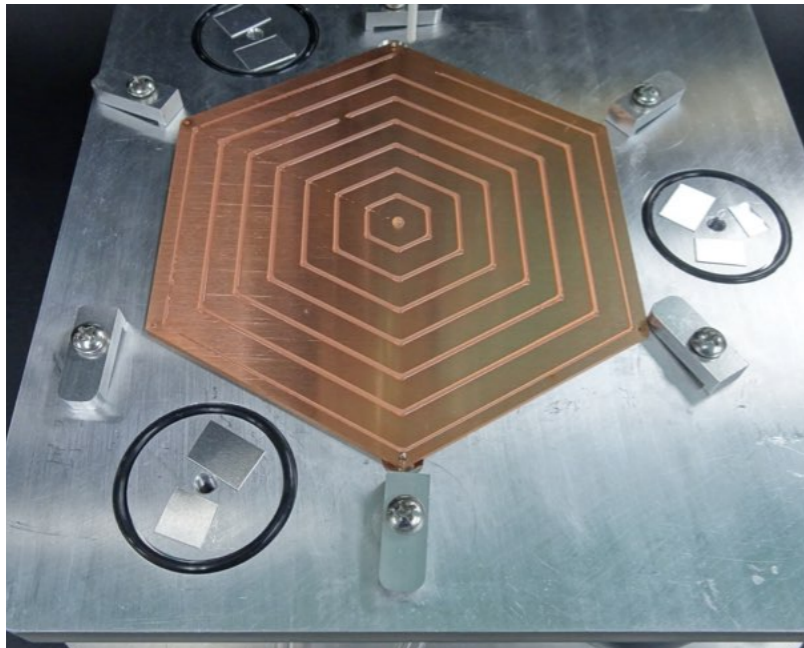
- Aerotech 1.25x1.25 m² robotic gantry with Labview control.
- OGP optical 3d measurement
- Hesse BJ820 automatic Bondjet and DAGE 4000 Bondtester(puller)
- Manual probe station and picoprobes (not visible in this pic)
- glue dispensers, mini-gantry, microscope, degassing chamber, Keithley 2410 and tools ...



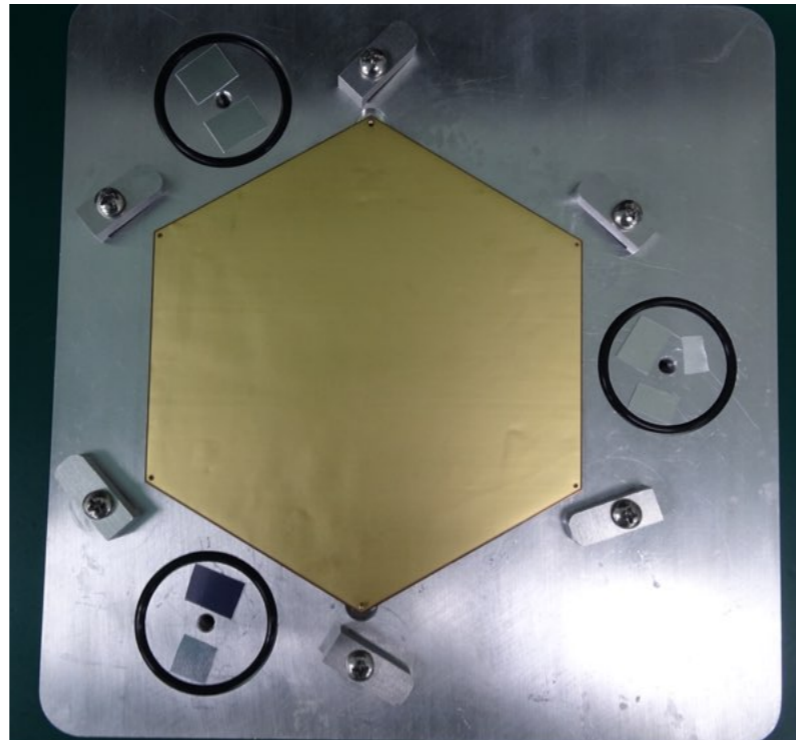
A set of jigs and tooling for 6-inch HGCal module assembly



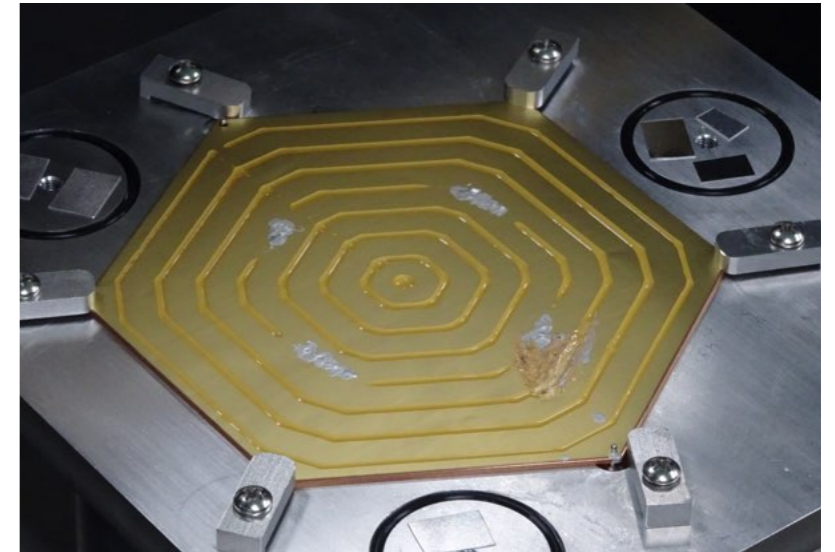
1. Deposit epoxy on Cu baseplate



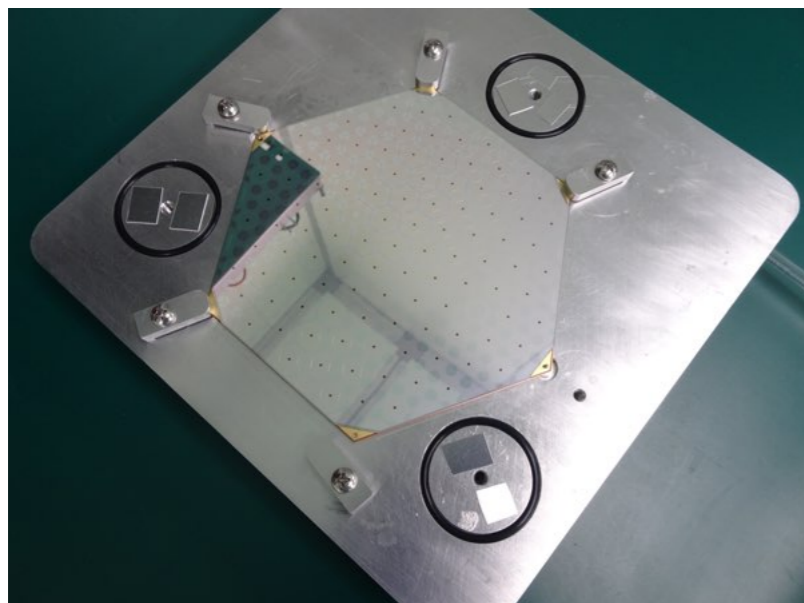
2. Place gold-plated Kapton film



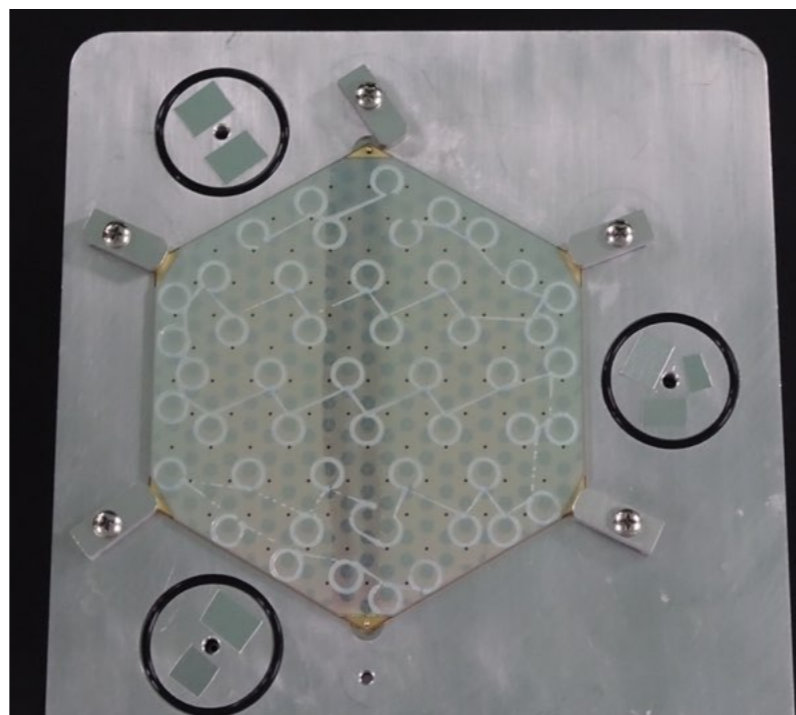
3. Deposit epoxy and silver epoxy on Kapton



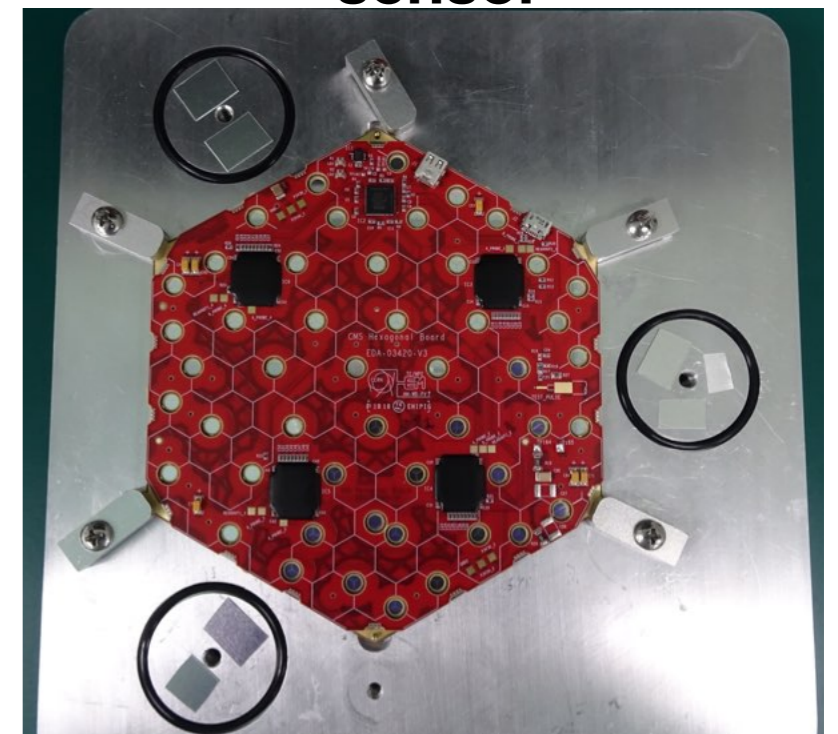
4. Place sensor on top of Kapton



5. Deposit epoxy on sensor, avoiding opening bond pads

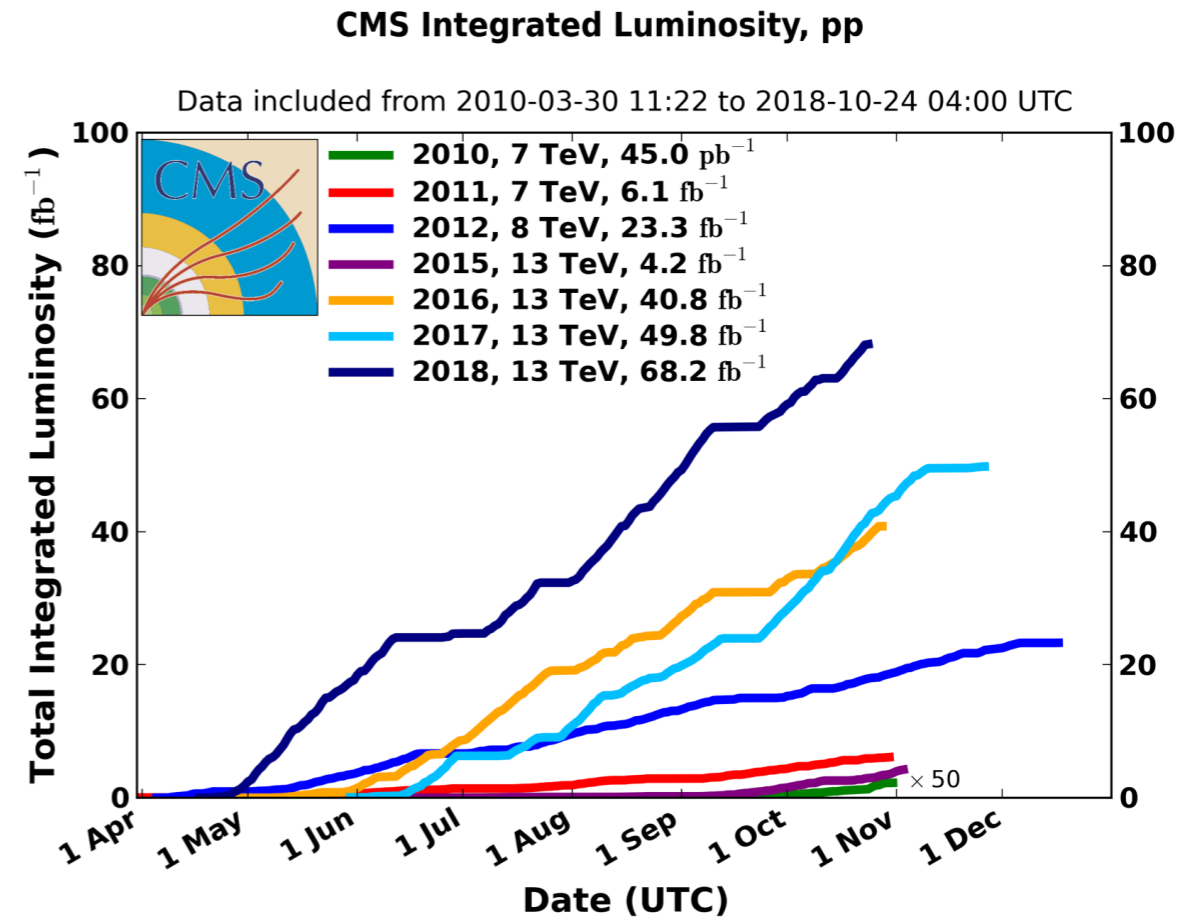


6. Place PCB on top of sensor



Conclusion and Outlook

- CMS has an extensive dark matter program, including both searches for mediators and searches in mono-X channels
- 137 fb⁻¹ of full Run II data are yet to be analyzed



- Moving towards more advanced/sophisticated models
 - t-channel production
 - spin-2 mediators, long-lived mediators or intermediate “dark” particles
- Detector upgrade going on and Taiwan (NCU/NTU) is playing a major role in the endcap calorimeter