

A Potential Astrophysical Test of Quantum Gravity

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Overview

- ▶ First pulsar orbiting black discovered in 2013!
- ▶ Hawking Information problem
- ▶ Fuzzballs
- ▶ Pulsar lensing

Magnetar PSR J1745-2900



(credit: MPIfR/Ralph Eatough) mysterious discovery in 2013,
orbiting galactic center black hole. Rafikov-Lai (2006): precision
GR test?

Black Hole Information



Hawking (1974): black holes radiate: $T = hc/k_B r_s \sim \mu\text{K}$. High entropy: number of photons emitted $S/k_B \sim 10^{77}$. Very slow: one photon of $\lambda \sim r_s \sim \text{km}$ each $\lambda/c \sim \text{ms}$. (1981): information loss? Evaporation is a Schwinger mechanism, does not depend on inside of black hole.

Dilemma

- ▶ No Hair: all black holes look identical after a short time (hour?)
- ▶ radiation only depends on outside of BH
- ▶ emitted radiation does not depend on formation history
- ▶ leads to microscopic time irreversibility of physics!
- ▶ breakdown of causality/unitarity?
- ▶ An initial pure state evolves into mixed state after a Page time (half the mass is lost).

Entropy solution

- ▶ string theory to the rescue!
- ▶ Strominger-Vafa (1996): counting of microstates
- ▶ unitarity saved?
- ▶ Stringy counting not possible in classical limit: what happens with Hawking's argument?

Fuzzballs



Samir Mathur+ (2002+): solutions to Hawking problem must be either non-local or hairy.

Orders of orders of magnitude

- ▶ Saha, partition function: $\frac{P(n_1)}{P(n_0)} = \frac{g_1}{g_0} \exp\left(-\frac{\Delta E}{k_B T}\right)$
- ▶ probability to observe in substantially non-schwarzschild state:
- ▶ $\Delta E \sim mc^2$
- ▶ $\frac{\Delta E}{k_B T} \sim 10^{77}$
- ▶ $S_1 \sim k_B \log g_1 \gtrsim 10^{77}$
- ▶ no-hair may be a great mis-estimate, off by 10^{77} orders of magnitude!

Fuzzballs

- ▶ round black holes have minimum surface, are most unlikely!
- ▶ constructive stringy solutions of some eigenstates: no horizon for no entropy
- ▶ classical Black Holes are superpositions of “naked” microstates
- ▶ evades Hawking’s argument: no schwarzschild background
- ▶ multipole deviation from GR $\sim (r_S/r)^{l+2}$

Landscape

- ▶ Firewall: aging of BH, destruction of observer
- ▶ remnants (Cornucopions)
- ▶ loops
- ▶ loss of unitarity/causality
- ▶ scientific test?

Lens

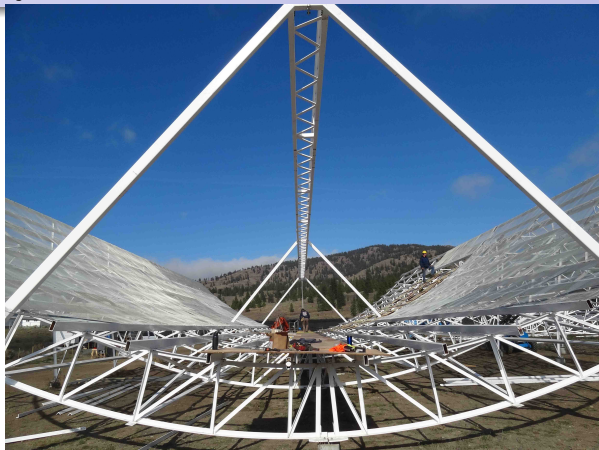


(credit: wikipedia) multiple imaging of pulsars: Boyle+ (2011+), Pen+ (2011+): interference of lensed images. Measure space-time metric to \sim mm at Einstein radius.

Lensing Prospects

- ▶ ideal setup: pulsar orbiting BH at $\sim 10,000 r_S$
- ▶ inclination similar to Einstein radius $\sim 1^\circ$
- ▶ two main images form double slit interferometer (Young) near conjunction
- ▶ quantum lens: expect image decoherence $\sim 10^{-6} r_S \sim \text{cm}$
- ▶ order unity effect in scintillation pattern

New surveys



11 pulsar-neutron star binaries, 1 pulsar-BH binary known. New surveys (e.g. SKA, CHIME+) will increase number 10-fold.

Conclusions

- ▶ Astrophysical test of quantum gravity?
- ▶ Promising future if high inclination BH-PSR binaries are discovered
- ▶ large cylinder telescopes (e.g. CHIME+) for searching
- ▶ scientific test for some scenarios of quantum gravity:
coherence of pulsar scintillation